

## III Lepton - Sector

### 1. Neutrinos

#### 1.1 Postulation and discovery

##### a) Nuclear $\beta$ -decay

After the observation of discrete energy lines in  $\alpha$  and  $\gamma$  decays the continuous spectrum for  $\beta$ -decays was surprising. To explain the spectrum Pauli proposed a new undetectable light neutral spin- $\frac{1}{2}$  particle (he called it neutron) - later called neutrino by E. Fermi which carries away spin, momentum and energy. (1930)

In the 1930's Fermi developed the first Theory of Neutrons and the  $\beta$ -decay:



After the postulation of the neutrino in 1930 by Pauli (Hypothesis was presented only in 1933 during 17<sup>th</sup> Solway Conf. in Brussels) it took until 1956 before the neutrino was discovered by Cowan & Reines (Nobel Prize 1995):

They used (anti)-neutrons from a nuclear reactor and detected the neutrino via the inverse neutron decay:



##### Experiment

Water tank w/  $\text{CdCl}_2$  and 2 liquid scintillators to detect the  $2\gamma$  from the  $e^+e^-$  annihilation

as well as the  $p$  from the neutron capture reaction:



Finally they observed neutron induced reactions with an average cross section of:

$$\bar{\sigma} = (11.0 \pm 2.6) \cdot 10^{-44} \text{ cm}^2$$

→ Figure of Experiment

### b.) Neutrino and Anti-Neutrino

If neutrinos  $\nu_e$  and anti-neutrinos  $\bar{\nu}_e$  were the same particles, the reaction



should be possible with the Reactor neutrinos ( $\bar{\nu}_e$ , anti-neutrinos).  
(This experiment goes back to an idea of Bruno Pontecorvo  
! 1946 - at that time  $\mu$  and  $\bar{\nu}$  were thought to be the same particle.)

R. Davis (Nobelprize 2002 for the detection of  $\nu$  with the same method) excluded the reaction for reactor neutrinos with a cross section limit:

$$\sigma < 0.9 \cdot 10^{-45} \text{ cm}^2$$

(theoretical estimation  $\approx 2.6 \cdot 10^{-45}$ )

This detection technique was later used to measure the neutrino flux from the sun!

### c.) Helicity of the neutrino (M. Goldhaber, 1957)

With his ingenious experiment M. Goldhaber was able to measure the neutrino helicity:



Only the  $\gamma$  in the direction of the  $\text{Sm}^*$  recoil will undergo resonant scattering:

- Momentum of these photons  $\parallel \nu$

- Spins of  $\nu$  and  $\gamma$  must be opposite  $\rightarrow$  Helicity the same  
 $\chi(\nu) = \chi(\gamma)$

$\rightarrow$  Measurement of the photon polarization via Compton scattering in magnetized iron.

$\rightarrow \chi(\nu) = -\frac{1}{2}$ , i.e. neutrinos are LH particles

Together with the observation of the helicity of the electron in  $\beta$ -decays this observation led to the (V-A)-theory of charged-current interactions.

d) Difference between  $\nu_e$  and  $\nu_\mu$ :

80

(Lederman, Schwartz, Steinberger et al., 1962)  
 = Nobel prize, 1988 =

Use  $\nu$ 's from  $\pi$  decays at Brookhaven AGS:

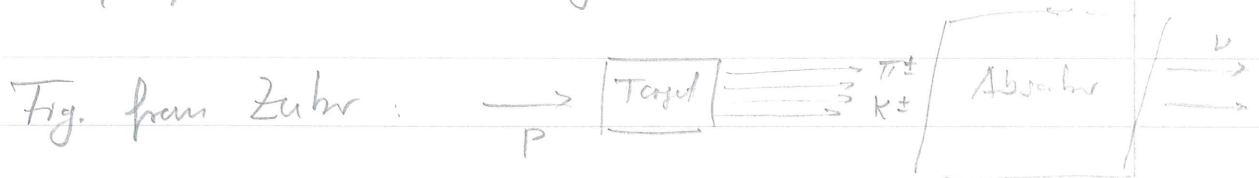


→ these neutrinos in the interaction with matter



produce only  $\mu$ 's and are therefore different from the neutrinos involved in  $\beta$ -decays/nuclear reactors.

→ the observed production rate of muon-like events is what one expects from  $\pi$ -like. There are no  $\nu_e \neq \nu_\mu$ . No indication for e-like events.



Discovery of the  $\tau$ -Neutrino (Donut Exp, 2000)

Direct Observation of  $\nu_\tau$

800 GeV p beam from Tevatron on Beam-Dump,

( $8 \times 10^{12}$  p/20s  $\approx$  20kW beam power)

97%  $\nu$  are  $\nu_e, \nu_\mu$  from  $\pi, K$  decays

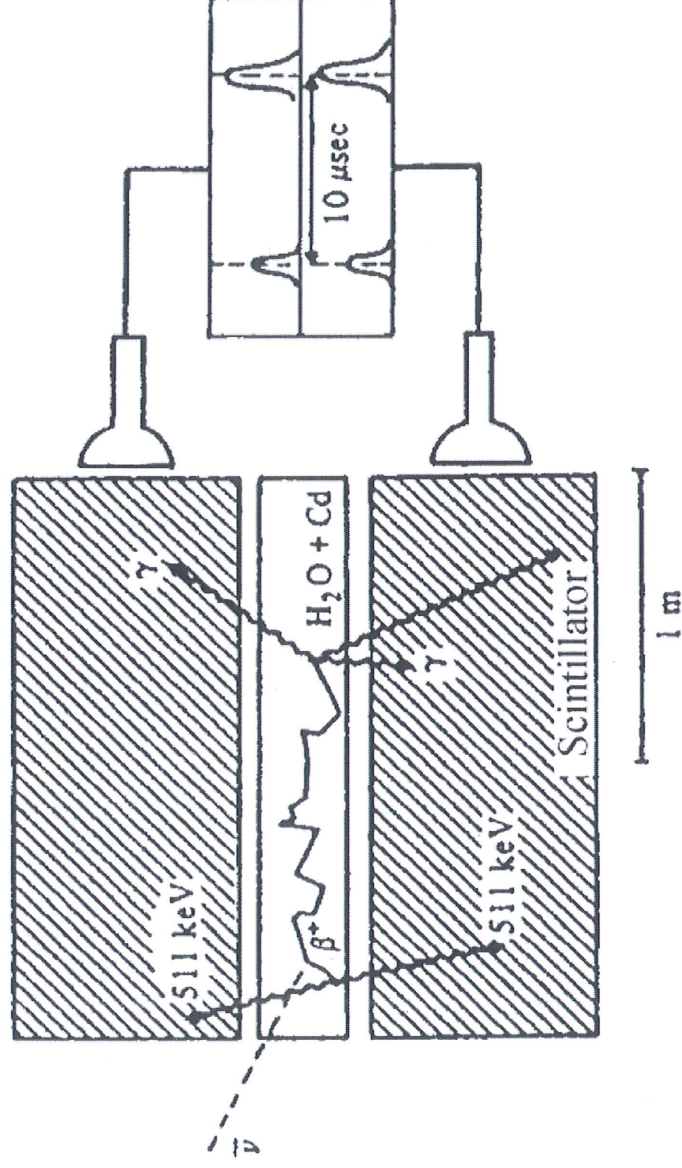
3%  $\nu_\tau$  mostly from  $D_s \rightarrow \nu_\tau \tau$



→ 9  $\tau$  events observed.

# Neutrino Discovery

Cowan & Reines, 1956 (Nobel prize 1995)



# Project Poltergeist & Herr Auge

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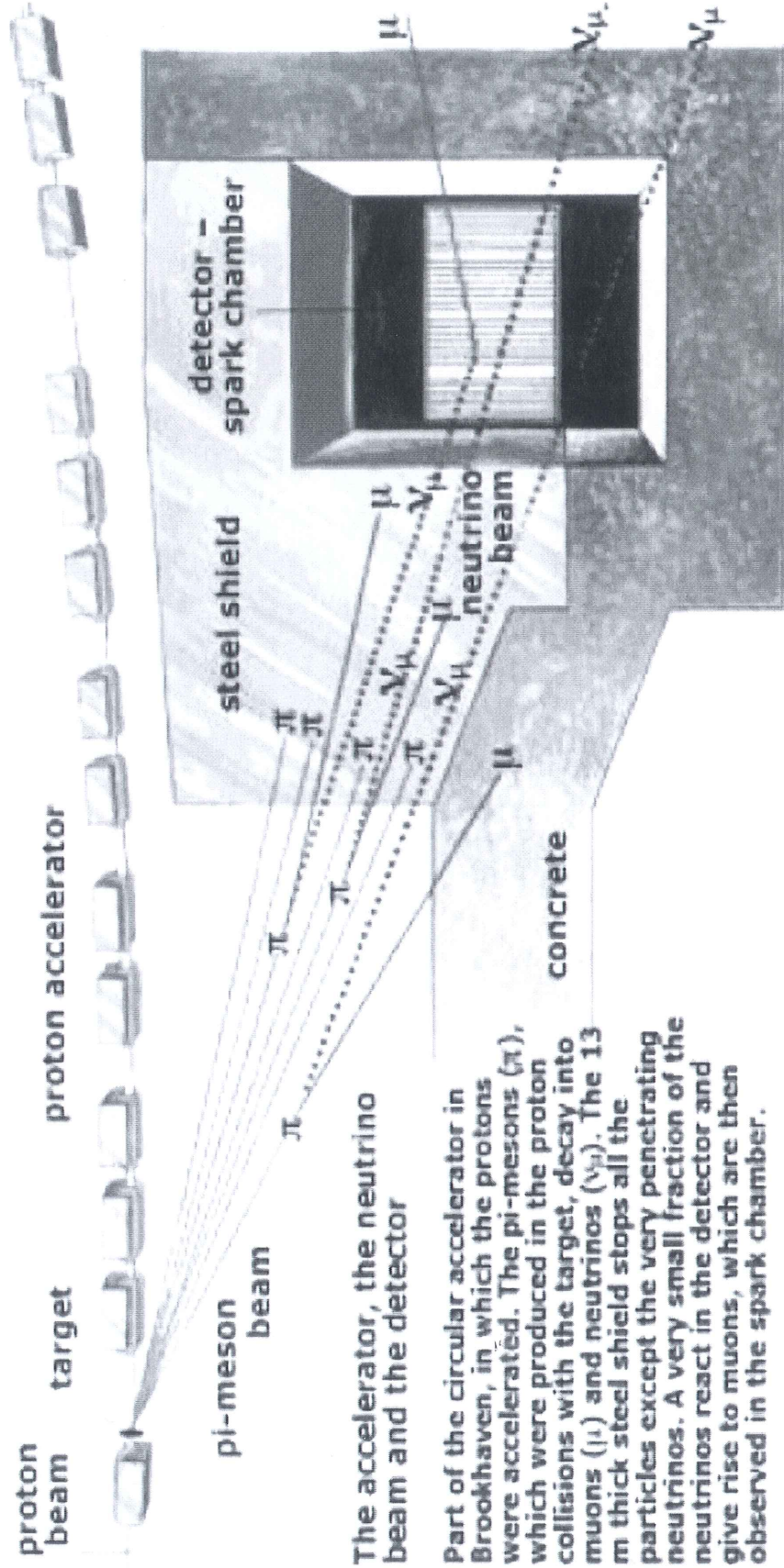


C. Cowan

F. Reines



# Electron and Muon Neutrino

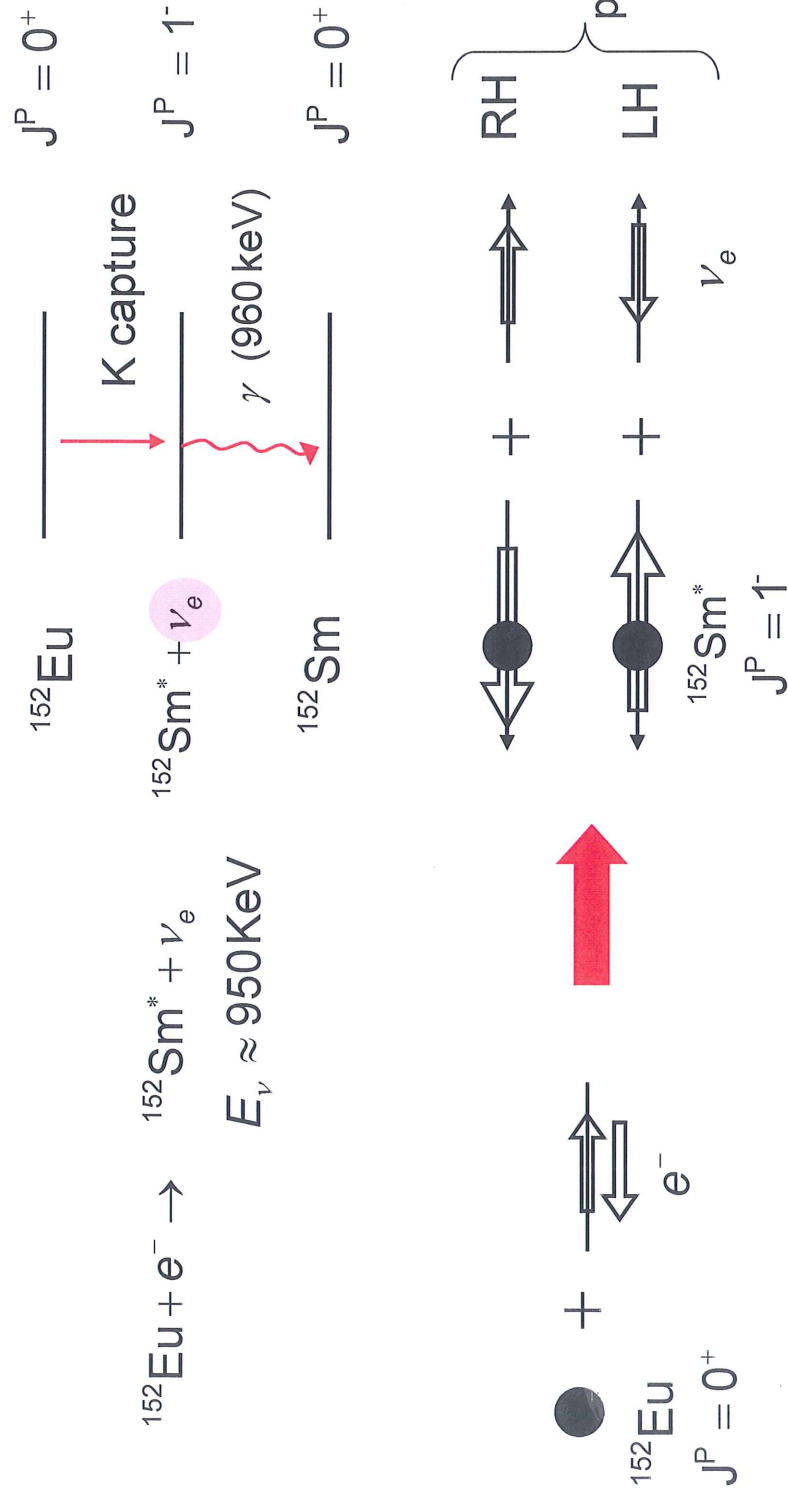


The accelerator, the neutrino beam and the detector

Part of the circular accelerator in Brookhaven, in which the protons which were accelerated. The pi-mesons ( $\pi$ ), which were produced in the proton collisions with the target, decay into muons ( $\mu$ ) and neutrinos ( $\nu_\mu$ ). The 13 m thick steel shield stops all the particles except the very penetrating neutrinos. A very small fraction of the neutrinos react in the detector and give rise to muons, which are then observed in the spark chamber.

# Neutrino Helicity

M. Goldhaber (1957)

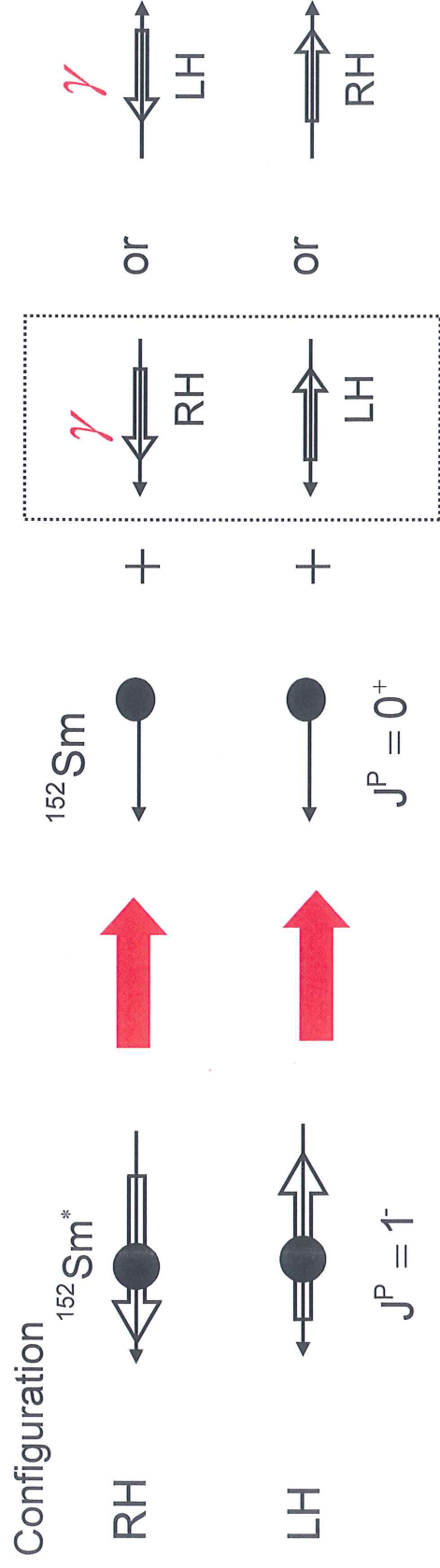


Sm undergoes small **recoil** ( $p_{\text{recoil}} = 950 \text{ KeV}$ ). Because of angular momentum conservation Spin  $J=1$  of  $\text{Sm}^*$  is opposite to neutrino spin. Important: **neutrino helicity is transferred to the Sm nucleus.**



# Neutrino Helicity

2.  $\gamma$  emission:  $^{152}\text{Sm}^*(J^P = 1^-) \rightarrow ^{152}\text{Sm}(J^P = 0^+) + \gamma$



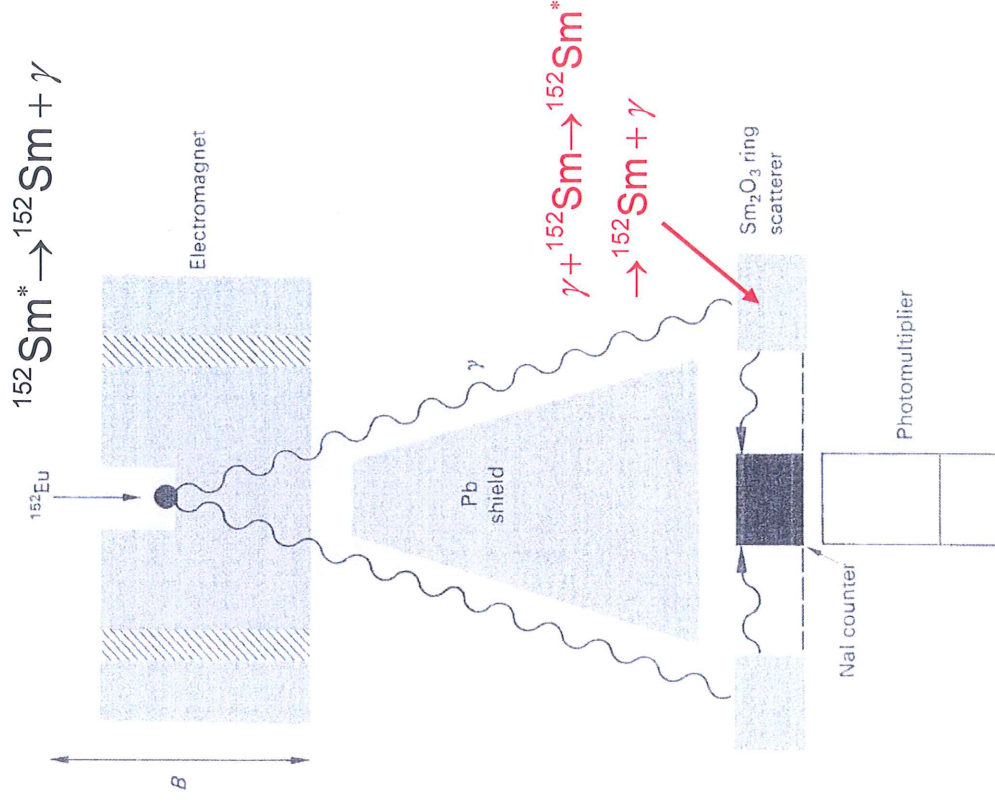
Photons along the Sm recoil direction carry the polarization of the  $\text{Sm}^*$  nucleus

- How to select photons along the recoil direction?  $\Rightarrow$  3
- How to determine the polarization of these photons?  $\Rightarrow$  4

# Neutrino Helicity

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3. Resonant photon scattering:  $\gamma + {}^{152}\text{Sm} \rightarrow {}^{152}\text{Sm}^* \rightarrow {}^{152}\text{Sm} + \gamma$



4.

Determination of the photon polarization

Exploit that the transmission index through magnetized iron is polarization dependent:  
Compton scattering in magnetized iron

$$P_\gamma = -0.66 \pm 0.14 \quad (\text{expect. } 0.75)$$