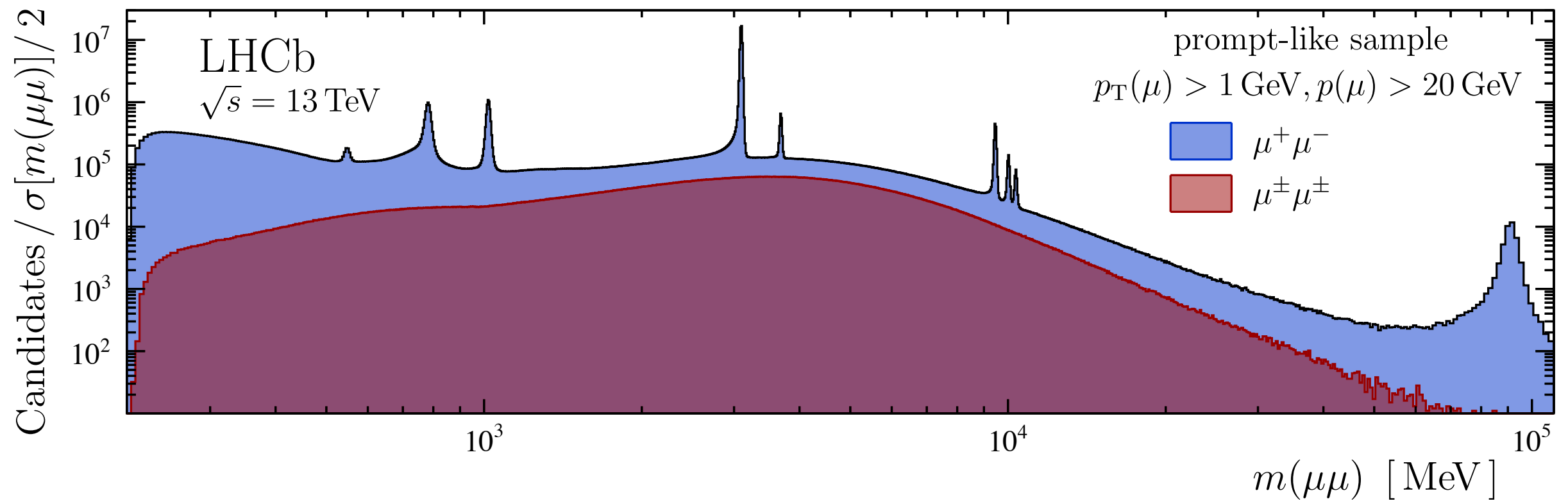




*Physics at the LHC - WS 2021-2022*

# *Real time analysis and searches at the intensity frontier at LHCb*

Martino Borsato

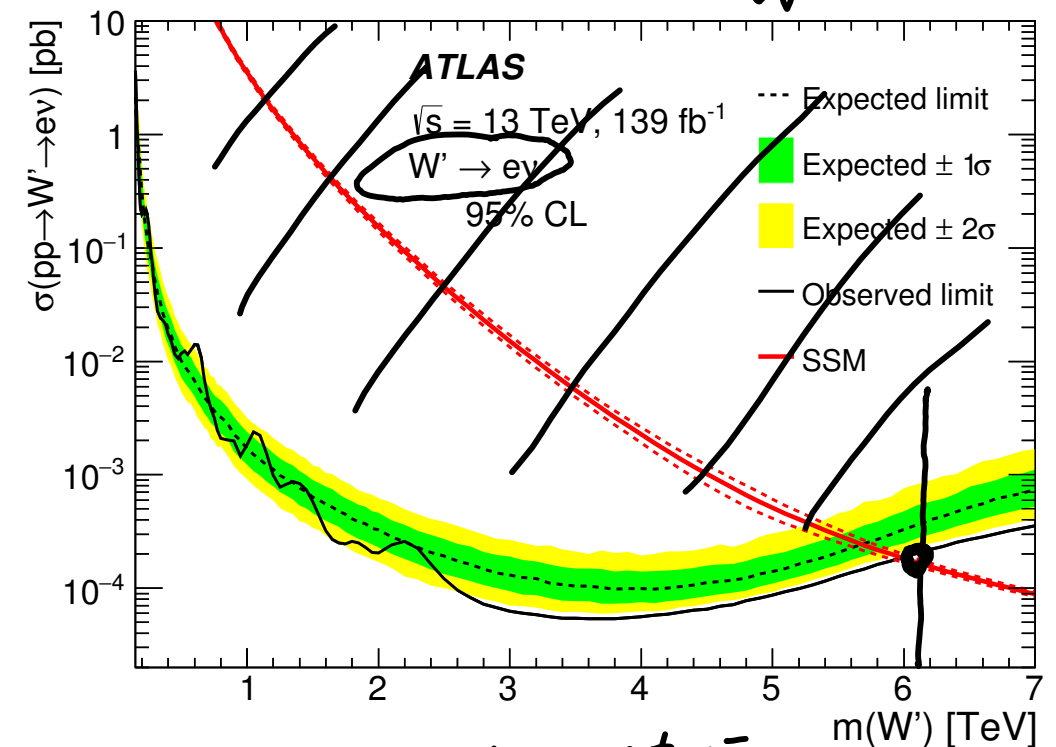
*Lecture 4/4 - Friday Feb 4<sup>th</sup> 2022*

# The energy frontier

$$W'^+ \rightarrow \ell^+ \nu$$

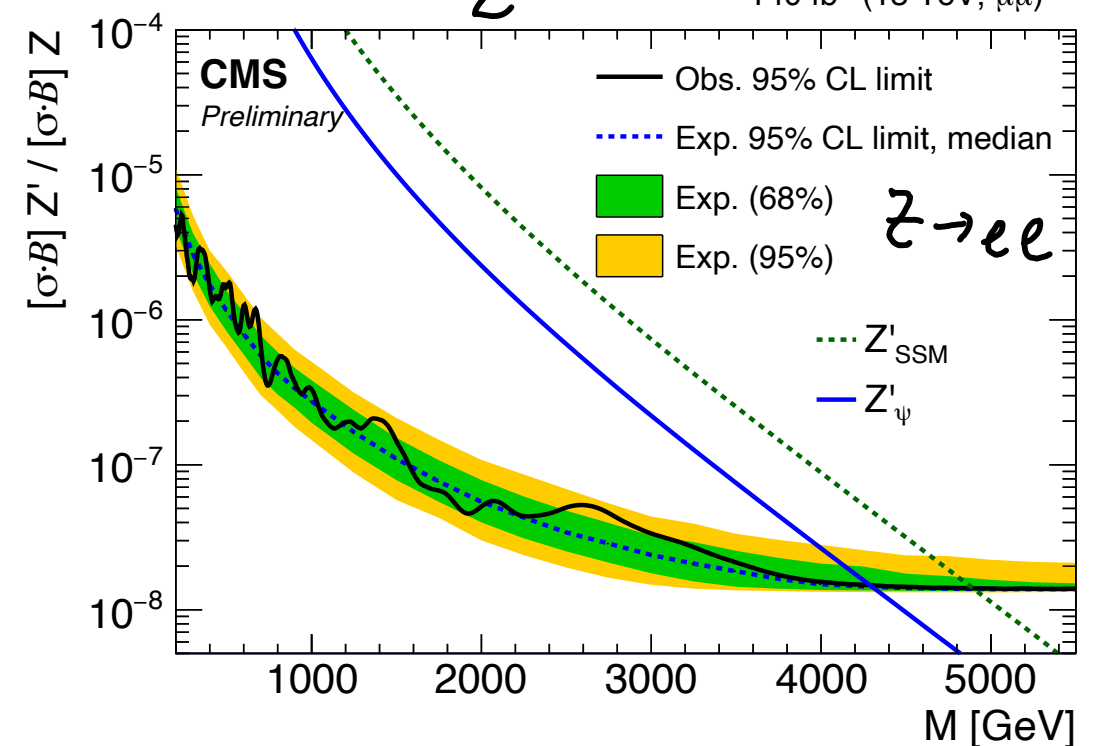
- **Energy frontier exploration** has greatly advanced particle physics
  - Often guided by indirect signs in lower energy processes
  - Example: CPV  $\rightarrow$  third family
- The LHC is our front runner
  - **Discovery of the Higgs**
  - Exploration of the multi-TeV range
  - So far no hint of new heavy objects
  - Lower limits up to several TeV

● **What if the new particles have smaller couplings to the SM?**



$$Z' \rightarrow \ell^+ \ell^-$$

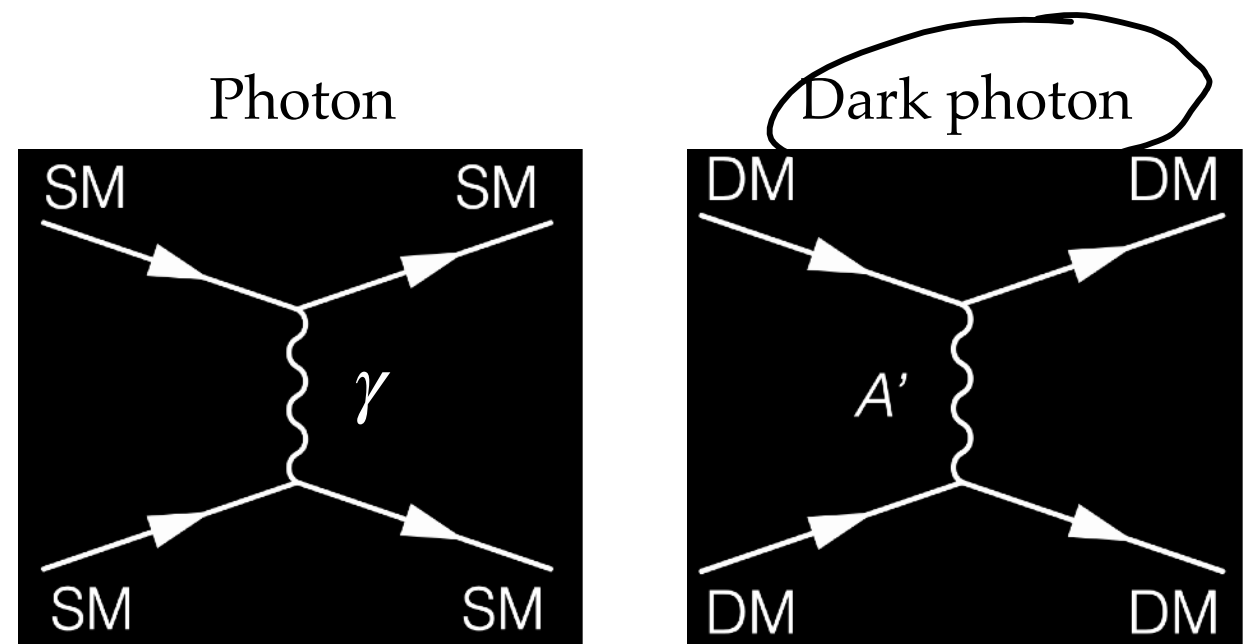
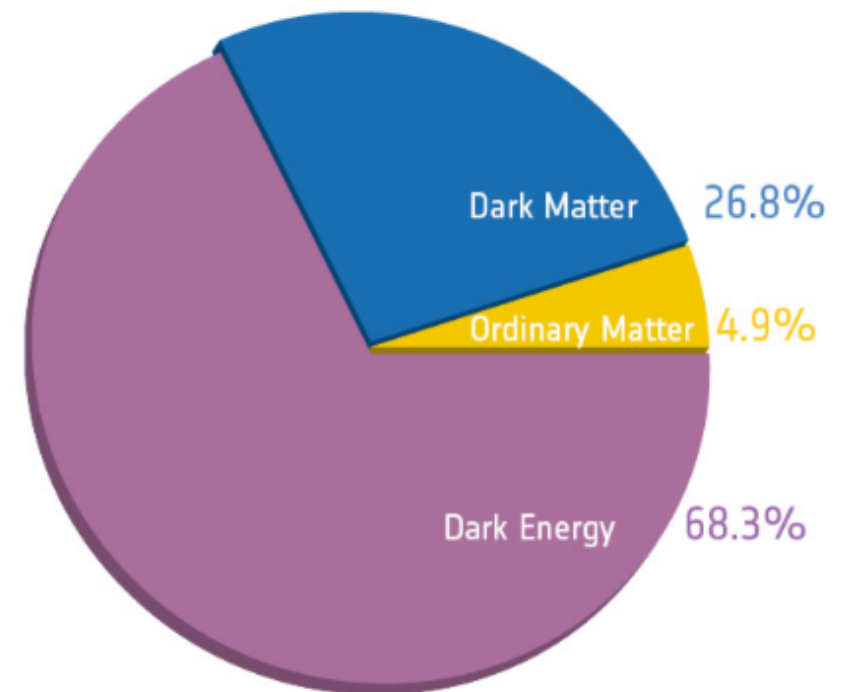
140 fb<sup>-1</sup> (13 TeV,  $\mu\mu$ )



$$Z \rightarrow e e$$

# Dark Forces

- Astrophysics + Cosmology → there must be Dark Matter
  - Dark = only gravitational effect detected → does not interact e.m. or strongly
  - No candidate particle in the SM
- The weak solution:
  - Dark Matter interact via the weak force
  - Then its mass should be above the weak scale (100 GeV)
- Alternative solution:
  - Dark Matter interacts via new force
  - Couples indirectly to ordinary matter (through quantum loops)
  - **Very small couplings, lighter masses**



# Dark Forces

- **Small coupling**

- Small cross section
- Need high luminosity
- Could have escaped previous experiments (intensity frontier)

$$Rate = \sigma \cdot \mathcal{L}$$

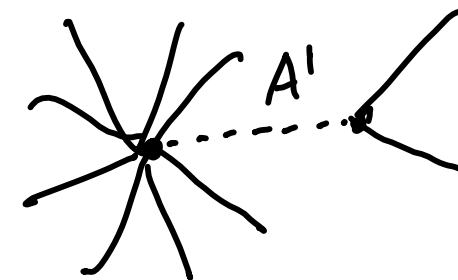
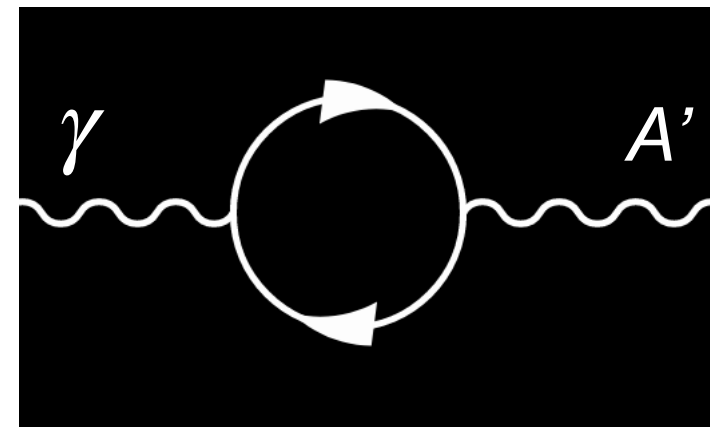
- **Lighter mass**

- Cheap to produce at LHC
- Production mostly in the forward region
- Comes with large background of SM particles

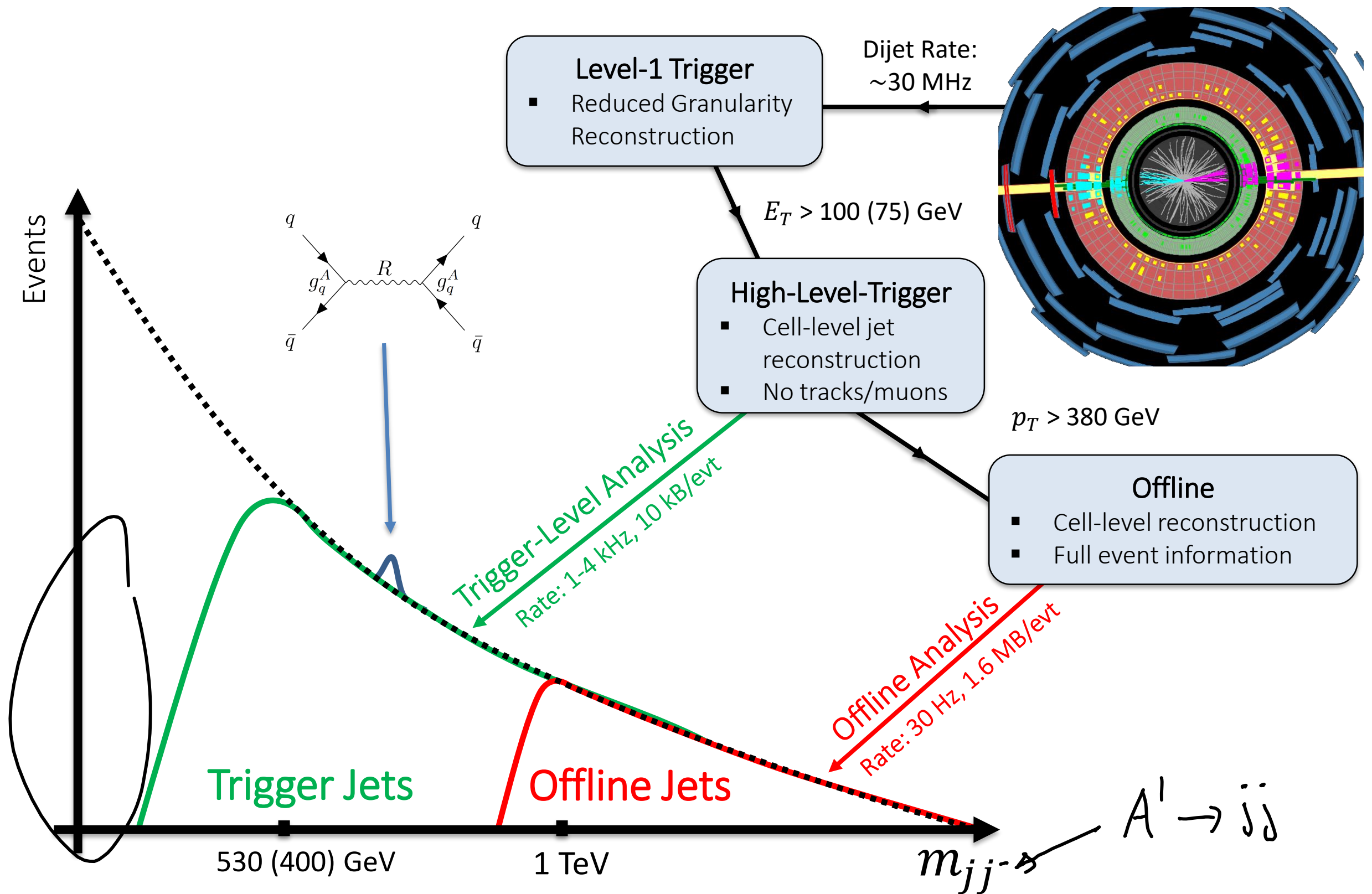
- **Small coupling + Light mass**

- Can have sizeable lifetime!
- Displaced vertex signature

Kinetic mixing

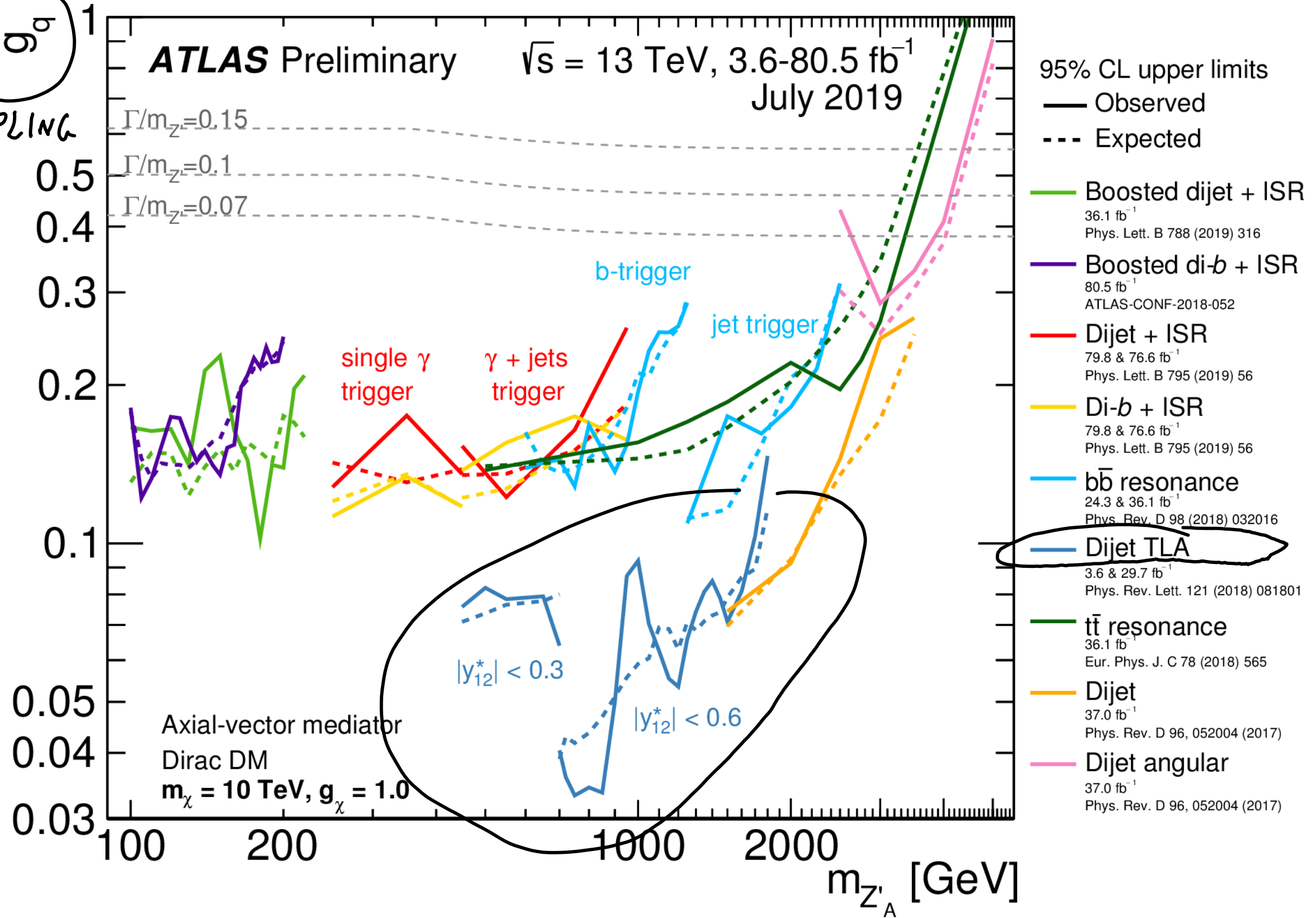


# Dijets : TLA spectrum



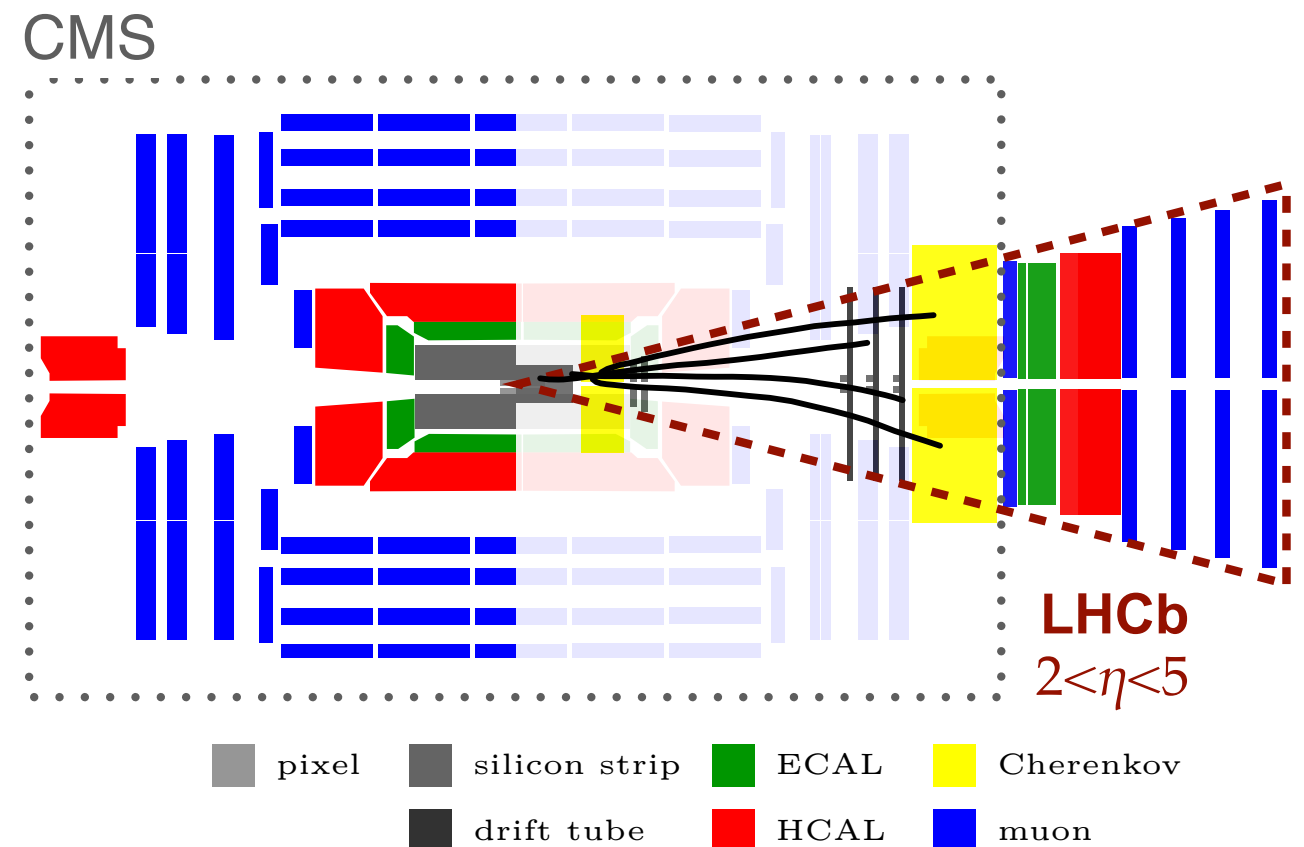
# All dijet searches

$\sigma$   
Coupling



# LHCb vs ATLAS/CMS

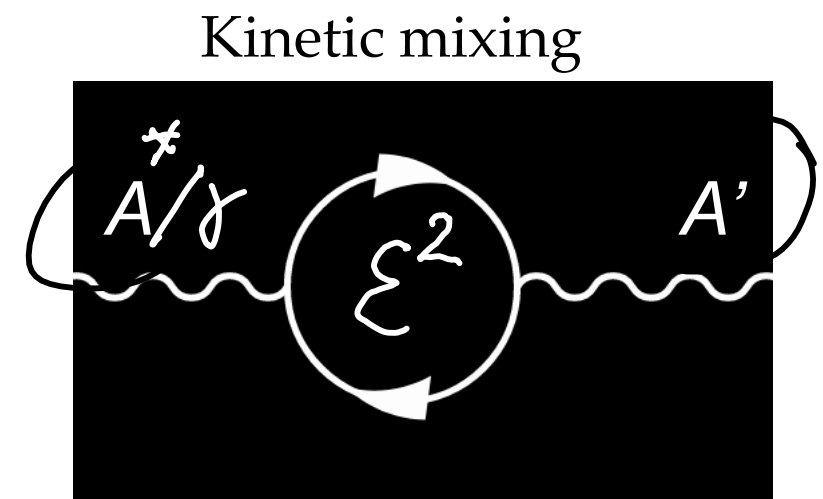
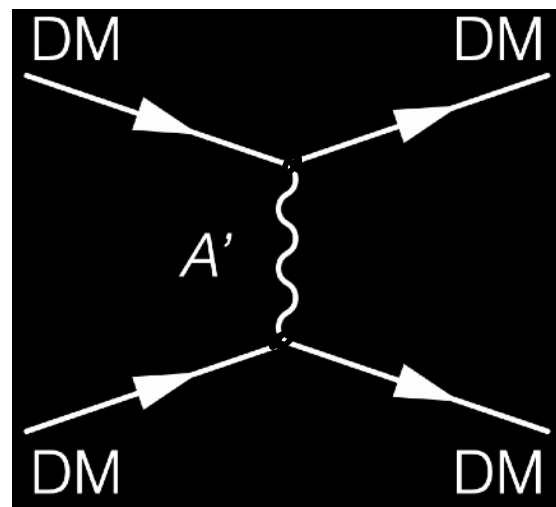
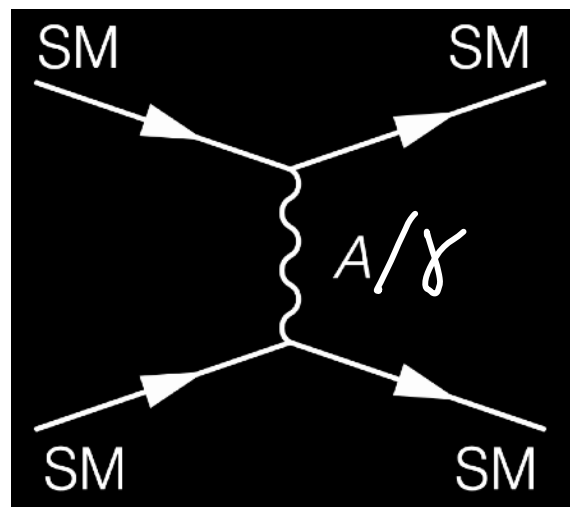
- **Lower luminosity** (low pile-up)
  - $\sim 1/8$  of ATLAS/CMS in Run 1
  - $\sim 1/20$  of ATLAS/CMS in Run 2
- **Capable of very soft triggers!**
  - At hardware level (L0):
    - ▶ Muons with  $p_T > 1.5$  GeV
    - ▶ Calo deposits with  $E_T > 3$  GeV
  - At Software level (HLT):
    - ▶ Displaced vertex triggers
    - ▶ Real-time calibration in Run 2:
      - ▶ Allowed “real time analysis” i.e. keep only interesting part of event



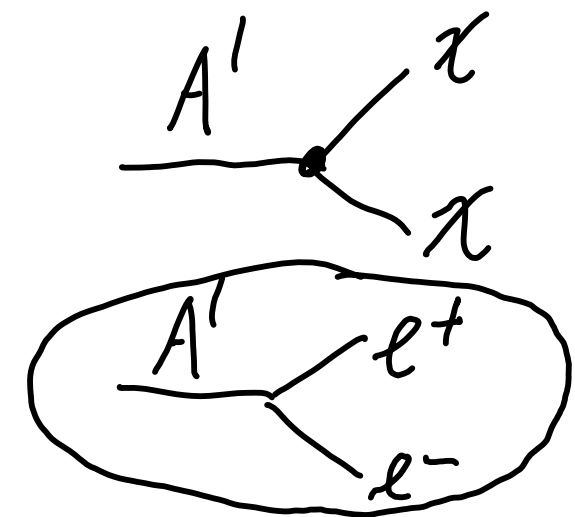
# *Dark photon search*

at LHCb

# Dark photons phenomenology

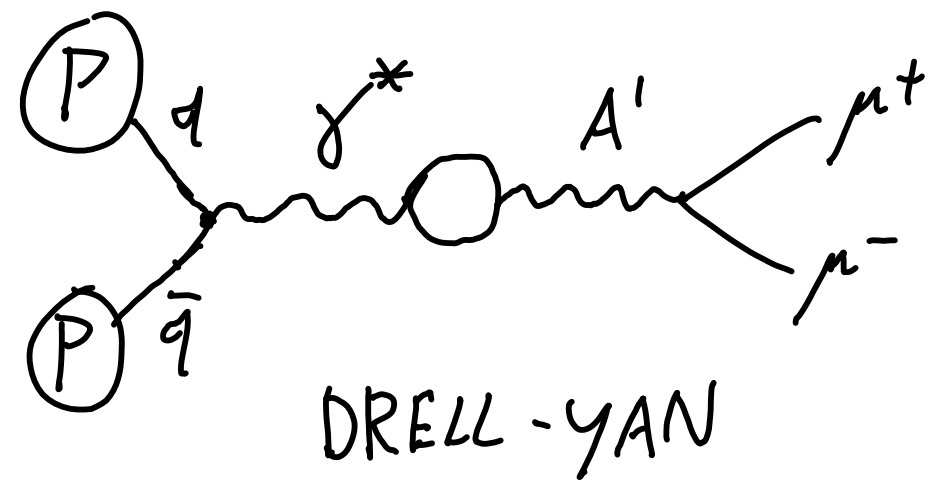
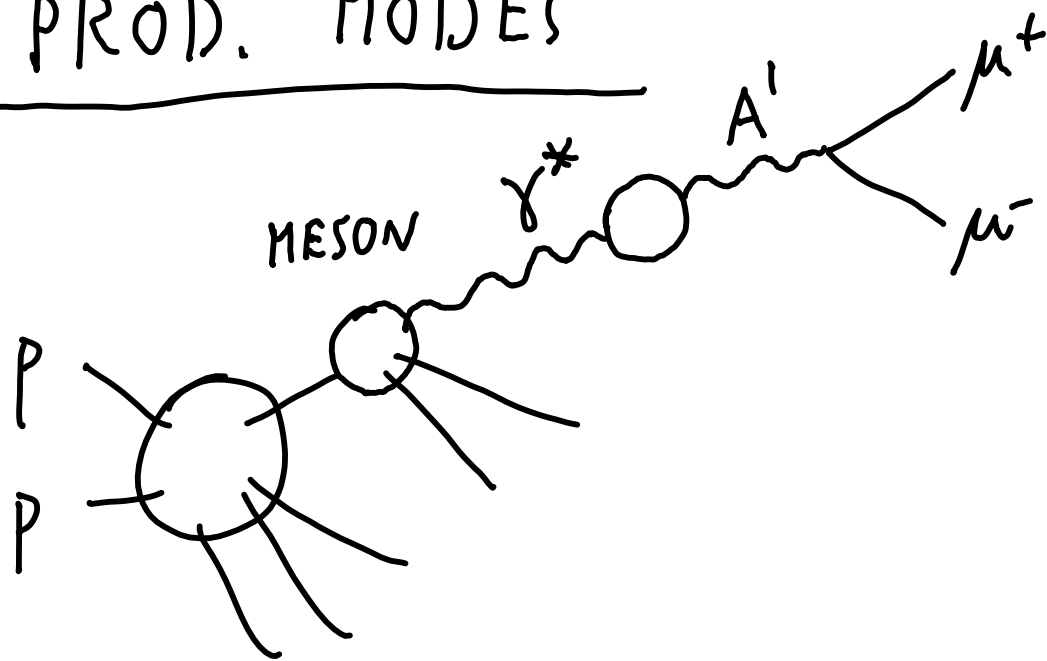


- Dark photon ( $A'$ ) mediates dark matter ( $\chi$ ) interaction
  - If  $m(A') > 2m_\chi$  then **invisible decay**  $A' \rightarrow \chi\chi$  dominant
  - If  $m(A') < 2m_\chi$  then **visible decay**  $A' \rightarrow \ell^+\ell^-$  dominant
- Production from mixing with virtual photon
  - Can oscillate to a dark photon with probability  $\epsilon^2$
- Dark photon lifetime proportional to  $1/(\epsilon^2 m_{A'})$ 
  - Light, rarely produced dark photons are displaced

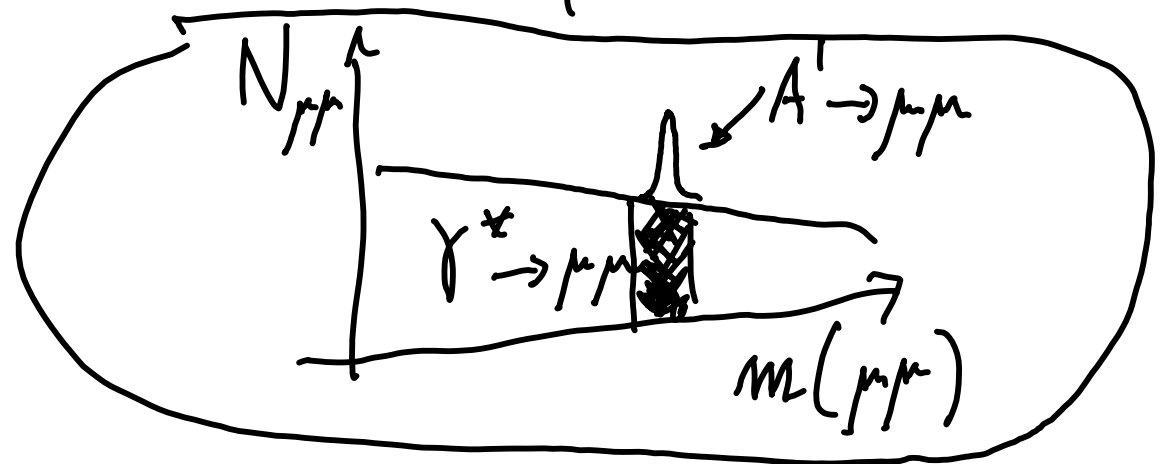


# Visible $A' \rightarrow \text{LHCb}$

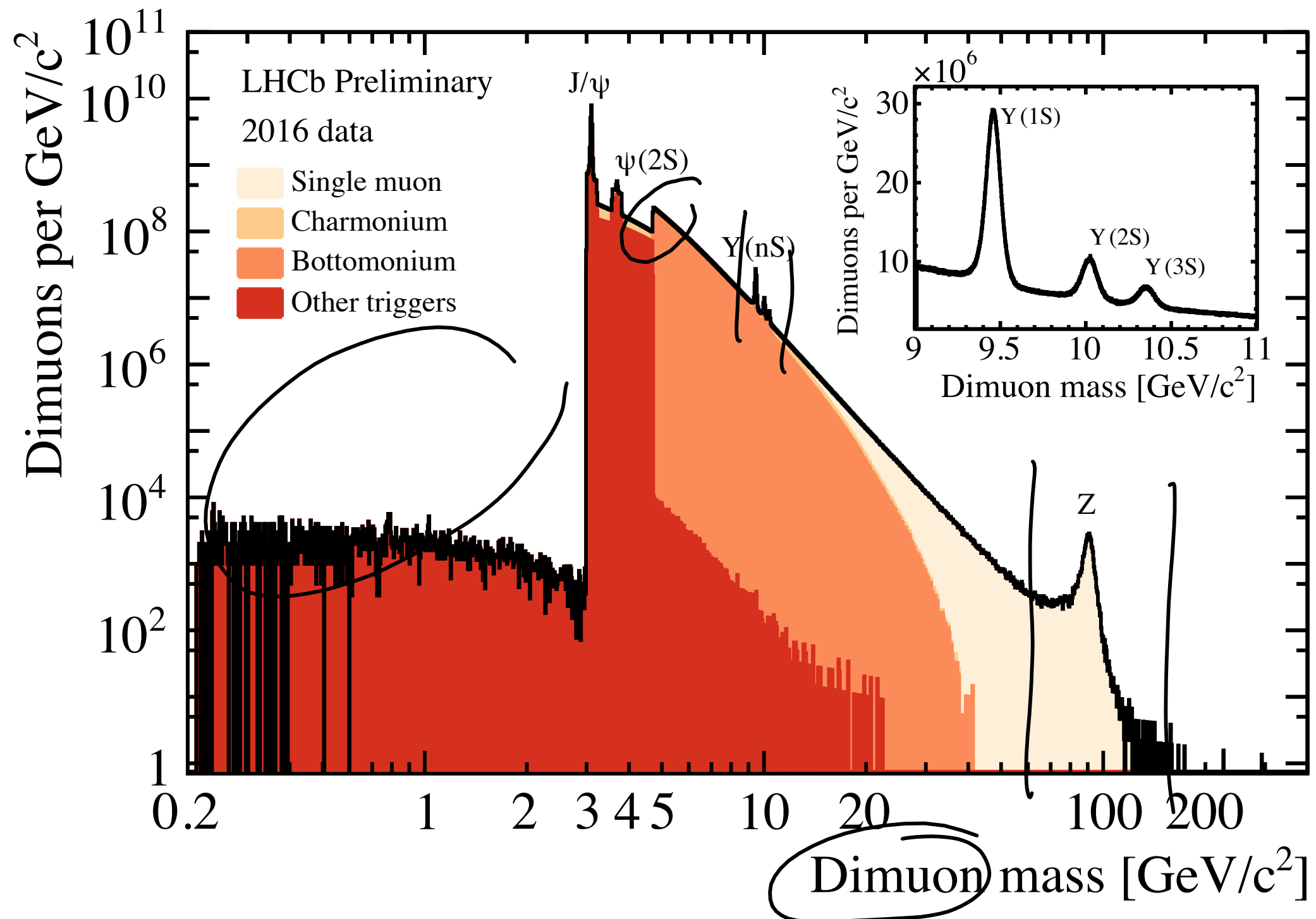
2 PROD. MODES



$$\text{expected } N^{A'}(m_{A'}, \varepsilon) = N^{\gamma^*}(m_{A'}) \cdot \varepsilon^2 \frac{\text{eff}(A' \rightarrow \mu\mu)}{\text{eff}(\gamma^* \rightarrow \mu\mu)}$$

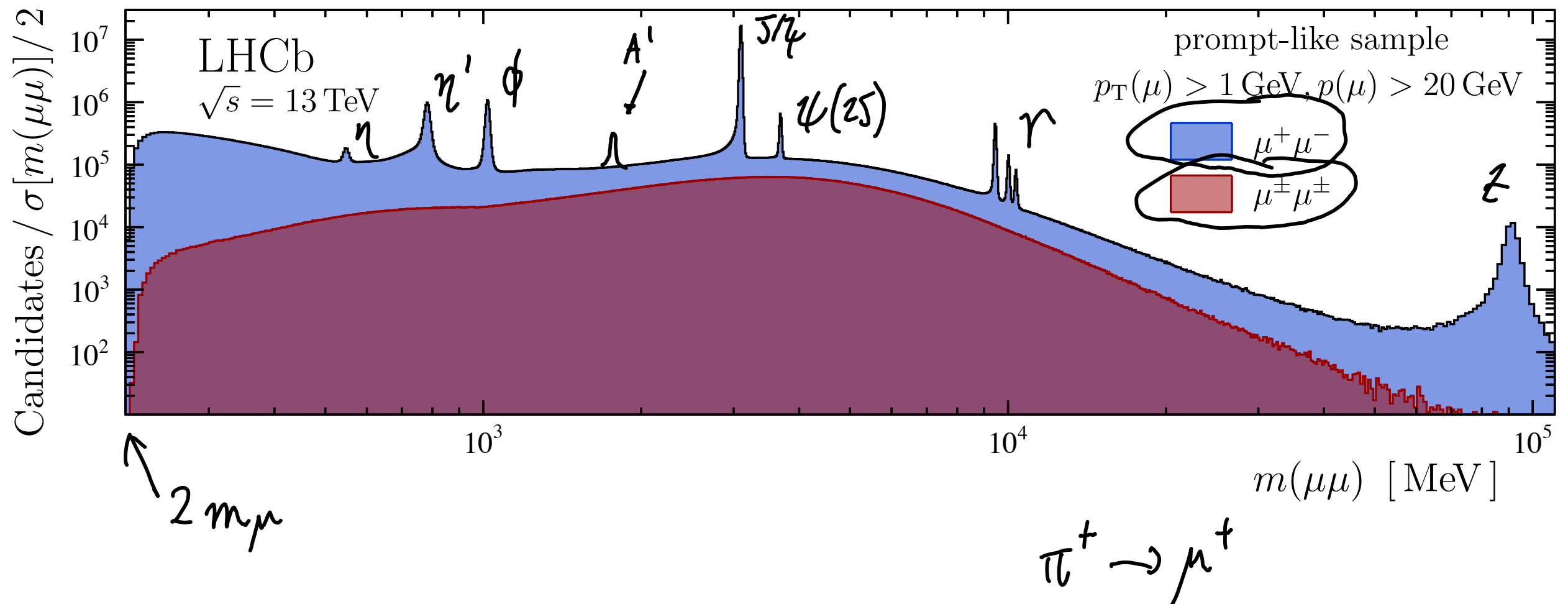


# Prompt $\mu\mu$ trigger



# Prompt $\mu\mu$ trigger

- Removed prescale thanks to real-time alignment and reduced event size (turbo stream)

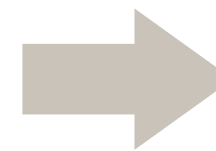


# Measuring $N\gamma^*$

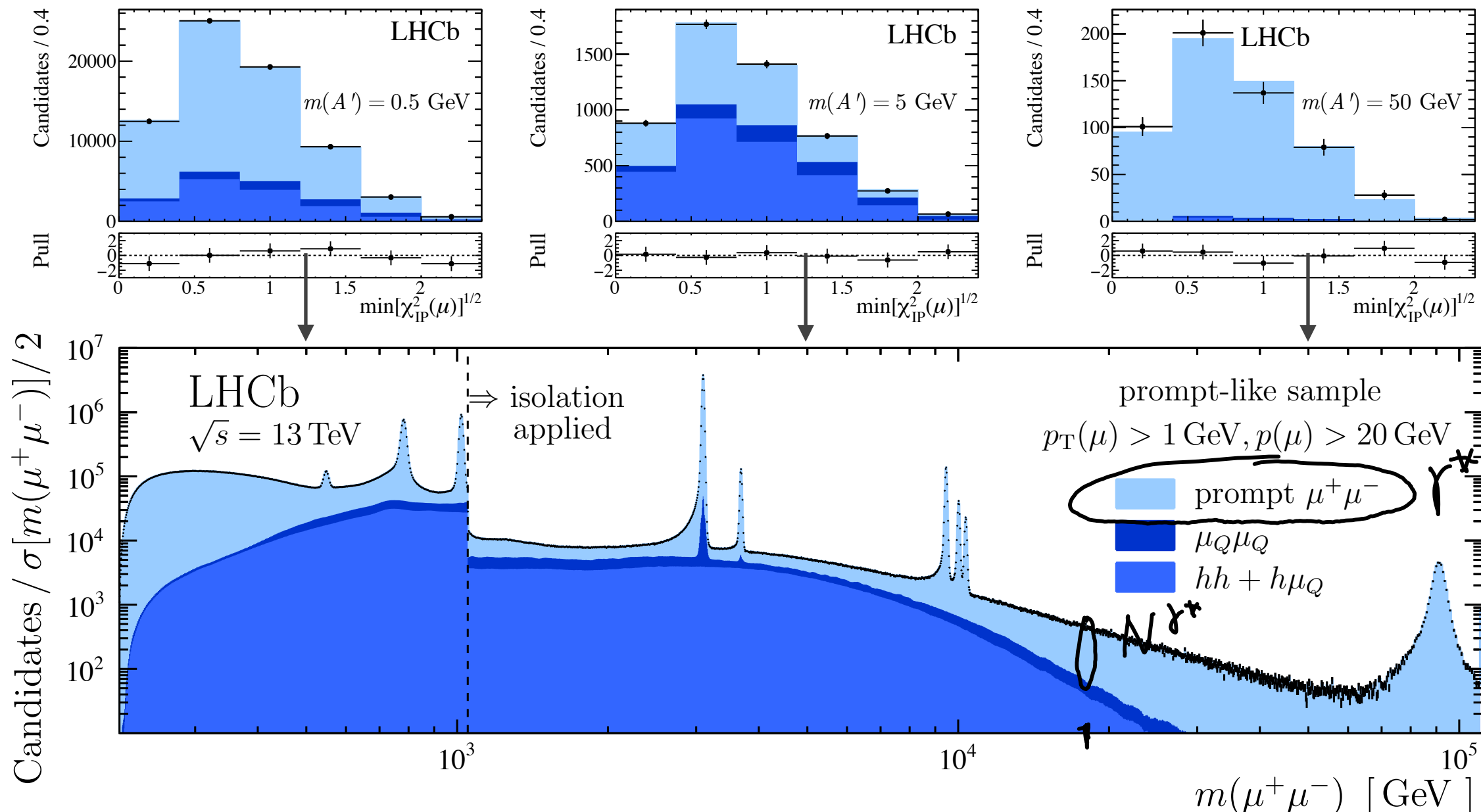
*BKG*

- prompt  $\mu^+\mu^-$
- $\mu_Q\mu_Q$
- $hh + h\mu_Q$

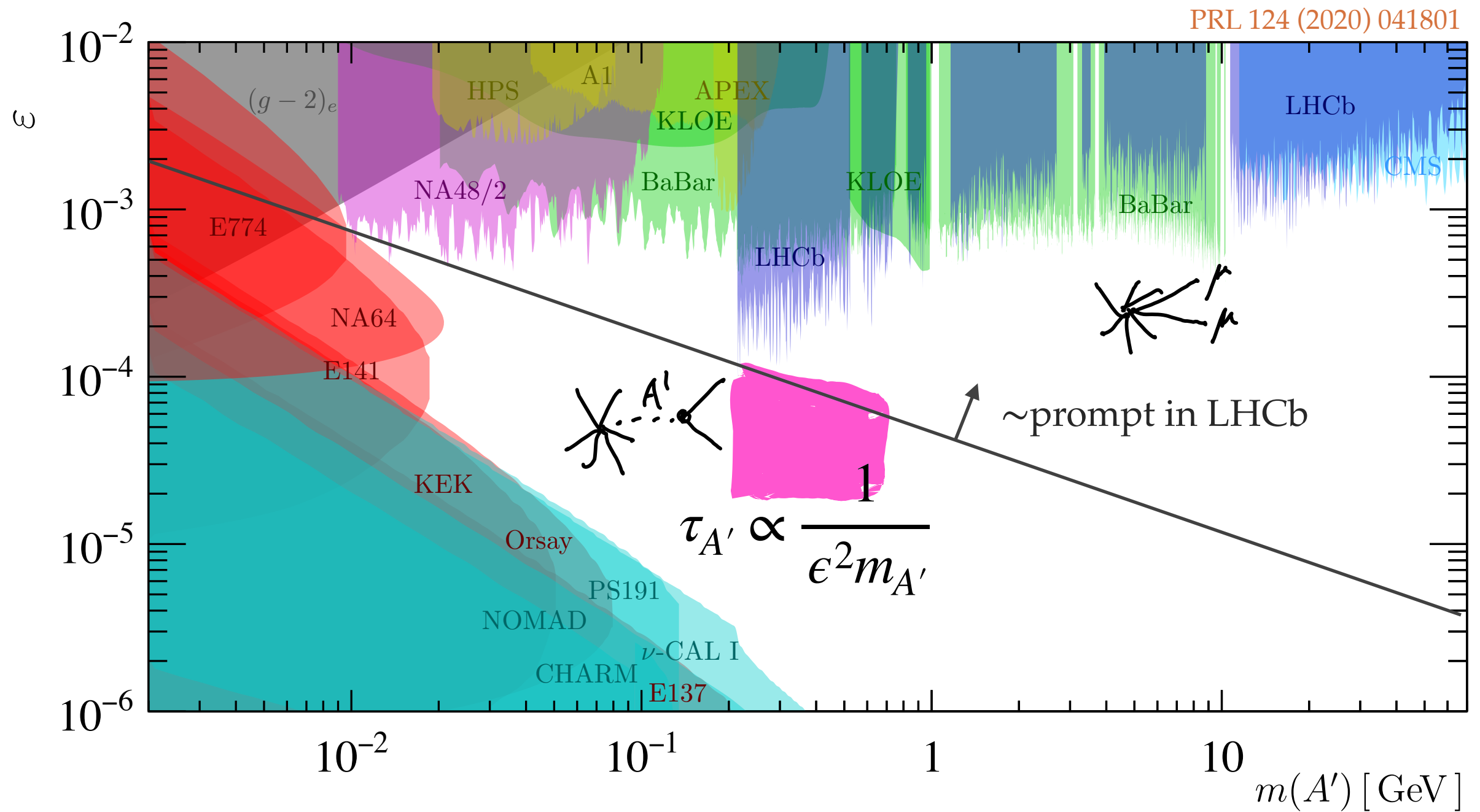
→ the  $N\gamma^*$  we want to measure  
 → muons from heavy flavour decays  
 → mis-identified pions as muons



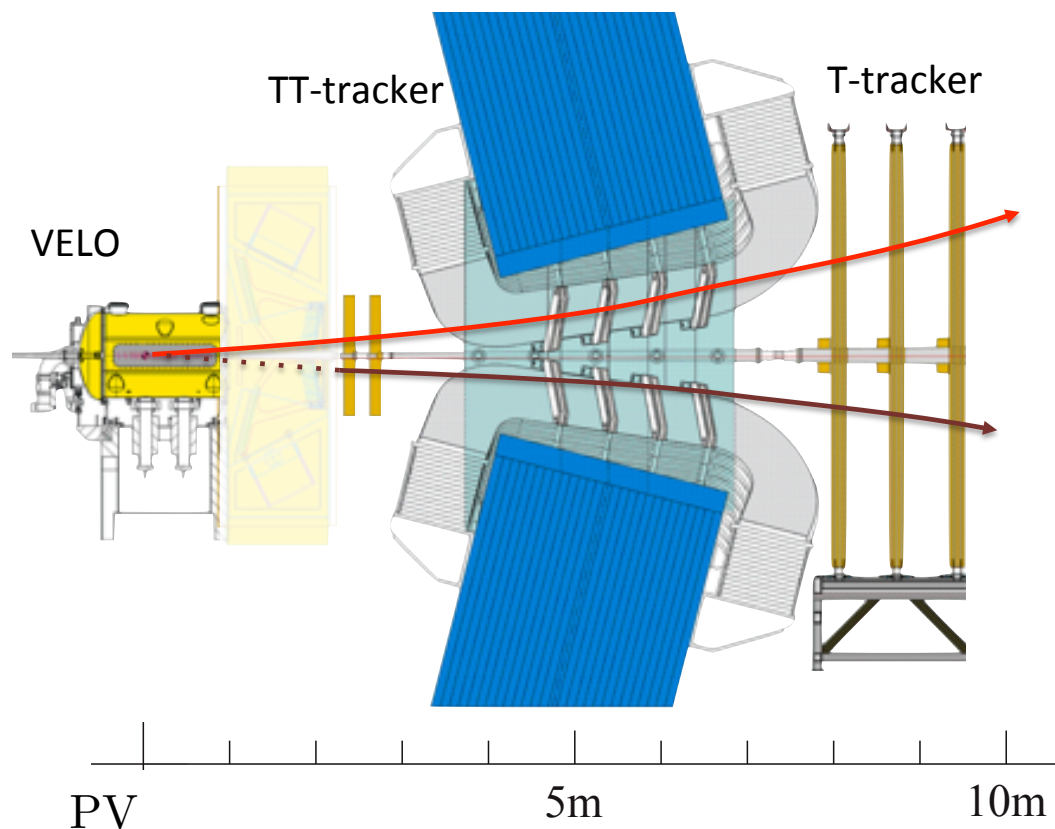
Using templates  
 for  $\min[\chi^2_{\text{IP}}]$   
 (small mass dep)



# Visible $A'$ limits

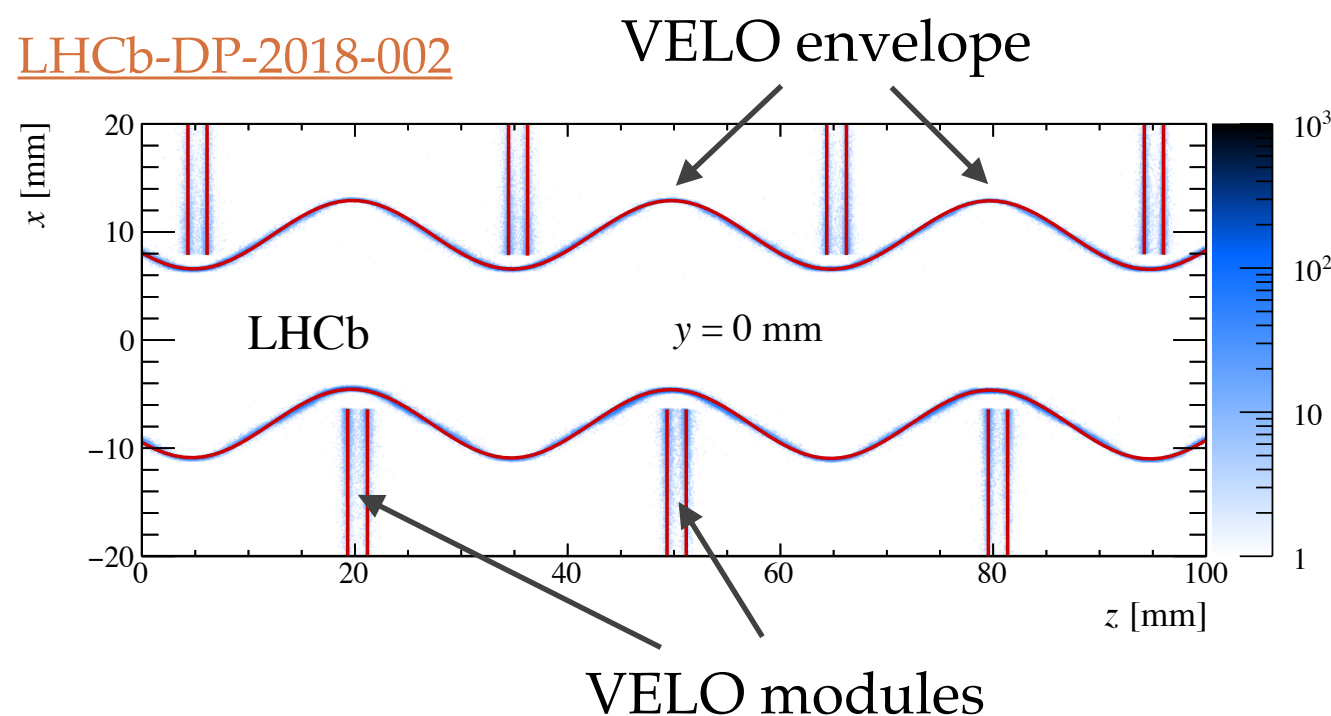


# Displaced Vertices at LHCb



- Currently only **within VELO**
  - Displacement  $< 20$  cm (but with boost)
- Could extend to **upstream tracker**
  - Displacement  $< 200$  cm
  - Worse vertex and  $p$  resolution ( $m(\pi\pi)$  resolution  $2\times$  larger)
  - Not yet available in trigger

[LHCb-DP-2018-002](#)



## Backgrounds in VELO

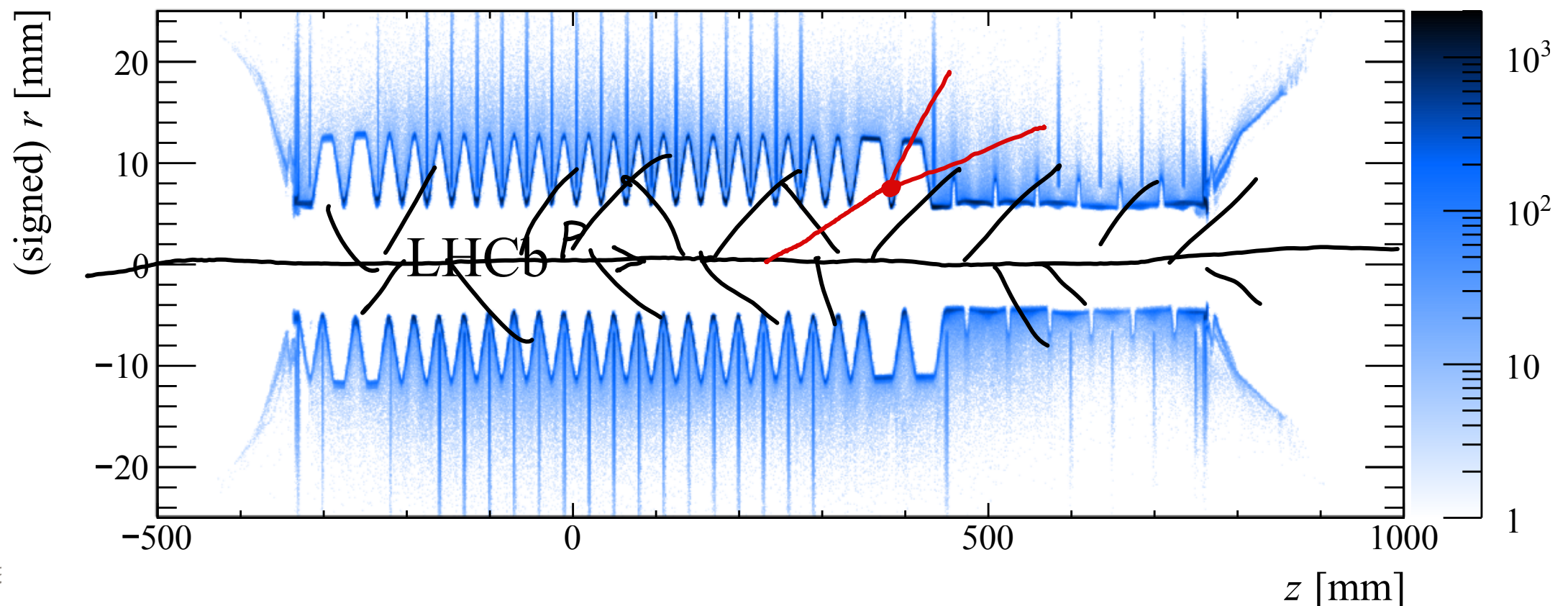
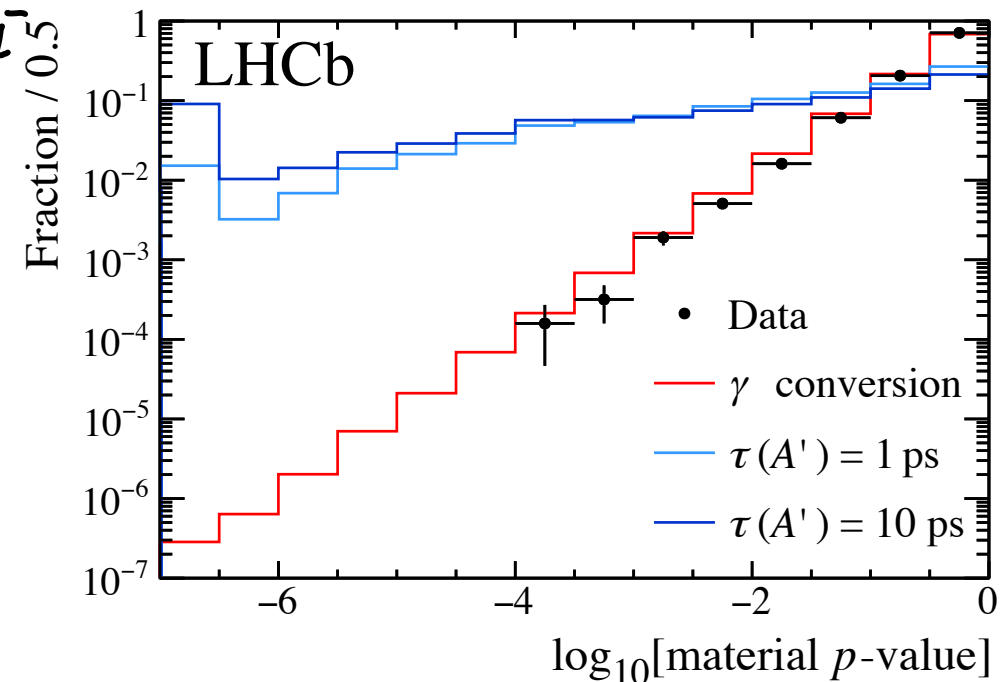
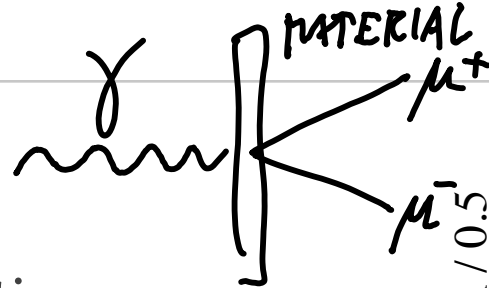
- Heavy Flavour displaced decays
  - $\tau(B) \sim 1.5$  ps,  $\beta\gamma \sim 10 \Rightarrow$  few mm
- Thin VELO envelope
  - $< 5$  mm: background mainly from heavy-flavour background
  - $> 5$  mm: background mainly from material interaction

# VELO material map

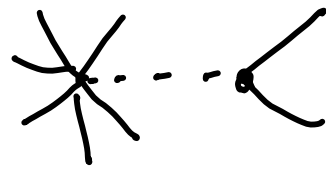
LHCb-DP-2018-002

## ● VELO material map

- based on material interactions from hadrons produced in beam-gas collisions
- Assign probability of material interaction hypothesis to each  $\mu\mu$



# Displaced dark photon

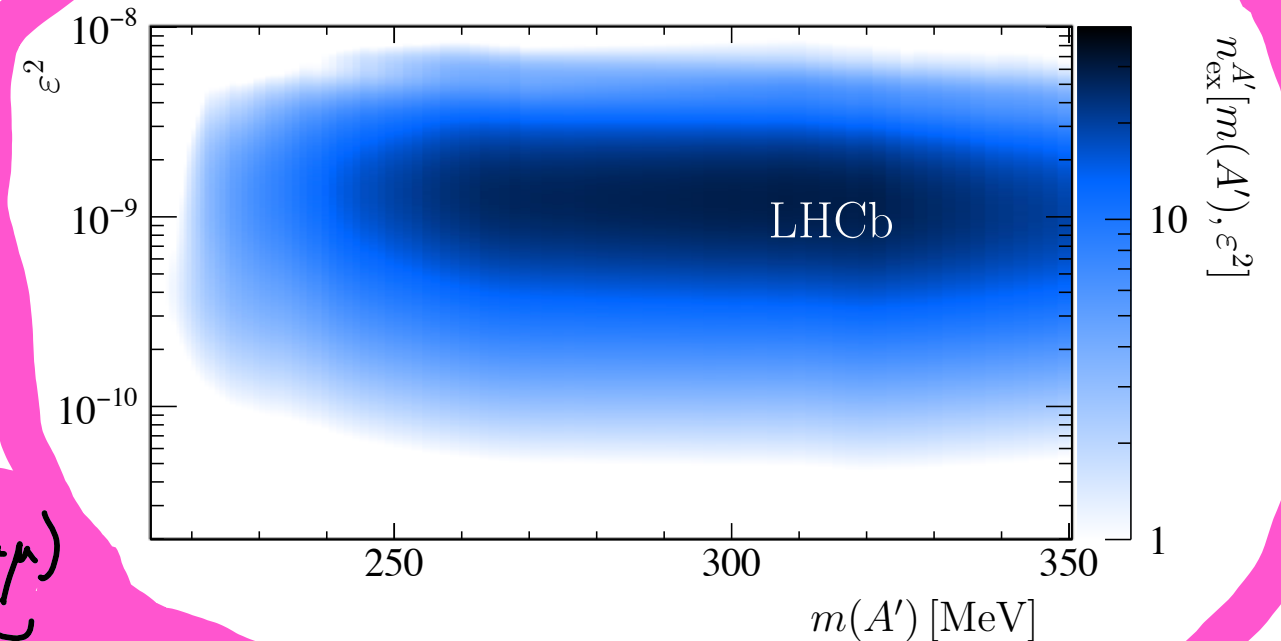


$$B \rightarrow \mu\mu \quad 10^{-9}$$

## ● Much lower background

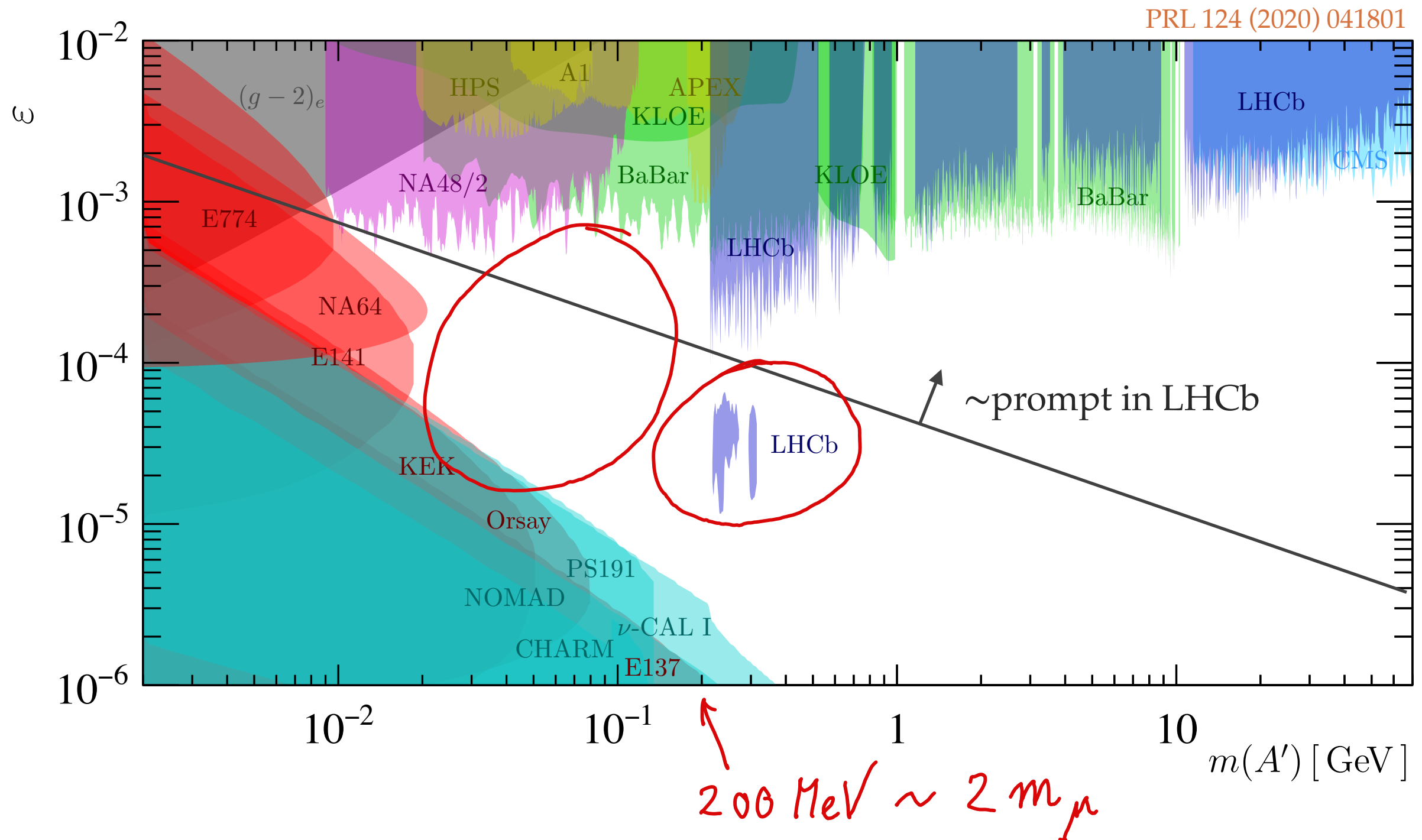
- No long-lived SM particle decaying to  $\mu\mu$
- Main background from  $\gamma \rightarrow \mu\mu$  conversions in the VELO vetoed

$$N^{A'}(m_{A'}, \epsilon) = N^{\gamma^*}(m_{A'}) \cdot \epsilon^2 \frac{\text{eff}(A' \rightarrow \mu\mu)}{\text{eff}(\gamma^* \rightarrow \mu\mu)}$$

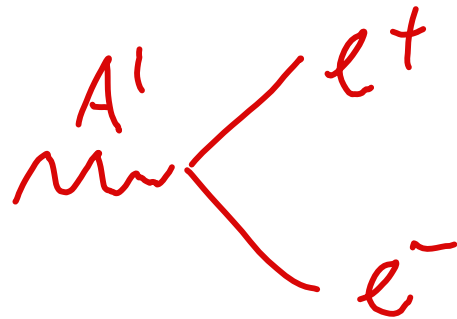


SIMULATION

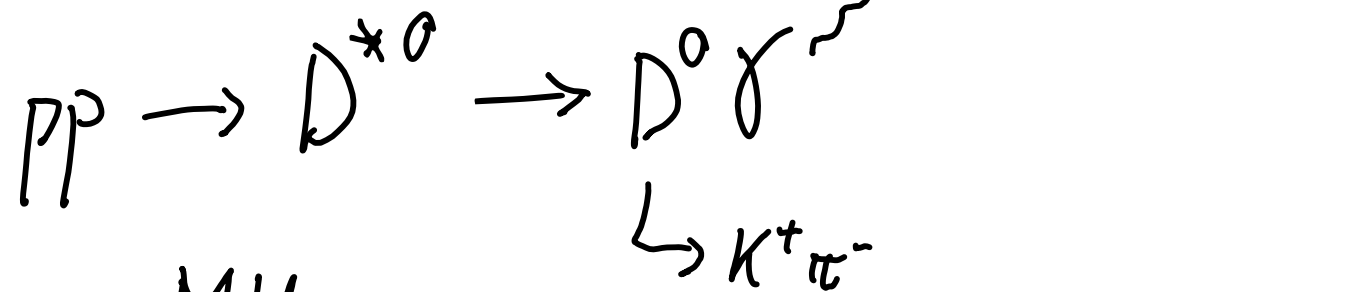
# Visible $A'$ limits



# Dark Photons below $2m_\mu$

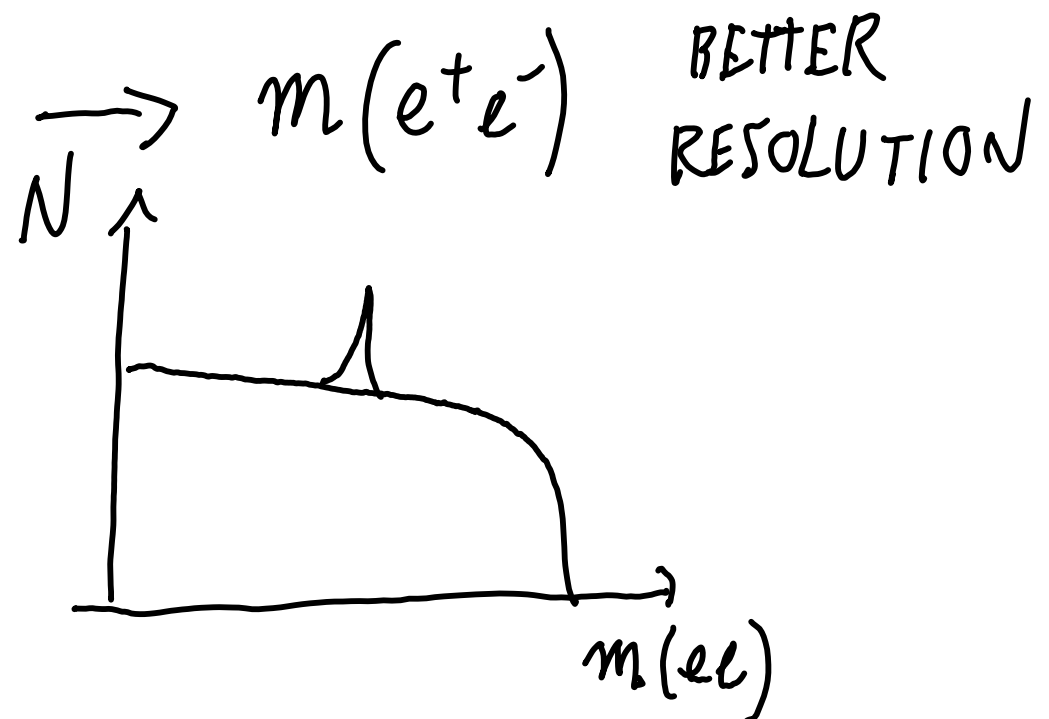


EXPERIM. TRICK



$$\text{Rate}(pp \rightarrow D^{*0} \rightarrow D^0 \gamma)_{\text{LHCb}} = 0.7 \text{ MHz}$$

CONSTRAINT  $m(D^{*0})$   $m(D^0)$



# Dark Photons below $2m_\mu$

