

# CRNS - new insights from modeling soil moisture measurement at the hectometer scale

07.10.2021 Markus Köhli<sup>1,2</sup>, Jannis Weimar<sup>1</sup>, Martin Schrön<sup>3</sup> and Ulrich Schmidt<sup>1</sup>

- <sup>1</sup> Physikalisches Institut, Heidelberg University, Heidelberg, Germany
- <sup>2</sup> Physikalisches Institut, University of Bonn, Bonn, Germany
- <sup>3</sup> Helmholtz Centre for Environmental Research UFZ, Leipzig, Germany



Physikalisches Institut Heidelberg University Germany



**DFG FOR 2394** 

Sens





Many studies were carried out for **finding a sensor calibration routine** and to compare the performance to conventional instruments... Rivera Villarreyes et al., 2011; Franz et al., 2012a; Hawdon et al., 2014; Almeida et al., 2014; Coopersmith et al., 2014

with a good agreement between measured neutron flux and soil moisture values



 $\Phi$ 



However, unexplained features in the CRNS data could not be described by the Desilets equation

$$\theta(N) = \frac{0.0808}{\left(\frac{N}{N_0}\right) - 0.372} - 0.115$$

Then,

groups tried to **fit the parameters** of the hyperbola to their data Rivera Villarreyes et al., 2011; Lv et al., 2014; Heidbüchel et al., 2016; Sigouin and Si, 2016

with a better correlation at the cost of site-specific calibrations





One reason: different weighting strategy needed

Footprint revision
 Köhli et al. 2015; Schrön et al. 2017











R. ROSOLEM, et al. "The Effect of Atmospheric Water Vapor on Neutron Count in the Cosmic-Ray Soil Moisture Observing System." *Journal of Hydrometeorology* 14(5), 2013, 1659–1671.





5

## Cosmic NS > Different Intensity Models







J. IWEMA et al. "Investigating temporal field sampling strategies for site-specific calibration of three soil moisture–neutron intensity parameterisation methods." *Hydrology and Earth System Sciences* 19(7), 2015, 3203–3216.



# Cosmic NS > Above-ground neutron flux

### Epithermal flux is not the result from one single process







#### Take into account the relative detection efficiency







Understand the mathematics of neutron transport:



### NS> Theoretical considerations

The neutron intensity is proportional to the number of collisions within the water-sensitive domain





NS> Theoretical considerations

The neutron intensity is proportional to the number of collisions within the water-sensitive domain

$$I(\theta) \propto \Phi_{\text{epith}}(\text{H}_2\text{O}) \propto n_{\text{col}}^{\text{ground}} \propto \frac{\sum^{\text{soil}} + w \sum^{\text{water}}}{\sum^{\text{soil}} \xi^{\text{soil}} + w \sum^{\text{water}} \xi^{\text{water}}}$$

$$(10)$$

Hyperbola-shaped



NS> Theoretical considerations

The neutron intensity is proportional to the number of collisions within the water-sensitive domain

 $\Phi \propto \exp(-r)/r$ 

$$I(\theta) \propto \Phi_{\text{epith}}(\text{H}_2\text{O}) \propto n_{\text{col}}^{\text{ground}} \propto \frac{\sum^{\text{soil}} + w \sum^{\text{water}}}{\sum^{\text{soil}} \xi^{\text{soil}} + w \sum^{\text{water}} \xi^{\text{water}}}$$
Hyperton the second s

Hyperbola-shaped

exponentially shaped





Our proposition: the universal transport solution (UTS)

$$I(\theta, h) = N_D \left(\frac{p_1 + p_2 \theta}{p_1 + \theta}\right)$$

$$+e^{-p_3\theta}$$



Our proposition: the universal transport solution (UTS)

$$I(\theta,h) = N_D \left(\frac{p_1 + p_2 \theta}{p_1 + \theta} \left(p_0 + p_6 h + p_7 h^2\right) + e^{-p_3 \theta} \left(p_4 + p_5 h\right)\right)$$



Our proposition: the universal transport solution (UTS)

$$I(\theta, h) = N_D \left( \frac{p_1 + p_2 \theta}{p_1 + \theta} \left( p_0 + p_6 h + p_7 h^2 \right) + e^{-p_3 \theta} (p_4 + p_5 h) \right)$$
  
Hyperbola  
But the second state of the second

# Cosmic NS > Simulating the above-ground intensity

The simulation toolkits used in this study are MCNP 6.2 and URANOS. The air medium consists of 78% nitrogen, 21% oxygen and 1% argon usually at a pressure of 1020 mbar. The soil consists of 50%Vol solids and a scalable amount of H2O. The solid domain is comprised of 75%Vol SiO<sub>2</sub> and 25%Vol Al<sub>2</sub>O<sub>2</sub> at a compound density of 2.86 g/cm<sup>3</sup>. The input spectrum used in this work relies on the cosmic-ray propagation models by Sato et al.







-N0 method is not steep enough in dry regions

- Exponential methods like Cosmic are more suitable for low soil moisture values - Using fixed lower and upper bounds for scoring leads to a significantly steeper soil moisture relation. The detector response icludes thermal and MeV neutrons, both scaling less good then the epithermal range.





-N0 method is not steep enough in dry regions

 Exponential methods like Cosmic are more suitable for low soil moisture values
 Using fixed lower and upper bounds for scoring leads to a significantly steeper soil moisture relation. The detector response icludes thermal and MeV neutrons, both scaling less good then the epithermal range. For humid compared to dry air the maximum achievable count rate is reduced by 20 %. This quantitatively agrees with Rosolem et al. (2013) who studied the change from dry to 22 g/m<sup>3</sup>. However, a strictly linear relationship for water vapor cannot be verified. The presented reduction rate of 0.0054 per gram air humidity seems to hold only for dry conditions.

### NS> Experimental evidence – dry conditions

M. Köhli et al. "Soil Moisture and Air Humidity dependence of the above-ground cosmic-ray neutron intensity" Front. Water, 2, 2021.



#### COSMOS Santa Rita ( $\theta \approx 7 \pm 3$ %, $h \approx 7 \pm 4$ g/cm<sup>3</sup>), daily resolution

Caution! Here: soil moisture converted to neutron counts

Data from Santa Rita (T. Franz)

### NS> Experimental evidence – dry conditions

M. Köhli et al. "Soil Moisture and Air Humidity dependence of the above-ground cosmic-ray neutron intensity" Front. Water, 2, 2021.

#### Zoom-In: air humidity scaling



### NS> Experimental evidence – wet conditions



Data from Rollesbroich (R. Baatz)

### **NS>** Experimental evidence – wet conditions

#### TERENO Fendt ( $\theta \approx 45 \pm 10$ %,h $\approx 8 \pm 3$ g/cm<sup>3</sup>), daily resolution



Data from Fendt (B. Fersch, KIT)

### Cosmic NS > Experimental evidence – wet conditions



TERENO Rollesbroich ( $\theta \approx 42 \pm 5$  %, h  $\approx 7 \pm 3$  g/cm<sup>3</sup>), daily resolution

Data from Rollesbroich (R. Baatz, FZJ)



#### 1.1 relative intensity MCNP6 full + Detector Response 1 MCNP6 full + Energy Window 0.9 URANOS + Detector Response 0.8 Nomethod 2010 COSMIC 2013 ----0.7 0.6 0.5 0.4 1 g/m<sup>3</sup>- 35 g/m <sup>3</sup> 0.3 0.2 30 50 10 20 40 0 soil moisture [%]



Cosmic NS> Thermal Shielding







#### Gadolinium Coating

approx. 90 % absorption efficiency to cadmium cutoff (0.5 eV) equals ~95 % in thermal regime ( < 0.1 eV)





• Site-specific calibrations are most probably a result of not optimal fit functions





- Site-specific calibrations are most probably a result of not optimal fit functions
- New CRNS intensity function found by understanding the contributions and improving the corrections of CR neutrons to the signal.

Combined function for soil moisture and humidity:

$$I(\theta,h) = N_D \left(\frac{p_1 + p_2 \theta}{p_1 + \theta} \left(p_0 + p_6 h + p_7 h^2\right) + e^{-p_3 \theta} \left(p_4 + p_5 h\right)\right)$$





- Site-specific calibrations are most probably a result of not optimal fit functions
- New CRNS intensity function found by understanding the contributions and improving the corrections of CR neutrons to the signal.

Combined function for soil moisture and humidity:

$$I(\theta,h) = N_D \left(\frac{p_1 + p_2 \theta}{p_1 + \theta} \left(p_0 + p_6 h + p_7 h^2\right) + e^{-p_3 \theta} \left(p_4 + p_5 h\right)\right)$$

First data blind validation very successfull

M. Köhli et al. "Soil Moisture and Air Humidity dependence of the above-ground cosmic-ray neutron intensity" Front. Water, 2, 2021.





- Site-specific calibrations are most probably a result of not optimal fit functions
- New CRNS intensity function found by understanding the contributions and improving the corrections of CR neutrons to the signal.

Combined function for soil moisture and humidity:

$$I(\theta,h) = N_D \left(\frac{p_1 + p_2 \theta}{p_1 + \theta} \left(p_0 + p_6 h + p_7 h^2\right) + e^{-p_3 \theta} \left(p_4 + p_5 h\right)\right)$$

• First data blind validation very successfull

M. Köhli et al. "Soil Moisture and Air Humidity dependence of the above-ground cosmic-ray neutron intensity" Front. Water, 2, 2021.

• Thermal shields can significantly improve the signal response





# Backup



Set	$p_0$	$p_1$	$p_2$	$p_3$	$p_4$	$p_5$	$p_6$	$p_7$
1: MCNP drf (full)	1.0940	0.0280	0.254	3.537	0.139	-0.00140	-0.0088	0.0001150
2: MCNP ewin (full)	1.2650	0.0259	0.135	1.237	0.063	-0.00021	-0.0117	0.0001200
3: URANOS drf	1.0240	0.0226	0.207	1.625	0.235	-0.00290	-0.0093	0.0000740
4: URANOS ewin	1.2230	0.0185	0.142	2.568	0.155	-0.00047	-0.0119	0.0000920

Results for the parameters p0 - p7

Intercomparison

