

g) Neutrino Masses

* In the SM neutrinos are exactly massless & occur only as ν_L (left-handed) within doublet $\begin{pmatrix} \nu_L \\ e_L \end{pmatrix}$.

* Neutrino oscillation experiments (mid 1990s) suggest: mass differences for neutrinos:

In mass eigenstates:

$$\Delta m_{12}^2 \sim 10^{-5} \text{ eV}^2, \quad |\Delta m_{23}^2| \sim 10^{-3} \text{ eV}^2$$

Note: only mass differences known, no absolute numbers

* Modelling neutrino masses requires introduction of a right-handed neutrino ν_R (at least a most straightforward model)

* Simplest scenario: mimic mass generation for up-quarks

$$L_N = \begin{pmatrix} \nu_L \\ e_L \end{pmatrix}_N, \quad R_{N-} = (e_R)_N, \quad R_{N0} = (\nu_R)_N$$

$$\mathcal{L}_{\text{left, II}} = -\sqrt{2} \left(\bar{L}_N f_{NM} \phi R_{N-} + \bar{L}_N f_{NM}^0 \phi^c R_{N0} + \text{h.c.} \right)$$

$$U(1)_Y : Y(\bar{L}_N) = +\frac{1}{2}, \quad Y(\phi^c) = -\frac{1}{2}$$

$$\rightarrow Y(R_{N0}) = 0 \quad \Leftrightarrow (\nu_R)_N \text{ completely neutral}$$

These masses are of the form:

$$\mathcal{L} = -(\bar{\nu}_L)_N (M_D)_{NM} (\nu_R)_N + \text{h.c.} \quad M_D: \text{Dirac mass matrices}$$

* Since ν_R is neutral, there exists another possible mass bilinear:

As for quarks, Dirac masses introduces mixing in family space in charged current

$$J^\mu = 2 \bar{L}_N \gamma^\mu \sigma_+ L_N$$

$$= 2 \left(\bar{\nu}_\tau \quad \bar{\nu}_\mu \quad \bar{\nu}_e \right)_L \gamma^\mu \begin{pmatrix} \nu' \\ \mu' \\ e' \end{pmatrix}_L$$

in terms of weak eigenstates

Diagonalizing Dirac mass terms for neutrinos & lepton masses
 \rightarrow mass eigenstates $\left(\nu_\tau, \nu_\mu, \nu_e \right)$ } as for quarks
 $\left(\tau, \mu, e \right)$

$$J^\mu = \left(\bar{\nu}_\tau, \bar{\nu}_\mu, \bar{\nu}_e \right) \gamma^\mu (1 + \gamma_5) \cdot V_{\text{left}} \begin{pmatrix} \tau \\ \mu \\ e \end{pmatrix}$$

* V_{left} governs Family mixing of neutrinos & can be measured in neutrino oscillation experiments.

* Like the CKM matrix V_{left} introduces CP violation. If only Dirac masses (as above) then: V_{left} has 3 angles & 1 phase — just as V_{CKM} .

2 Challenges:

a) Naturalness:

Why are the Dirac masses for neutrinos smaller by factor of $\sim 10^{-6}$ than those for down-type quarks?
 (Note however: only mass differences known experimentally)

b) There exists another possibility for neutrino masses — Majorana masses. One would need a mechanism to suppress / prevent them if neutrino masses are (only) of Dirac type

* Since ν_R is neutral, there exists another possible mass bilinear — the Majorana mass term:

Recall: For Dirac spinor $\psi(x)$:

$$\psi^c(x) = C \bar{\psi}^T(x) = C \gamma^0 \psi^*(x)$$

$$\overline{\psi^c}(x) = -\psi^T C^{-1} = \psi^T C^*$$

$$(\psi_R)^c = \frac{1}{2}(1 - \gamma^5)\psi^c = (\psi^c)_L$$

$$(\overline{\psi_R})^c = \overline{\psi^c} \cdot \frac{1}{2}(1 + \gamma^5)$$

Therefore: $(\overline{\psi_R})^c \cdot \psi_R \neq 0$

Note: If ψ_R carries charge, the $(\overline{\psi_R})^c \psi_R$ is not gauge invariant because

$$(\overline{\psi_R})^c \cdot \psi_R = \psi_R^T C^* \psi_R$$

Consider a family of neutrinos:

⇒ Majorana mass: $\mathcal{L}_M = -M_M (\overline{\nu_R^c} \nu_R + \overline{\nu_R} \nu_R^c)$

Dirac mass: $\mathcal{L}_D = -M_D (\overline{\nu_R} \nu_L + \overline{\nu_L} \nu_R)$

Consider a family of neutrinos.

Define: $\chi = \nu_L + i\nu_L^c \Rightarrow \chi^c = \chi \iff$ Majorana spinor

$\omega = \nu_R + i\nu_R^c \quad \omega^c = \omega$

then: $\mathcal{L}_M = -M_M \overline{\omega} \omega$ and

$$\begin{aligned} \mathcal{L}_{\text{mass}} &= -M_D (\overline{\nu_R} \nu_L + \overline{\nu_L} \nu_R) - M_M (\overline{\nu_R^c} \nu_R + \overline{\nu_R} \nu_R^c) \\ &= -(\overline{\chi}, \overline{\omega}) \begin{pmatrix} 0 & \frac{1}{2}M_D \\ \frac{1}{2}M_D & M_M \end{pmatrix} \begin{pmatrix} \chi \\ \omega \end{pmatrix} \end{aligned}$$

Note:

$$\frac{1}{2}(1-\gamma_5)\chi = \frac{1}{2}(1-\gamma_5)v_L \equiv v_L$$

$$\frac{1}{2}(1+\gamma_5)\omega \equiv v_R$$

$$\frac{1}{2}(1-\gamma_5)\omega \equiv v_R^c$$

* Diagonalise:

eigenvalues:

$$m_{1,2} = \frac{1}{2} \left(M_m \pm \sqrt{M_m^2 + M_D^2} \right)$$

eigenvectors:

$$\eta_1 = \cos \Theta \chi - \sin \Theta \omega \quad \leftrightarrow \text{lower sign}$$

$$\eta_2 = \sin \Theta \chi + \cos \Theta \omega \quad \leftrightarrow \text{upper sign}$$

$$\tan(2\Theta) = -M_D/M_m$$

→ For general M_D, M_m get admixture of χ/ω
 ~ mass eigenstates.

* See-saw mechanism:

To explain smallness of observed mass eigenvalues we take:

$$\left. \begin{array}{l} M_D \sim \text{order electro-weak masses} \\ M_m \sim \lesssim 10^{15} \text{ GeV} \end{array} \right\} M_D \ll M_m$$

$\leftarrow 10^3 \text{ GeV}$

→ $|m_1| \approx \frac{1}{4} \cdot \frac{M_D^2}{M_m} \sim \text{eV} \quad \Rightarrow \quad \eta_1 \approx \chi = v_L + (v_L^c)$

$$m_2 \approx M_m \quad \Rightarrow \quad \eta_2 \approx \omega = v_R + (v_R^c)$$

i.e.:

$$v_L = \frac{1}{2}(1-\gamma_5)\chi \approx (\eta_1)_L \quad \text{where } \eta_2 \text{ of small mass } \sim \frac{M_D^2}{M_m}$$

$$v_R = \frac{1}{2}(1+\gamma_5)\omega \approx (\eta_2)_R \quad \leftrightarrow \text{heavy eigenstate}$$

Note:

$M_H = 10^{15}$ GeV is upper limit in order to
eV - scale (light) mass eigenvalue

$$\frac{M_D^2}{M_H} \quad \text{for } M_D \leq \Theta \left(\begin{array}{l} \text{quark} \\ \text{mass} \end{array} \right)$$

Recall: $m_{\text{top}} \sim 170$ GeV $\Rightarrow M_D \leq 10^3$ GeV

$$\Rightarrow \frac{M_D^2}{M_H} \sim \frac{10^6}{10^{15}} \text{ GeV} \sim 10^{-9} \text{ GeV} \checkmark$$

More natural is $M_D \approx 10 - 10^2$ GeV

$$\Rightarrow M_H \sim 10^{11} \sim 10^{15} \text{ GeV}$$

Compare: $M_{\text{Planck}} = 10^{19}$ GeV

$M_{\text{GUT}} = 10^{16}$ GeV

\Rightarrow Majorana mass scale is 'New Physics'
& begs for theoretical explanation

In particular: Why is $M_H \neq M_{\text{GUT}}$?

\leadsto Beyond Standard Model Physics

Generalisations to more families

* Assume 3 families of ν_L^i as in $\begin{pmatrix} \nu_L^i \\ e_L^i \end{pmatrix}_i$ $i=1,2,3$
 ← EW eigenstates

* Add m copies of $(\nu_R^i)_\alpha$ $\alpha=1, \dots, m$
 $(\nu_R^i)_\alpha$: SM singlets (as above)

→ most natural choice: $m=3$

→ $m=2$ possible if only 2 neutrino families massive
 (Recall: only $\Delta m_{12}^2, |\Delta m_{13}^2|$ known)

→ $m \geq 4$: consistent → extended models w/ extra sterile (singlet) neutrinos

* Introduce: $\vec{\chi}^i \equiv (\nu_L^i + (\nu_L^i)^c)_i$ $i=1,2,3$

$\vec{\omega}^\alpha \equiv (\nu_R^\alpha + (\nu_R^\alpha)^c)_\alpha$ $\alpha=1, \dots, m$

$$\mathcal{L}_{\text{mass}} = - \begin{pmatrix} \vec{\chi}^i, \vec{\omega}^\alpha \end{pmatrix} \begin{pmatrix} 0 & \frac{1}{2} M_D \\ \frac{1}{2} M_D & M_M \end{pmatrix} \begin{pmatrix} \vec{\chi}^i \\ \vec{\omega}^\alpha \end{pmatrix}$$

\mathcal{H} : symmetric, but complex

Diagonalise: $\mathcal{H} = \mathcal{V}^T \tilde{\mathcal{H}}_D \mathcal{V}$ \mathcal{V} : unitary

$$\tilde{\mathcal{H}}_D = \begin{pmatrix} \varepsilon^2 & \mathcal{H}_{\nu_L} & 0 \\ 0 & & \mathcal{H}_{\nu_R} \end{pmatrix} \quad \varepsilon: \text{parameter} \ll 1$$

$(\mathcal{O}(M_D/M_M))$

$$\mathcal{V} = \begin{pmatrix} U_{11} & \varepsilon U_{12} \\ \varepsilon U_{21} & U_{22} \end{pmatrix} \quad \mathcal{V}^T = \mathcal{V}^{-1}$$

Charged current:

$$\vec{\nu}_L^i = \frac{1}{2}(1 - \gamma_5) \vec{\chi}^i = \frac{1}{2}(1 - \gamma_5) \begin{pmatrix} 1 & 0 \\ 0 & 0 \end{pmatrix} \mathcal{V}^{-1} \begin{pmatrix} \vec{\chi} \\ \vec{\omega} \end{pmatrix}$$

mass eigenstates

$$\begin{pmatrix} 1 & 0 \\ 0 & 0 \end{pmatrix} v^{-1} \cdot \begin{pmatrix} \vec{\nu}_L \\ \vec{\nu}_R^c \end{pmatrix} = U_{11}^+ \vec{\nu}_L + \epsilon U_{21}^+ \vec{\nu}_R^c$$

$$\vec{\nu}_L = U_{11}^+ \vec{\nu}_L + \underbrace{\epsilon U_{21}^+ \vec{\nu}_R^c}_{\text{small admixture from heavy right-handed neutrino mass eigenstates}}$$

$\vec{\nu}_R^c$: heavy neutrino eigenstates

$\vec{\nu}_L$: light neutrino eigenstates

\Rightarrow charged current:

$$J^\mu = \vec{\nu}_L \gamma^\mu (1 - \gamma^5) V_{\text{PMNS}} \vec{\nu}_L \quad \left. \vphantom{\vec{\nu}_L} \right\} \text{primary mixing}$$

$$+ \vec{\nu}_R^c \gamma^\mu (1 - \gamma^5) \epsilon \cdot \vec{\nu} \vec{\nu}_L \quad \left. \vphantom{\vec{\nu}_R^c} \right\} \text{suppressed}$$

\Rightarrow good approximation to truncate mixing to (3×3) Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix

complete mixing depends on

$$3 \left(\frac{3 + 2n - 1}{2} \right)$$

$$3 \left(\frac{3 + 2n - 1}{2} \right)$$

mixing angles

phases

(Details: Schechter, Valle Phys Rev D 22, 2227 (1980))

V_{PMNS} is analogue of V_{CKM} in Majorana models.

For Majorana masses phase shifts of fermions do not leave the mass matrix intact — unlike for Dirac masses!

→ cannot eliminate the $(6-1)$ phases as in V_{CKM}
 but only 3 phases from rotation of \vec{e}_L

→ $V_{PMNS} =$

$= V_{23} \cdot V_{31} \cdot V_{12} \cdot K$

$V_{23} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_2 & s_2 \\ 0 & -s_2 & c_2 \end{pmatrix}, V_{31} = \begin{pmatrix} c_3 & 0 & s_3 e^{-i\delta} \\ 0 & 1 & 0 \\ s_3 e^{i\delta} & 0 & c_3 \end{pmatrix}, V_{12} = \dots$

$K = \text{diag}(1, e^{i\phi_1}, e^{i\phi_2})$

← not present in CKM & not present if masses are of Dirac type!

Exp.: $\left. \begin{aligned} \sin^2(2\theta)_{23} &> 0.87 \\ 0.7 &\leq \sin^2(2\theta)_{12} \leq 0.94 \\ \sin^2 \theta_{13} &< 0.051 \end{aligned} \right\}$

2 large mixings
 1 small mixing

$\left. \begin{aligned} |\Delta m_{13}^2| &\sim 10^{-3} \text{ eV}^2 \\ \Delta m_{12}^2 &\sim 10^{-5} \text{ eV}^2 \end{aligned} \right\}$

2 possibilities
 a) usual hierarchy $m_1 < m_2 < m_3$
 b) inverted hierarchy $m_3 < m_1 < m_2$

only δ relevant for neutrino oscillations

ϕ_1, ϕ_2 relevant for neutrinoless double-beta decay

* Models w/ Majorana masses also lead to $\vec{\nu}_L / \vec{\nu}_R^c$ mixing \leadsto weak & current interactions

* Majorana masses explicitly break lepton number

\rightarrow cosmological implications as "Leptogenesis"
& Baryogenesis via (B-L) breaking through sphalerons