

# Statistical Hadron Production from AGS to Collider Energies

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- Statistical Model
- Fixed Target Data
- RHIC Data
- Chemical Freeze-Out and the Phase Boundary
- Chemical Freeze-Out vis-a-vis Initial Condition and Thermal Freeze-Out
- Outlook and Open Questions

# Choice of Statistical Ensemble

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- **Grand Canonical Ensemble (GC):** in large system, with large number of produced particles, conservation of additive quantum numbers ( $B, S, I_3$ ) can be implemented on average by use of chemical potential  $\mu$ ; asymptotic realization of exact canonical approach
- **Canonical Ensemble (C):** in small system, with small particle multiplicity, conservation laws must be implemented locally on event-by-event basis (Hagedorn 1971, Shuryak 1972, Rafelski/Danos 1980, Hagedorn/Redlich 1985)  
→ severe phase space reduction for particle production “canonical suppression”
- **C relevant in**
  - low energy HI collisions (Cleymans/Redlich/Oeschler 1998/1999)
  - high energy hh or  $e^+e^-$  collisions (Becattini/Heinz 1996/1997)
  - peripheral HI collisions (Cleymans/Oeschler/Redlich 1998, Hamieh/Redlich/Tounsi 2000)

- **Connection C and GC:** recently shown by Tounsi/Redlich hep-ph/0111159 and 0111261:

$$n_k^C = n_k^{GC} F_s$$

where k stands for all hadrons with a given strangeness S

- for strangeness 1: <sup>1</sup>

$$F_s = \frac{I_1(x)}{I_0(x)}$$

x describes size of thermal phase space available for strange particles

$$x = 2S_1 = 2 \sum_k Z_k^1 \propto V = (V_0/2) N_{part}$$

ignoring multistrange hadrons, x is total of all particles with S=1 in GC

- for multistrange:

$$F_s = \frac{I_s(x)}{I_0(x)}$$

- limiting cases:

$$\lim_{x \rightarrow \infty} \frac{I_1(x)}{I_0(x)} = 1 \quad C \rightarrow GC$$

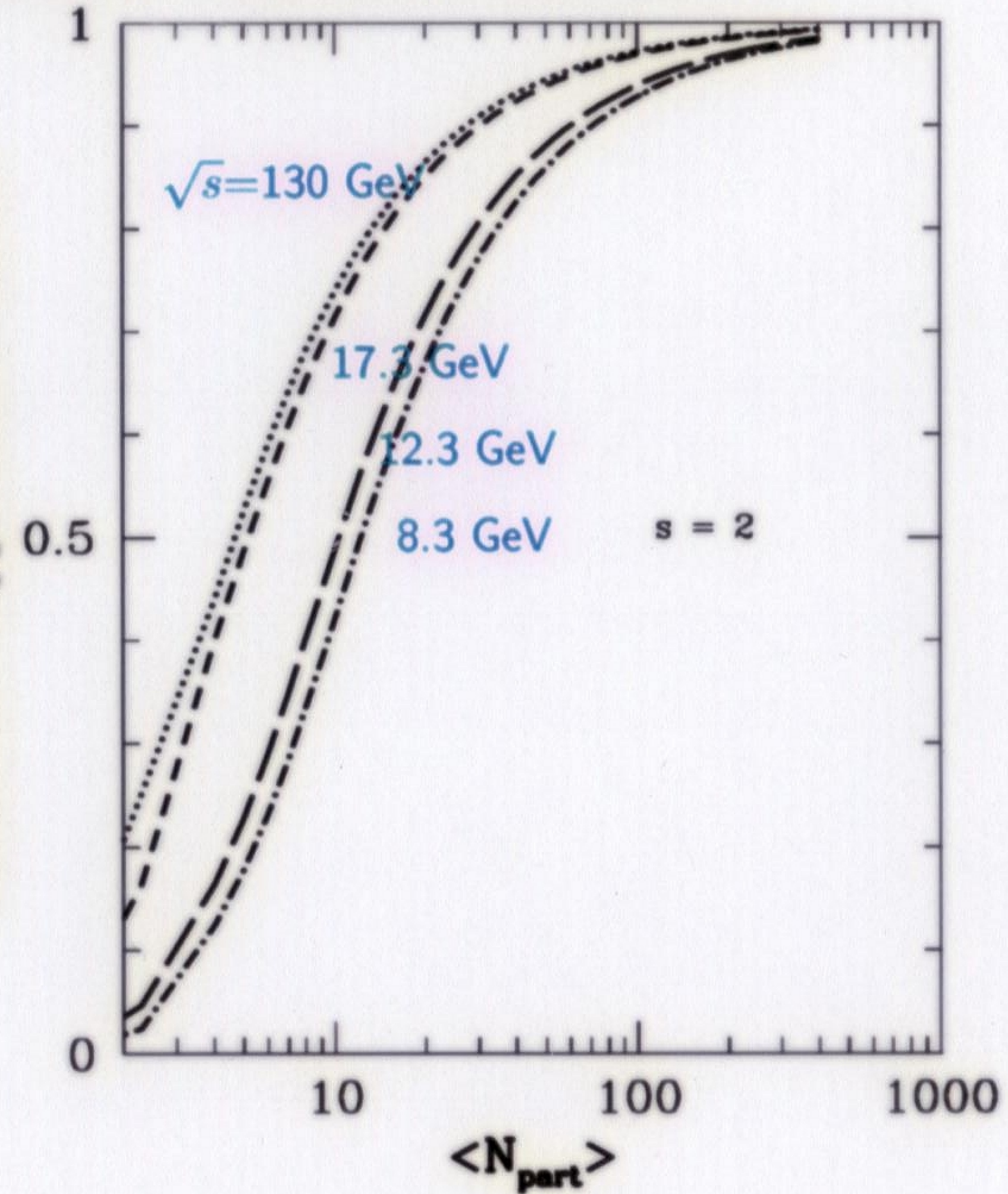
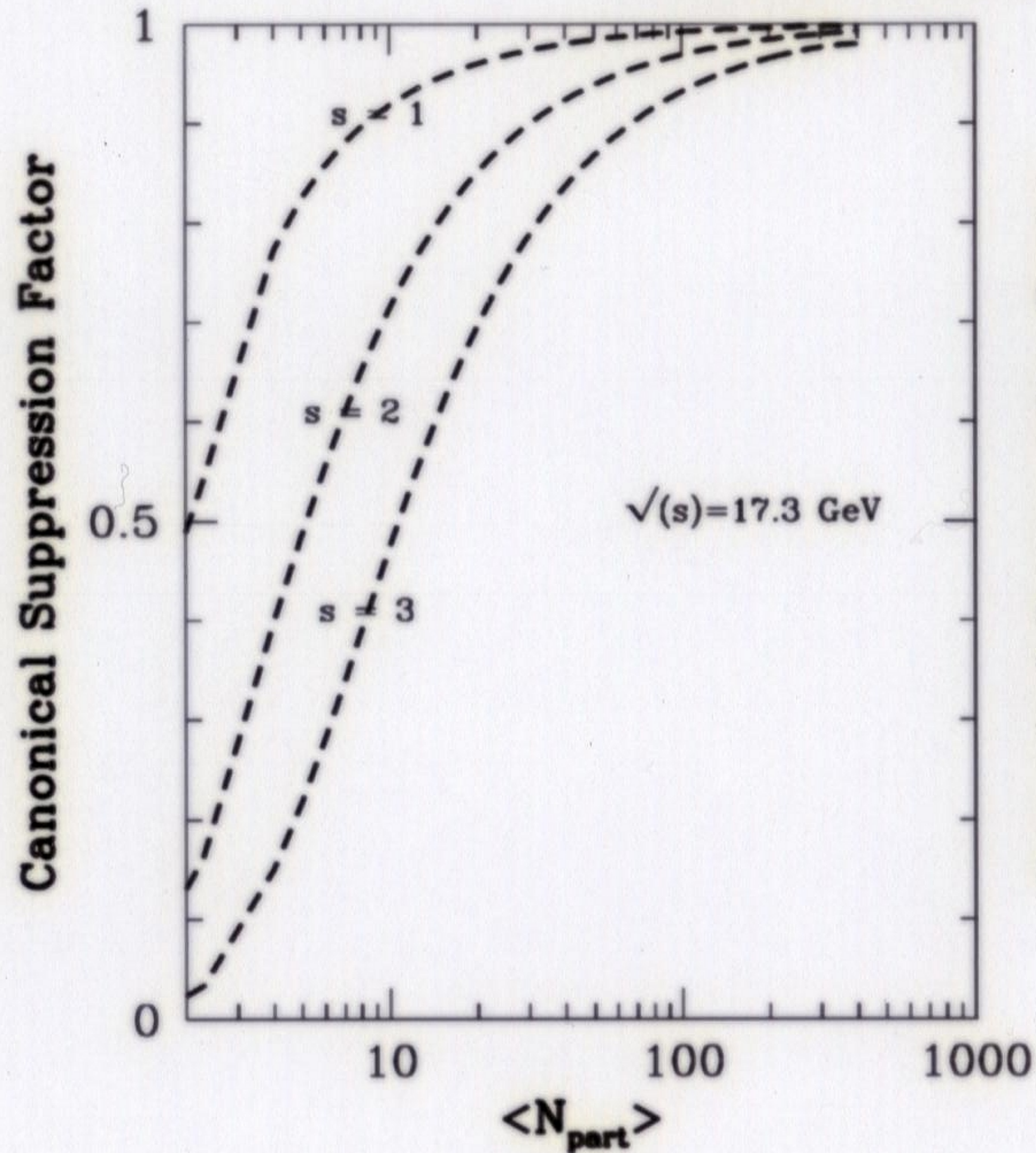
$$\lim_{x \rightarrow 0} \frac{I_n(x)}{I_0(x)} \rightarrow (x/2)^n \quad \propto V^n$$

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<sup>1</sup>first seen for charge conservation in pion production Horn/Silver 1970 and similar Shuryak 1972

# Canonical Suppression Factor

Tounsi and Redlich, hep-ph/0211159



for  $N_{part} \geq 60$  Grand Canonical ok to better 10%

# Grand Canonical Ensemble

$$\ln Z_i = \frac{V g_i}{2\pi^2} \int_0^\infty \pm p^2 dp \ln[1 \pm \exp[-(E_i - \mu_i)/T]]$$

$$n_i = N/V = -\frac{T}{V} \frac{\partial \ln Z_i}{\partial \mu} = \frac{g_i}{2\pi^2} \int_0^\infty \frac{p^2 dp}{\exp[(E_i - \mu_i)/T] \pm 1}$$

$$\mu_i = \mu_B B_i + \mu_S S_i + \mu_{I_3} I_i^3$$

for every conserved quantum number there is a chemical potential  $\mu$  but can use conservation laws to constrain:

- Baryon number:  $V \sum_i n_i B_i = Z + N \rightarrow V$
- Strangeness:  $V \sum_i n_i S_i = 0 \rightarrow \mu_S$
- Charge:  $V \sum_i n_i I_i^3 = \frac{Z - N}{2} \rightarrow \mu_{I_3}$

This leaves only  $\mu_b$  and  $T$  as free parameter when  $4\pi$  considered for rapidity slice fix volume e.g. by  $dN_{ch}/dy$

- **finite volume correction:** a la Balian-Bloch

$$f = 1 - \frac{3\pi}{4pR} + \frac{1}{(pR)^2}$$

- **interactions:** van der Waals type via excluded volume correction

thermodynamically consistent form Rischke/Gorenstein/Stöcker/Greiner 1990

$$p^{excl.}(T, \mu) = p^{id.gas}(T, \hat{\mu}); \quad \text{with } \hat{\mu} = \mu - v_{eigen} p^{excl.}(T, \mu)$$

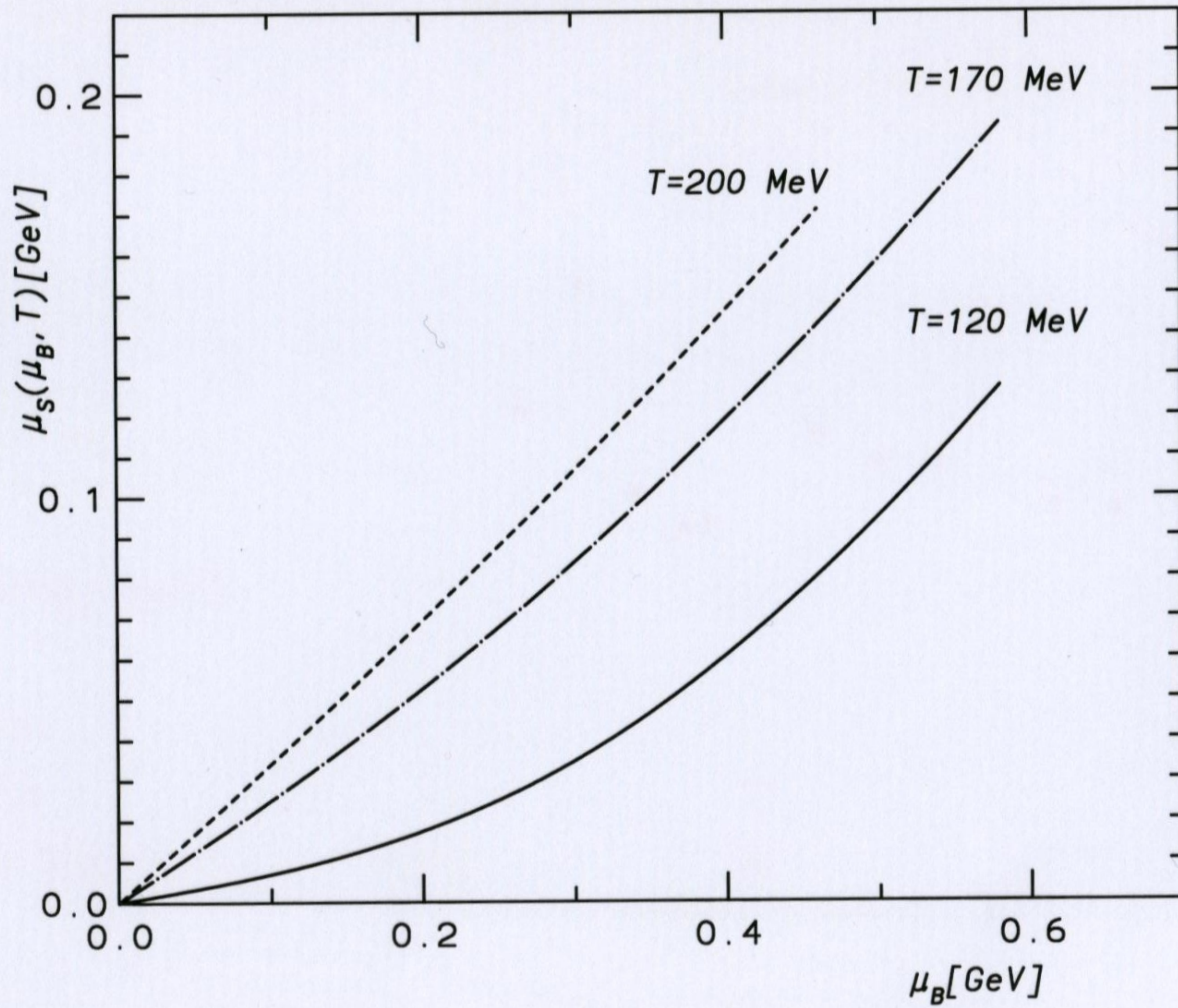
use  $r=0.3$  fm (hard core of nucleon-nucleon interaction)

- **resonance width:** e.g. Weinhold, Friman, Nörenberg, 1996

$$\ln Z_R = N \frac{V d_R}{2\pi^2} T \exp[\mu/T] \int_{s_{min}}^{s_{max}} ds s K_2(\sqrt{s}/T) \frac{1}{\pi} \frac{m_R \Gamma_R}{(s - m_R^2)^2 + m_R^2 \Gamma_R^2}$$

not important for large enough T

# Dependence of $\mu_S$ on $T$ and $\mu_B$



# CERN SPS Data and Thermal Model

P. Braun-Munzinger, I. Heppe, J. Stachel, Phys.Lett.**B465** (1999) 15 + reanalysis in 2003 with more data

central 158 A GeV/c Pb + Pb collisions

free parameters:

$$T = 0.170 \pm 0.005 \text{ GeV}$$

$$\mu_b = 0.255 \pm 0.010 \text{ GeV}$$

fixed by conservation laws:

$$\mu_s = 0.074 \text{ GeV from } \Delta S=0$$

$$\mu_{I_3} = 0.005 \text{ GeV from } \Delta Q=0$$

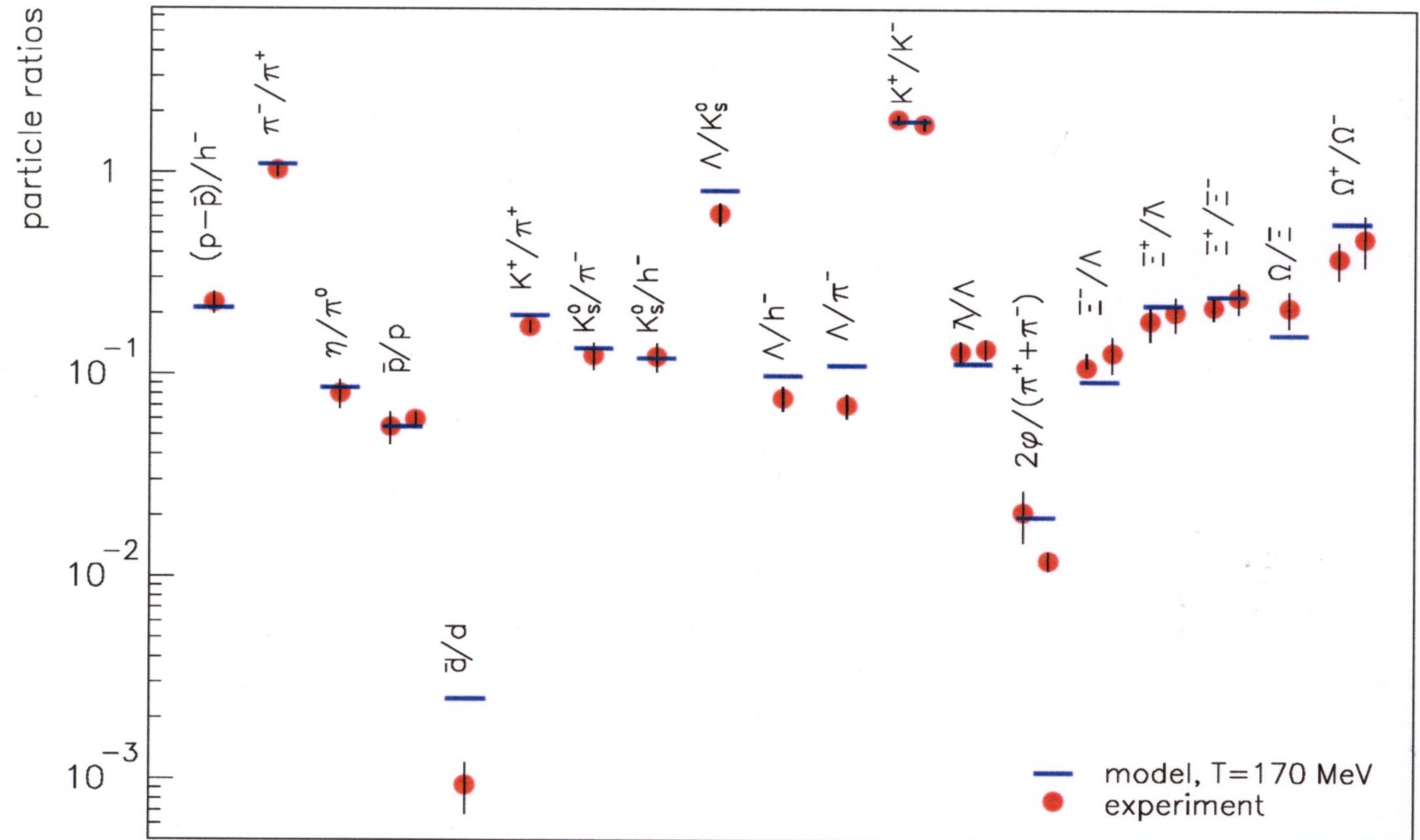
reduced  $\chi^2$  (excluding  $\phi$  and d)

2.0

largest contribution:

weak feeding of  $\pi$

affects  $\Lambda/\pi$  and  $\Lambda/h^-$

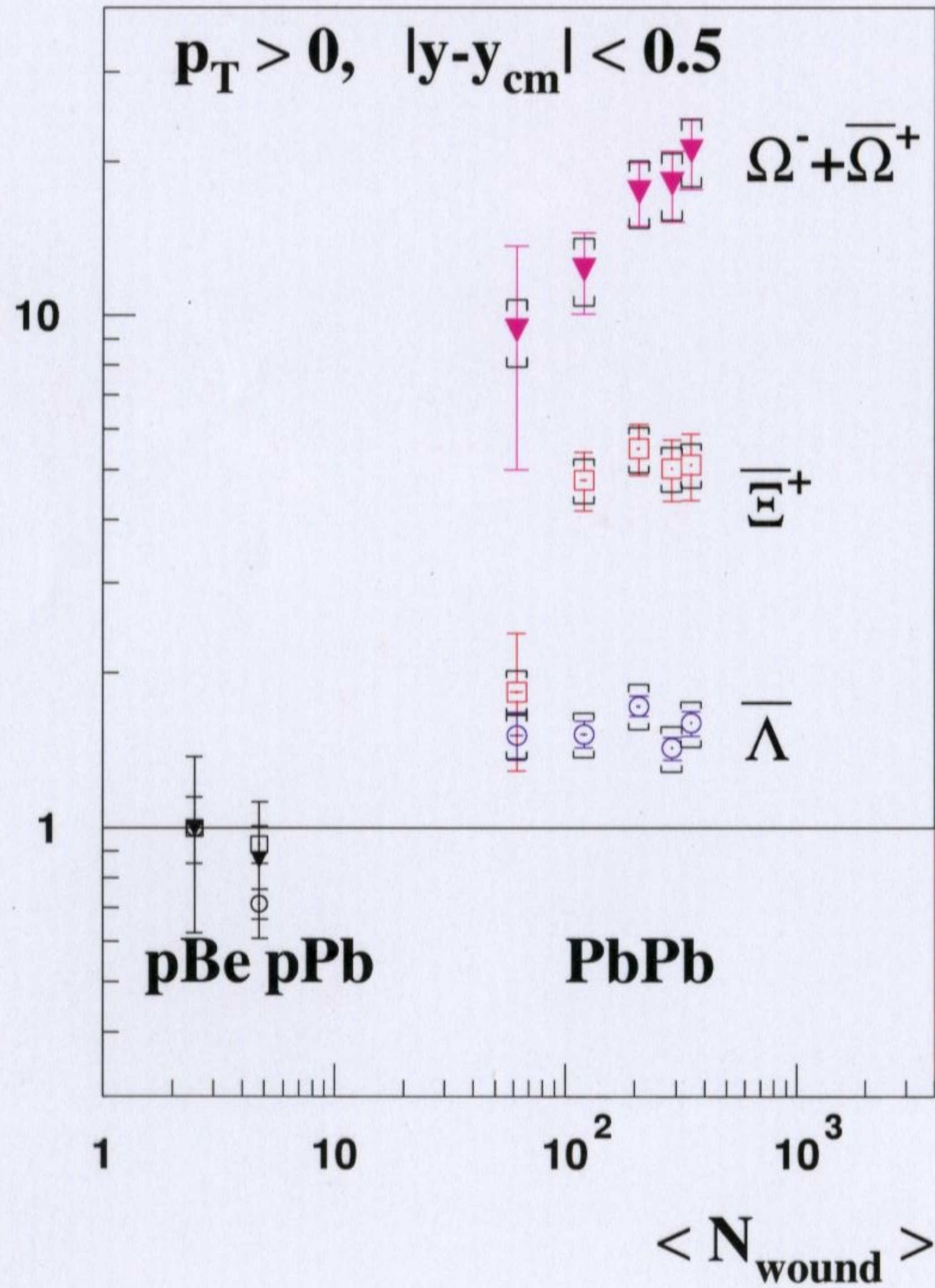




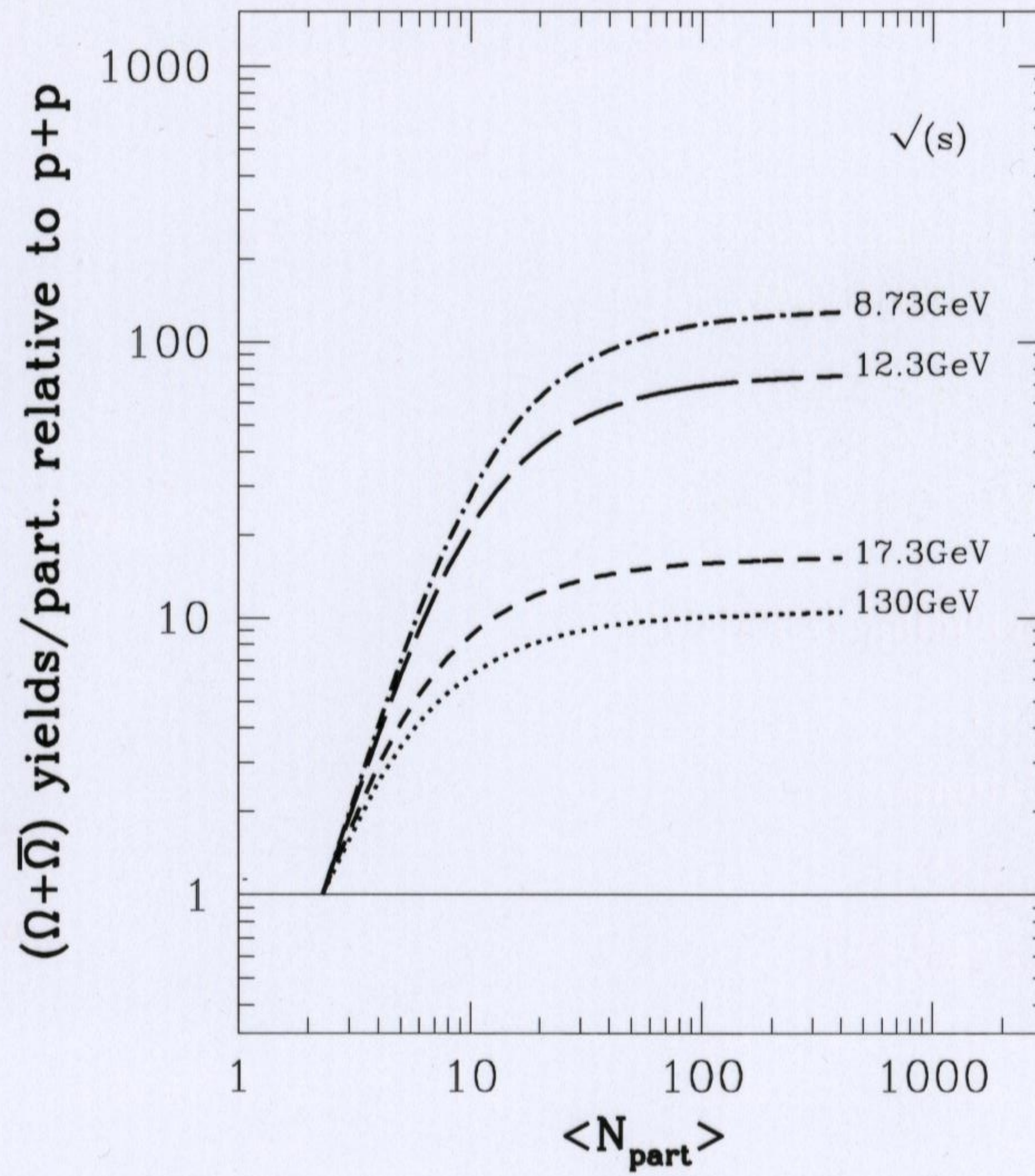
# Strangeness Enhancement in 158 A GeV/c Pb + Pb Collisions

NA57

Particle / event / w. nucl. relative to pBe



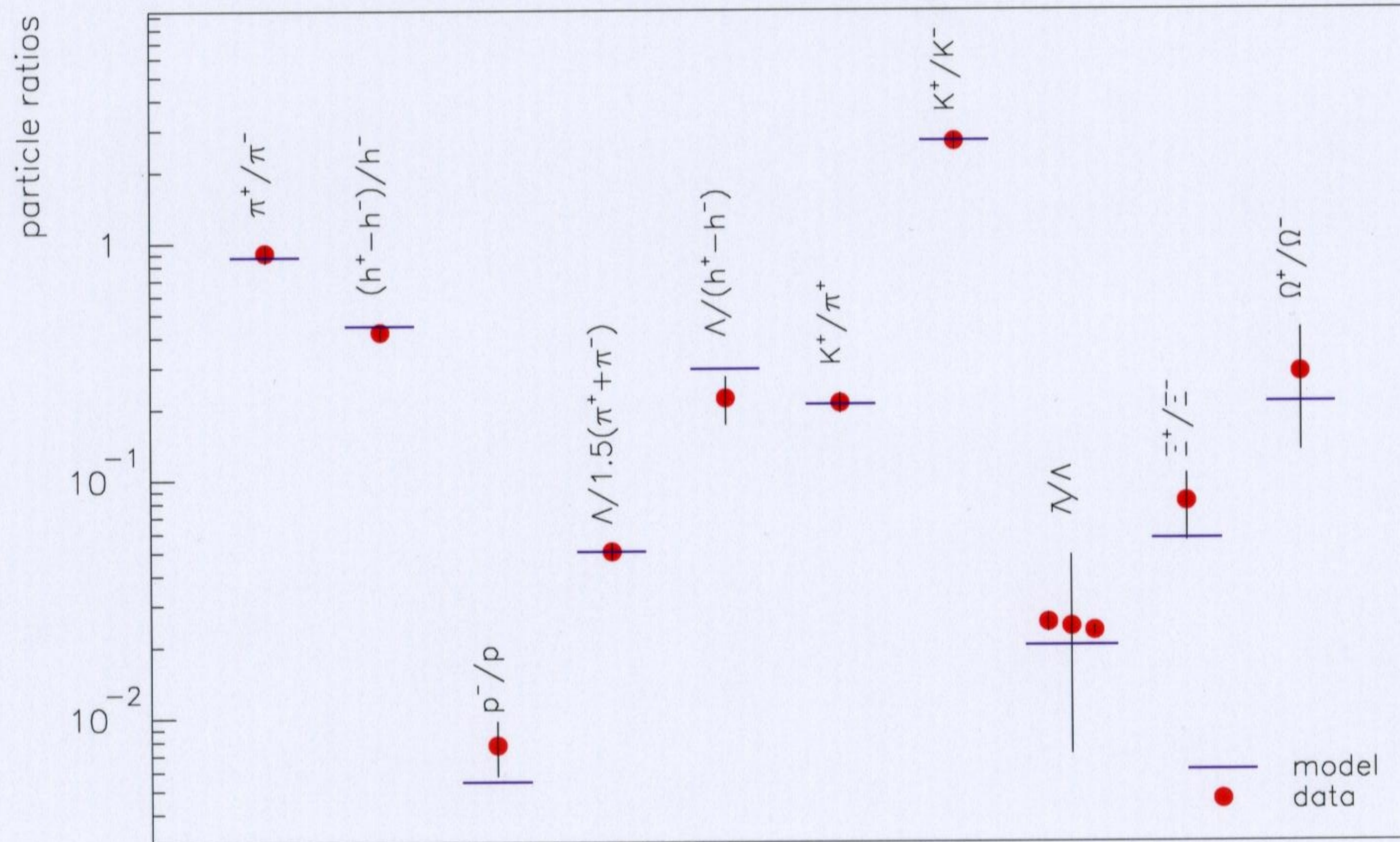
Tounsi/Redlich



# Hadron Yields at SPS and Thermal Model

P. Braun-Munzinger, D. Magestro, J. Stachel, Dec. 02

central 40 A GeV/c Pb + Pb collisions - thermal model parameters:  $T = 148$  MeV,  $\mu_b = 400$  MeV



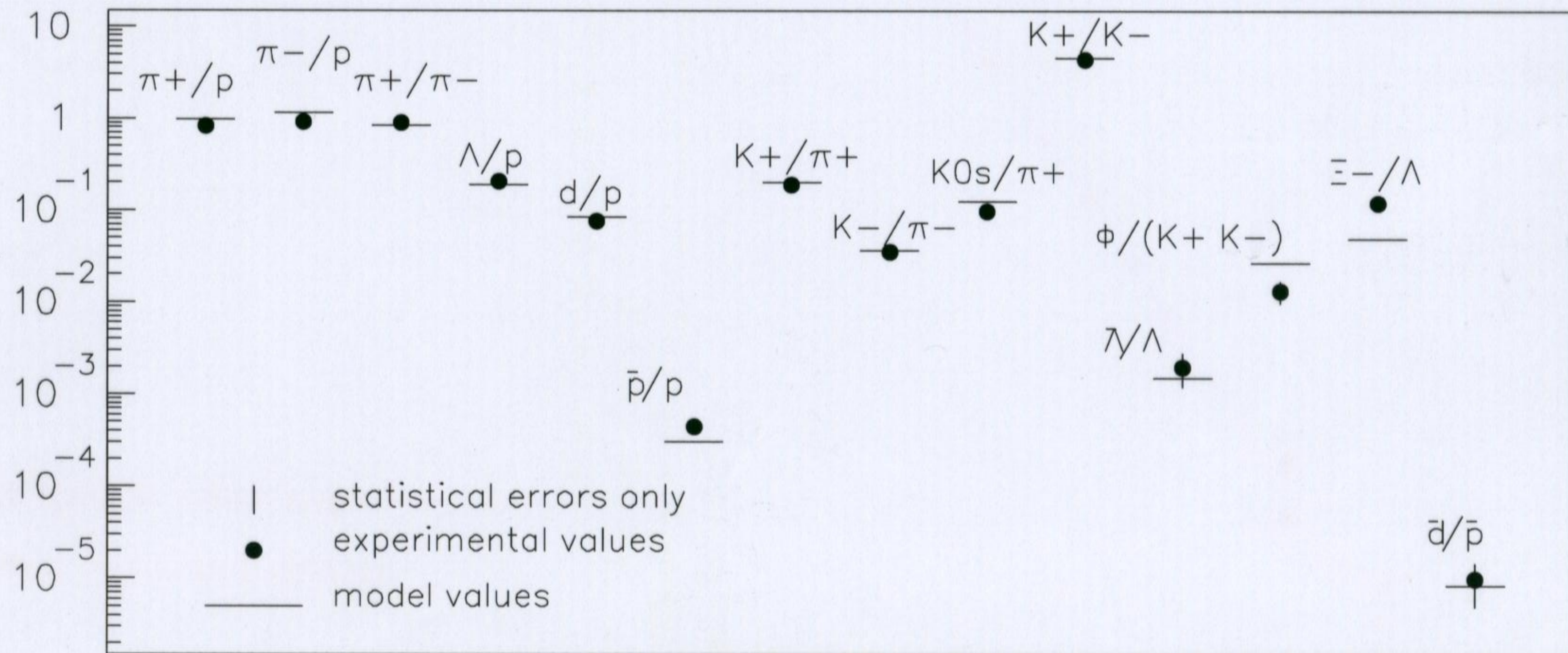
reduced  $\chi^2 = 1.1$

# Hadron Yields at AGS and Thermal Model

P. Braun-Munzinger, I. Heppe, J. Stachel, Phys. Lett. **B465** (1999) 5  
and I. Heppe, Diploma thesis, U. Heidelberg 1998

central 14.6 A GeV/c Si + Au collisions

thermal model parameters:  $T = 125$  MeV,  $\mu_b = 540$  MeV



*D. Prorok: also  $E_T$  and  $N_{ch}$  described by same parameters - consistency checks*  
yields for 11.5 A GeV/c Au + Au are very similar

# Yields of Light Nuclei at AGS and Thermal Model

Addition of every nucleon  $\rightarrow$  penalty factor  $R_p=48$

but data are at very low  $p_t$

$p_t$  int. with A-dependent slope  $\rightarrow R_p=26$

Grand Canonical Ensemble:

$$R_p \approx \exp[(m_n \pm \mu_b)/T]$$

for  $T=125$  MeV and  $\mu_b = 540$  MeV

$\rightarrow R_p=23$  good agreement!

also good for antideuterons

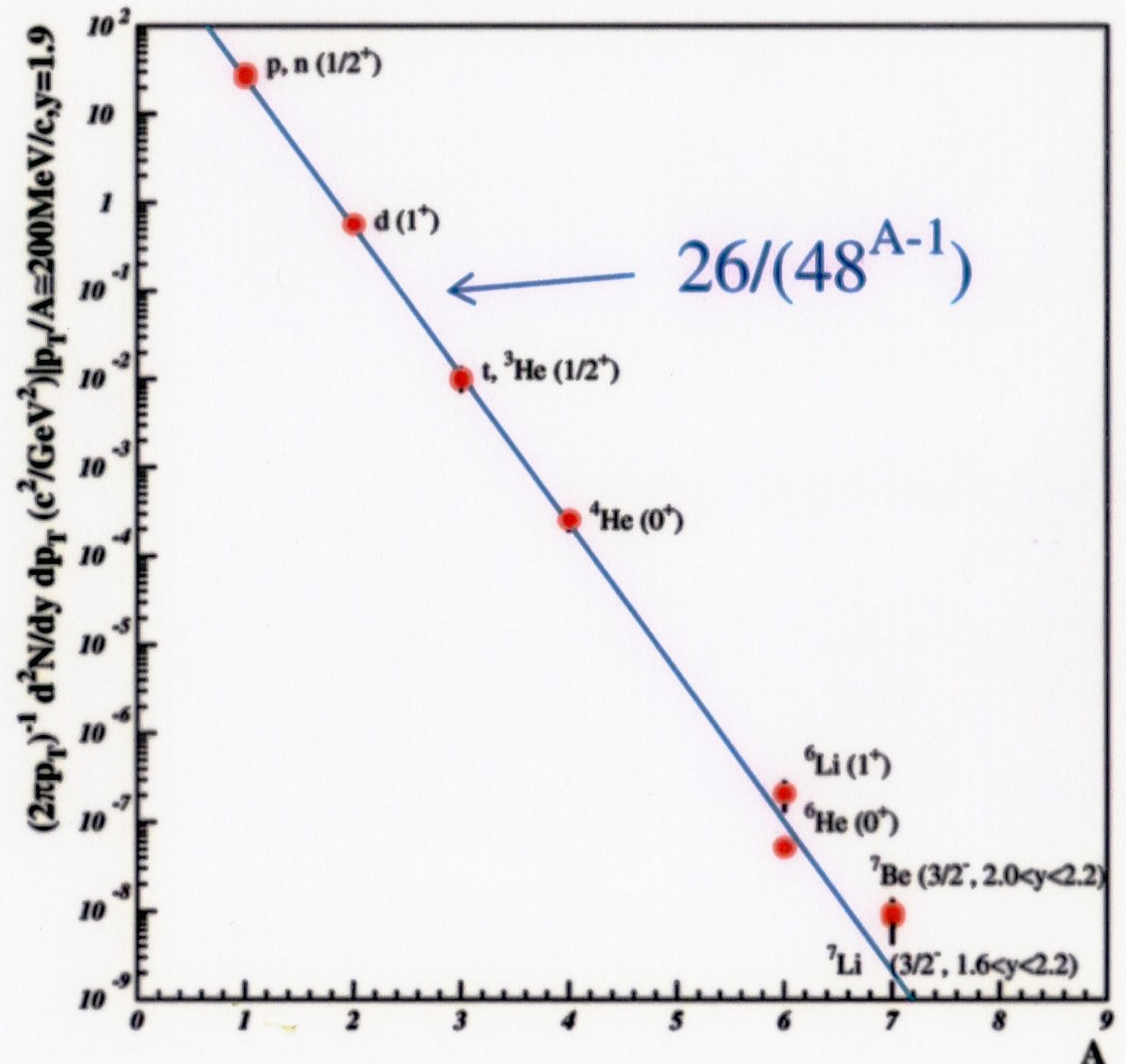
data:  $R_p=2 \pm 1 \cdot 10^5$  GC:  $R_p=1.3 \cdot 10^5$

P.Braun-Munzinger, J.Stachel

J.Phys.G28(2002)1971

Note: AGS may be special here since  
chemical and thermal freeze-out coincide

E864 Collaboration, Phys. Rev. C61 (2000) 064908



## The Perennial Question of Use of $\gamma_S$

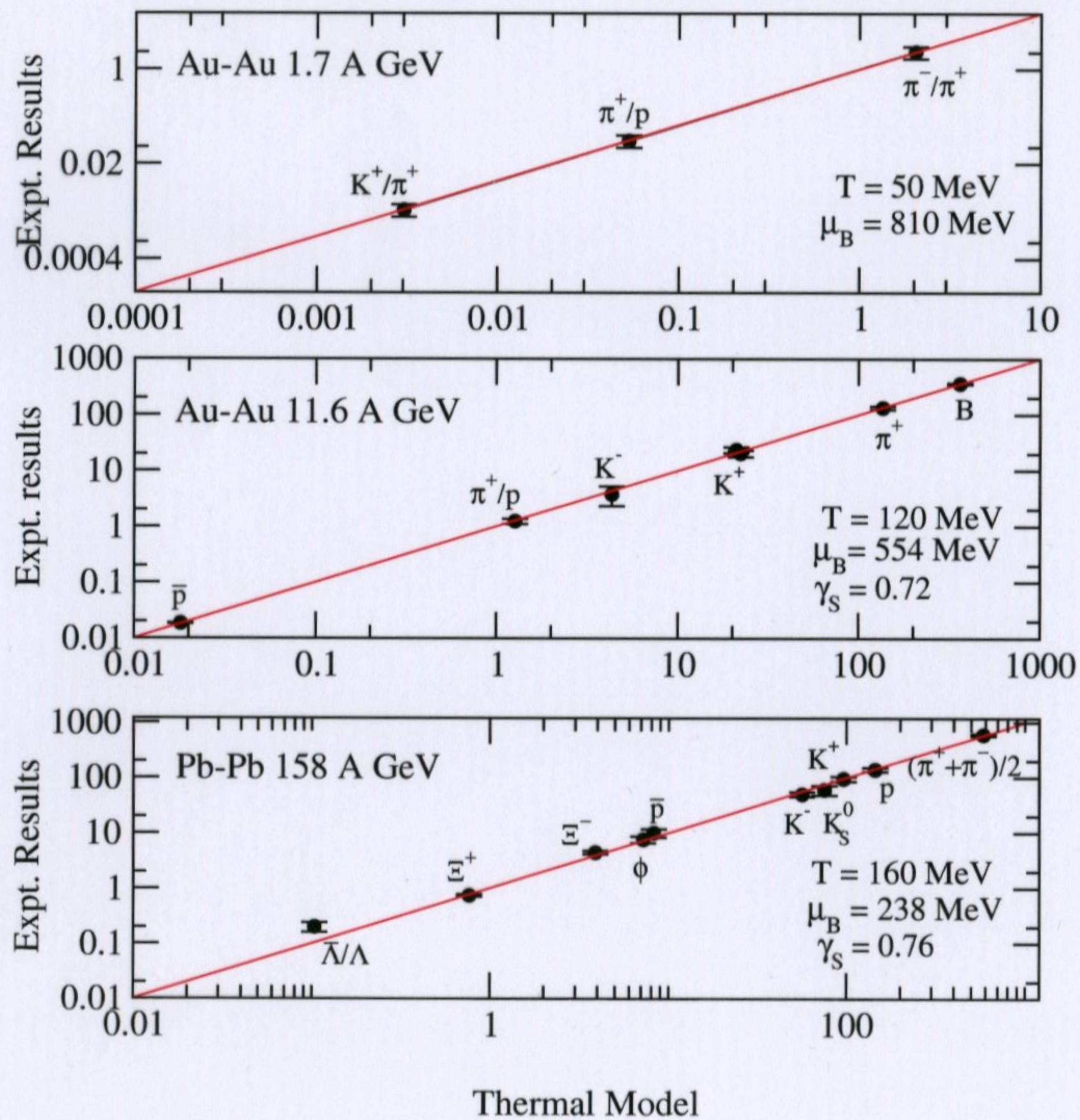
$$n_i(S) = \gamma_S \frac{g_i}{2\pi^2} \int_0^\infty \frac{p^2 dp}{\exp[(E_i - \mu_i)/T] \pm 1}$$

- Looks similar to canonical suppression but in that case dependence on  $V_0$  and  $N_{part}$  fixed.
- $\gamma_S$  as free fit parameter for every system and centrality has no thermodynamic justification, rather a measure of lack of equilibration
- but if really no thermalization, need dynamical treatment, doubtful that this could be replaced by factor  $\gamma_S$

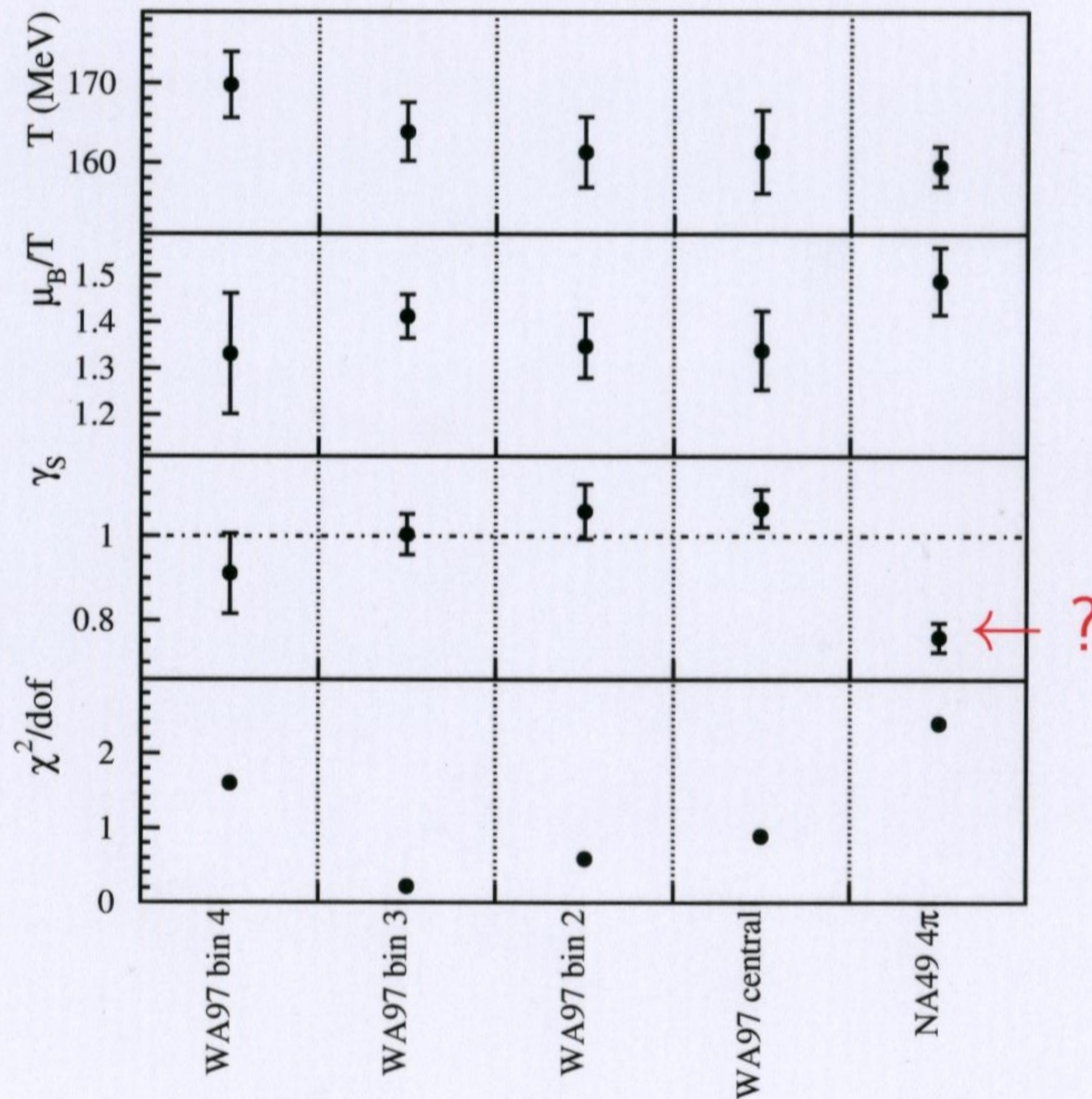
# Strangeness Undersaturation?

central Au + Au and Pb + Pb collisions, from SIS to SPS

F. Becattini, J. Cleymans, A. Keranen, E. Suhonen, K. Redlich, hep-ph/0011322



F. Becattini, J.Phys. G28 (2002) 1553

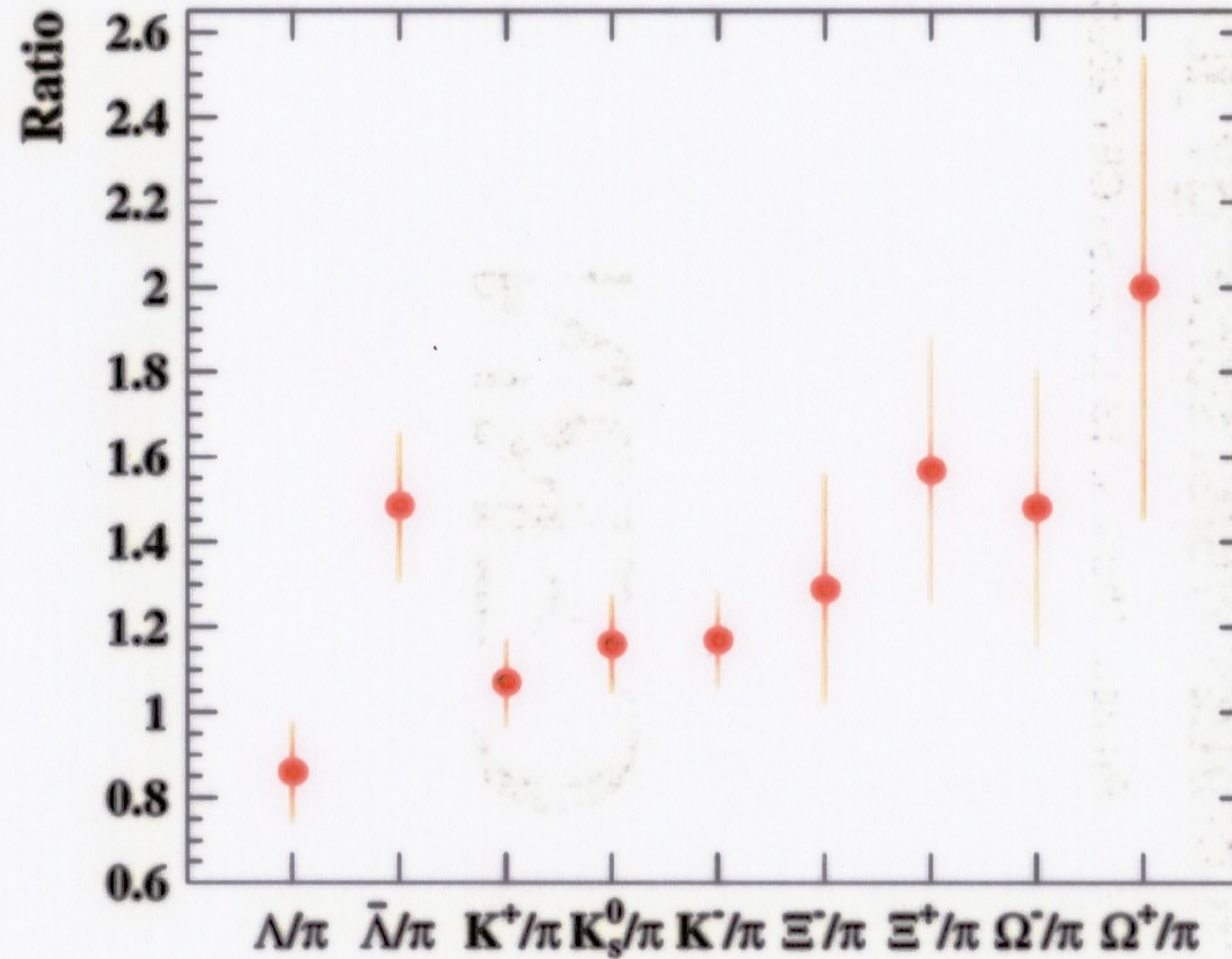


# Difference Mid-rapidity vs. $4\pi$ Data

all data NA49, all measured or scaled to top 10 % centrality

$\gamma=0 / 4\pi$

158 A GeV/c Pb + Pb collisions



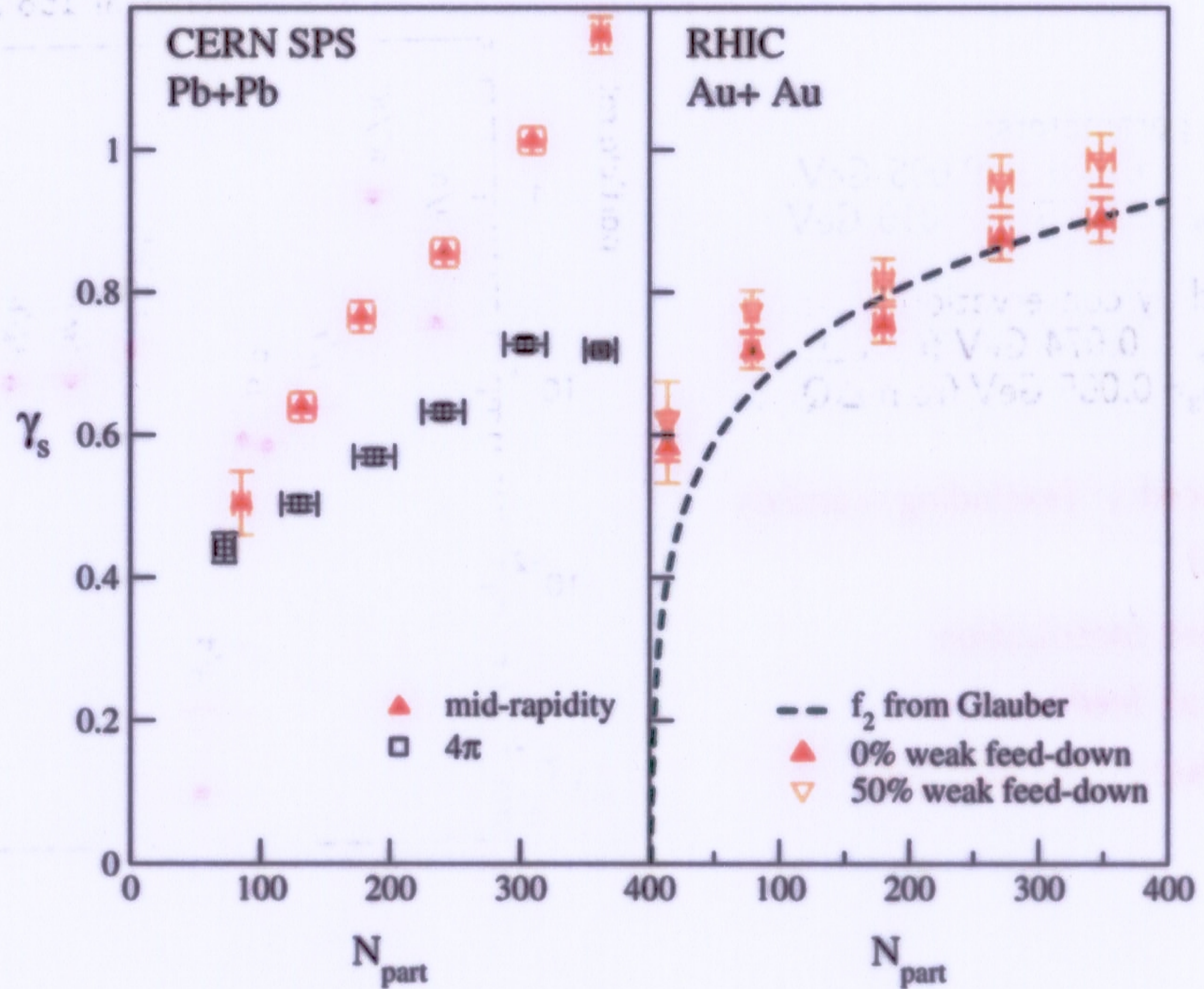
Fit of mid-rapidity data does not need strangeness suppression

# Centrality Dependence of Strangeness Saturation

Cleymans, Kämpfer, Steinberg, Wheaton, hep-ph/0212335

Fit  $\mu_B$  and  $\gamma_S$  to  $\pi$ , K, p yields

$f_2$  fraction of  $N_{part}$   
with multiple collisions



Central collisions reach strangeness saturation at mid-rapidity  
constant  $\phi/\pi$  from STAR does not support  $\gamma_S \leq 1$  for more peripheral  
*centrality dependence of R (NA57) not consistent w. above  $\gamma_S$*



## Lack of strangeness saturation for Light Systems and Peripheral Coll.

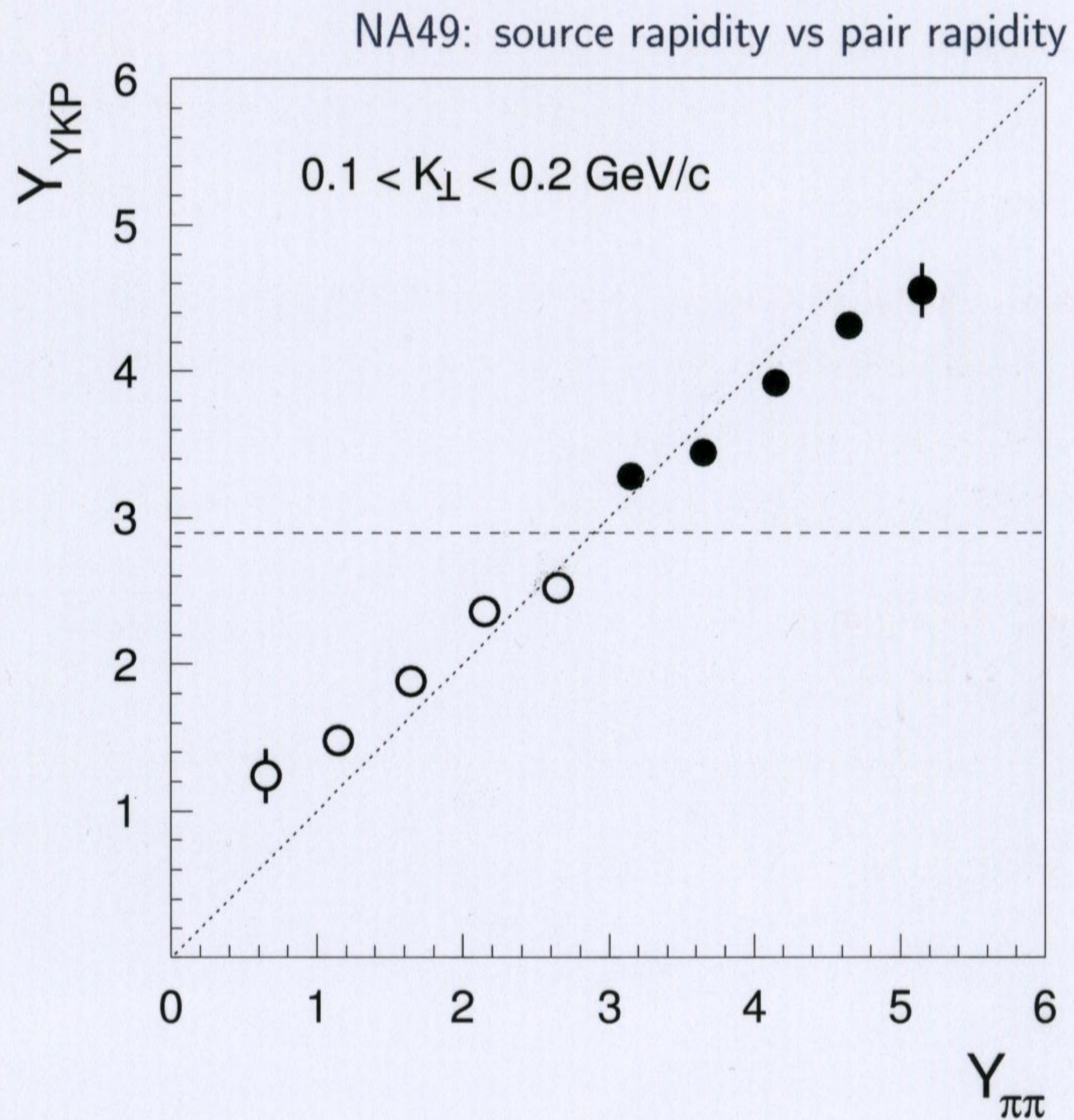
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centrality dependence of PbPb data and  
of CC and SiSi C. Höhne, NA49, QM02

- no simple  $N_{part}$  scaling as for  
e.g. canonical suppression
- scales rather with number of collisions  
or fraction of nuclei having undergone  
multiple collisions

# Choice of Rapidity Window

- **Bjorken expansion:** boost invariant  
→ any rapidity window ok
- **AGS and lower:** no separation of central and fragmentation regions  
→ choose  $4\pi$
- **RHIC:** rapidity plateau  $\Delta y \approx \pm 2$   
→ within this window any cut should be okay
- **Full energy SPS:** evidence that some transparency sets in, distinction central region - fragmentation region becomes meaningful  
→ assess width of plateau from HBT  
and get  $\Delta y \approx \pm 1$



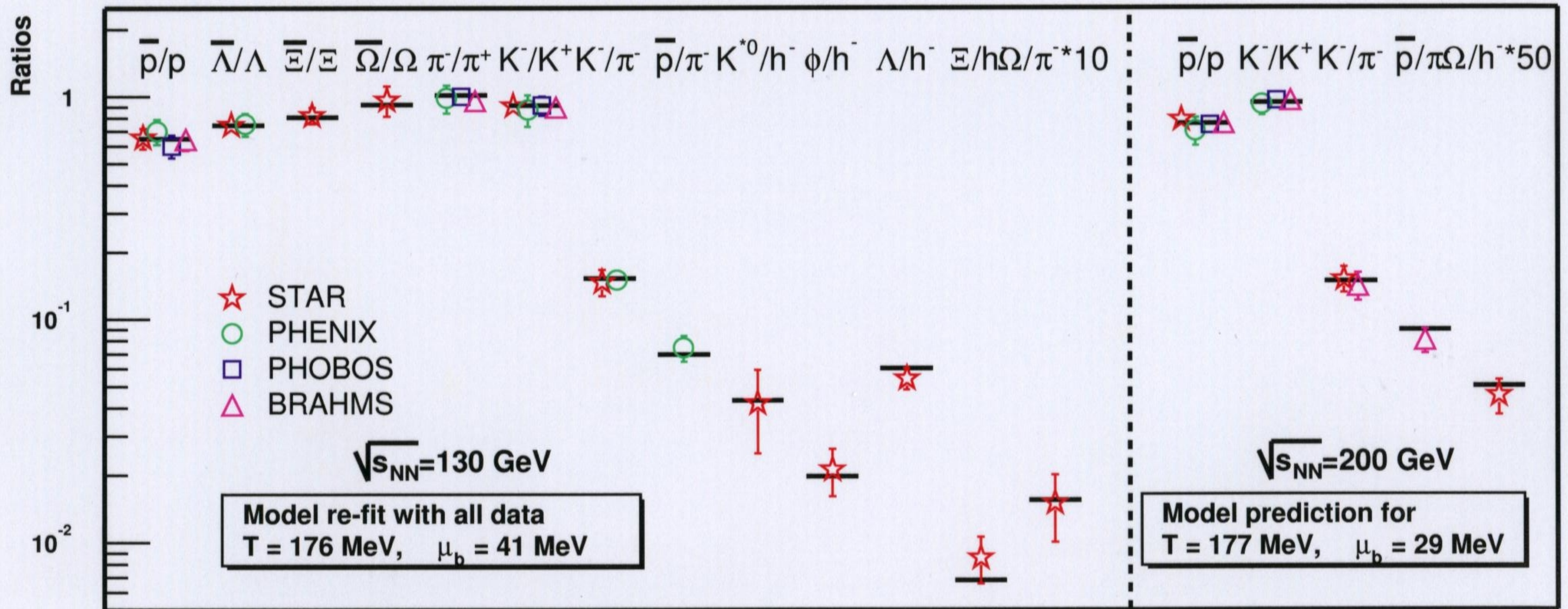
# RHIC Data and Thermal Model

P. Braun-Munzinger, D. Magestro, K. Redlich, J. Stachel, Phys. Lett. B518 (2001) 41

central Au + Au collisions, data from all experiments combined

$$\chi_r^2 = 0.8$$

$$\chi_r^2 = 1.1$$



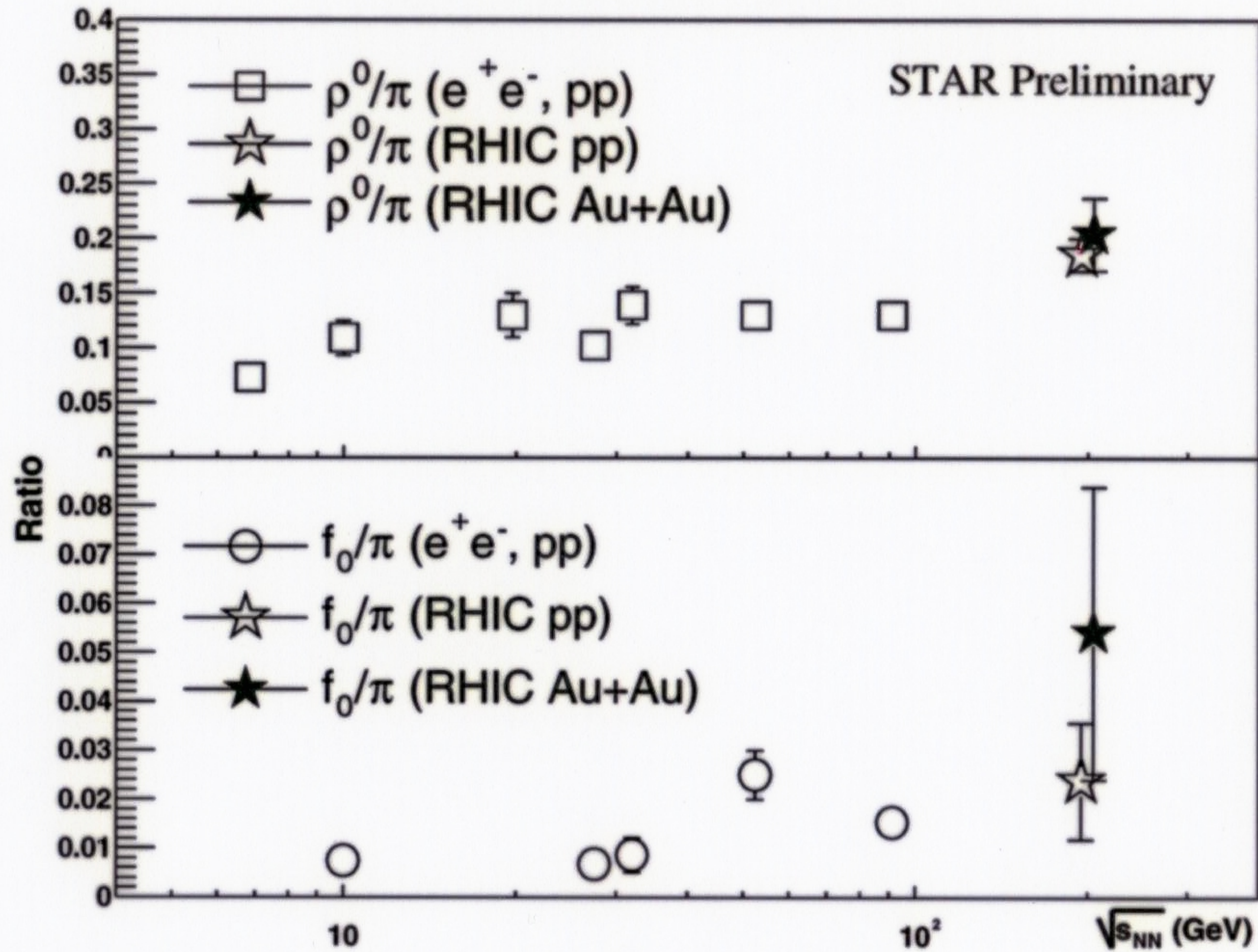
Braun-Munzinger et al., PLB 518 (2001) 41

D. Magestro (updated July 22, 2002)

*fit result confirmed by Becattini and Xu/Kaneta*

# $\rho^0$ and $f_0$ yield at RHIC

P. Fachini, QM2002, nucl-ex/0211001



statistical model  $T=177$  MeV:  $\rho^0/\pi^- = 0.11$  and  $T=120$  MeV:  $4 \cdot 10^{-4}$

even with growth of  $\mu_\pi$  to nearly  $m_\pi$  and  $\Delta m(\rho) \approx 60$  MeV difficult!

## Mass changes close to $T_c$ ?

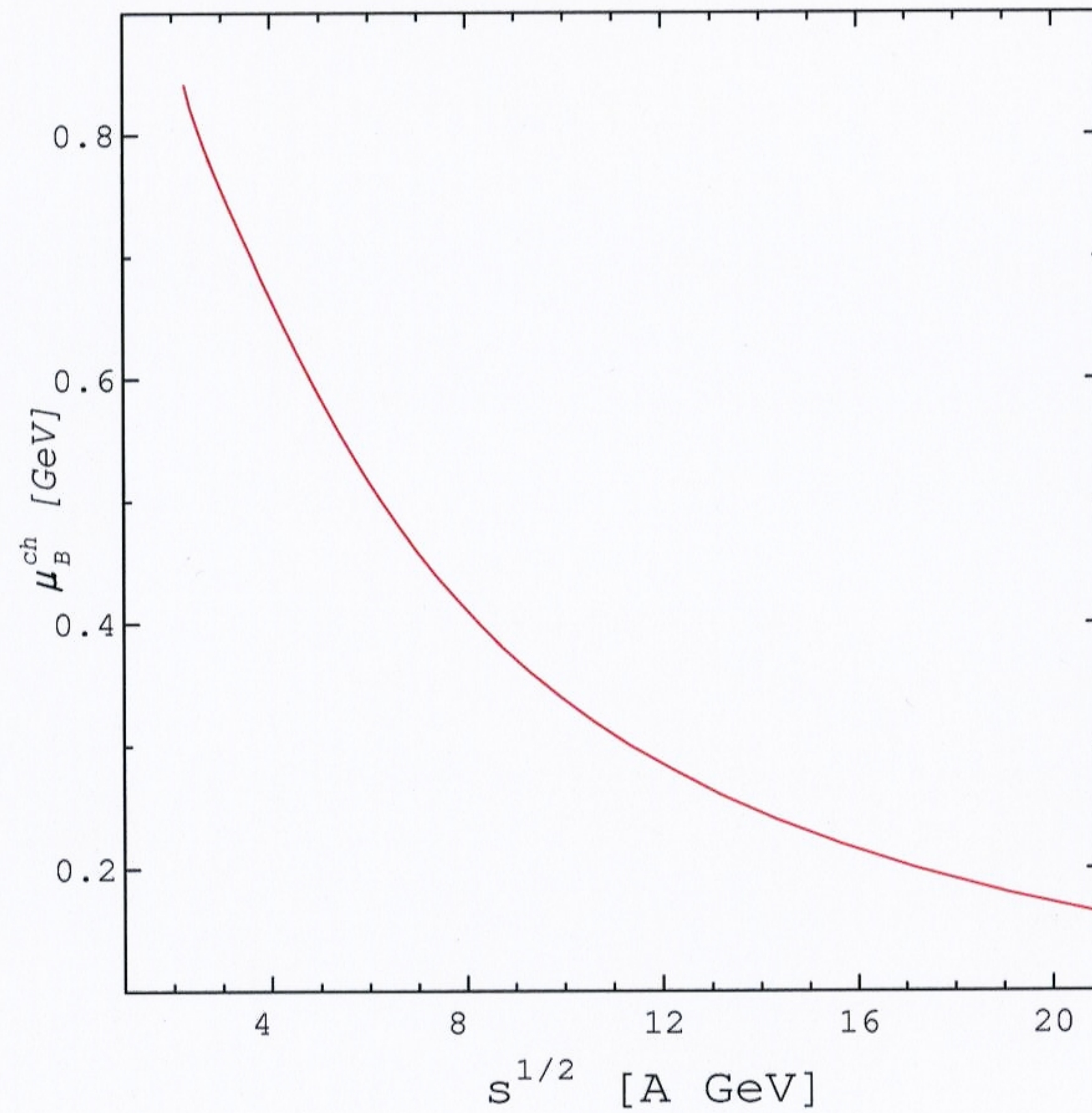
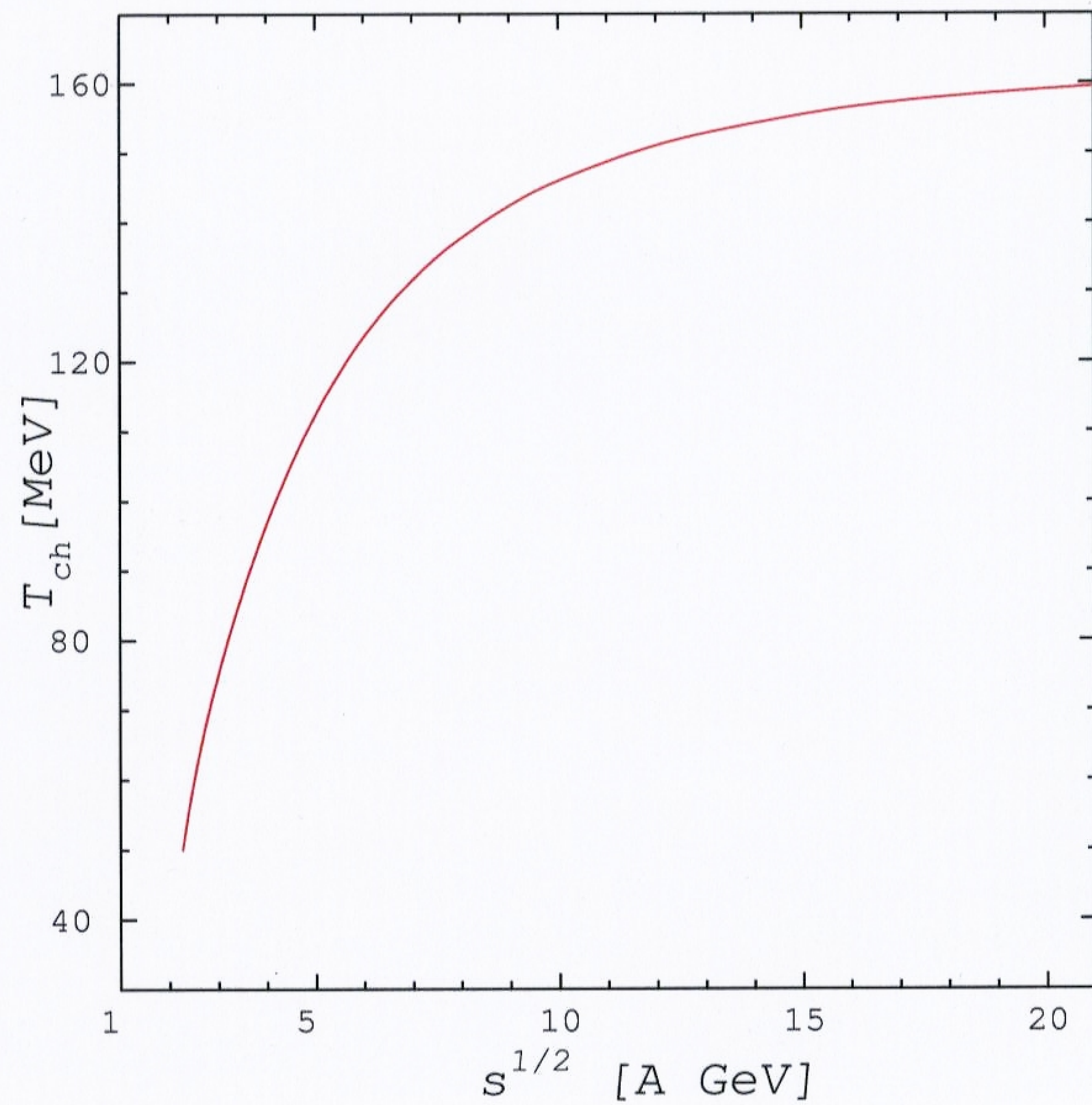
repeat fit of RHIC data with several assumptions

- change all masses by constant factor:  
similar fit quality if variation  $\leq 20\%$   
(see also Michalec, Florkowski, Broniowski  
nucl-th/0103029)
- reduce  $m_\phi$  by 5% : 38 discrepancy w. data
- reduce  $m_{K^0*}$  by 10% : 2.58 " "

no room for very significant changes

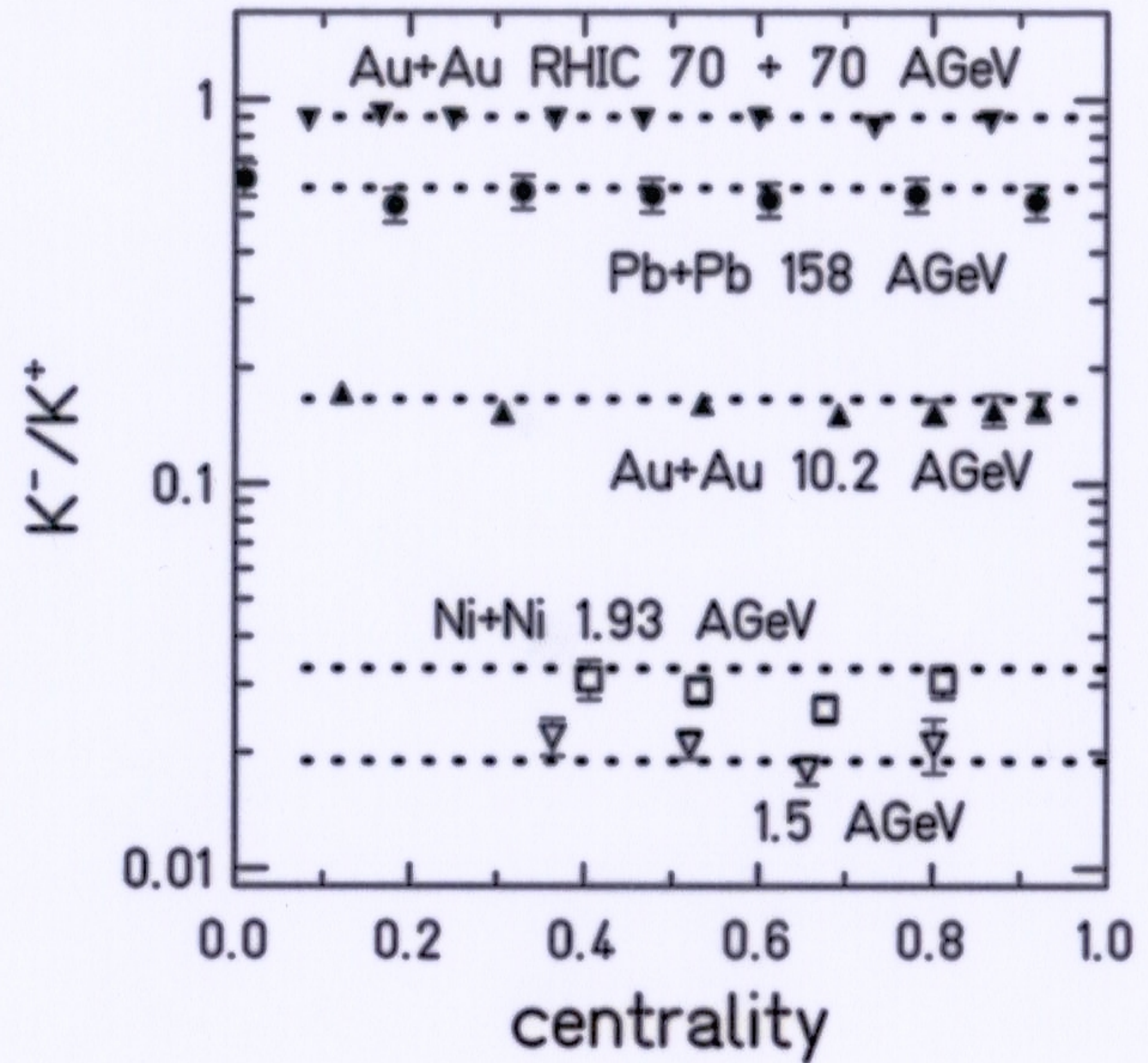
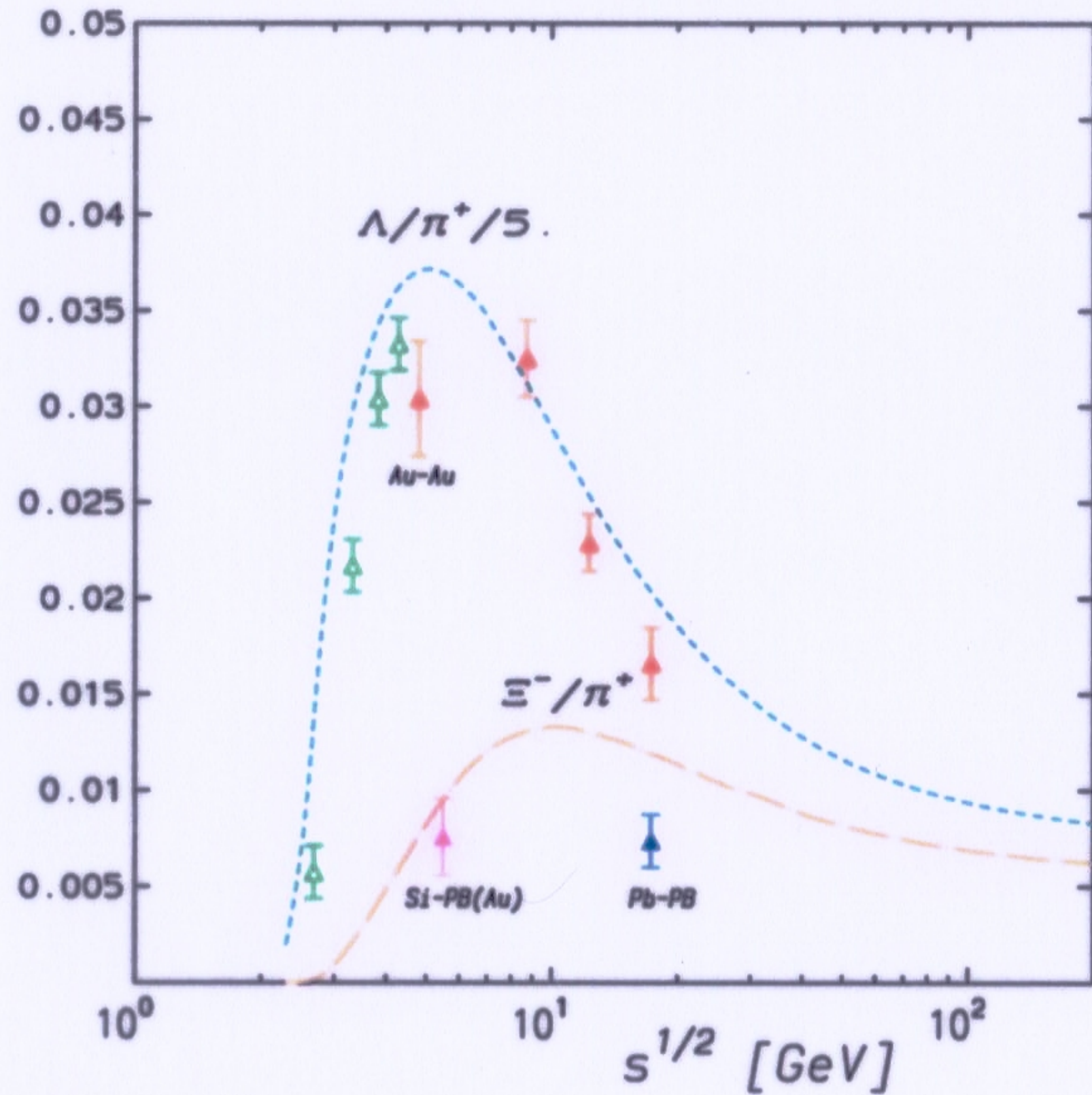
# Evolution of Thermal Parameters with $\sqrt{s}$

Cleymans/Redlich PR C60 (1999) 054908



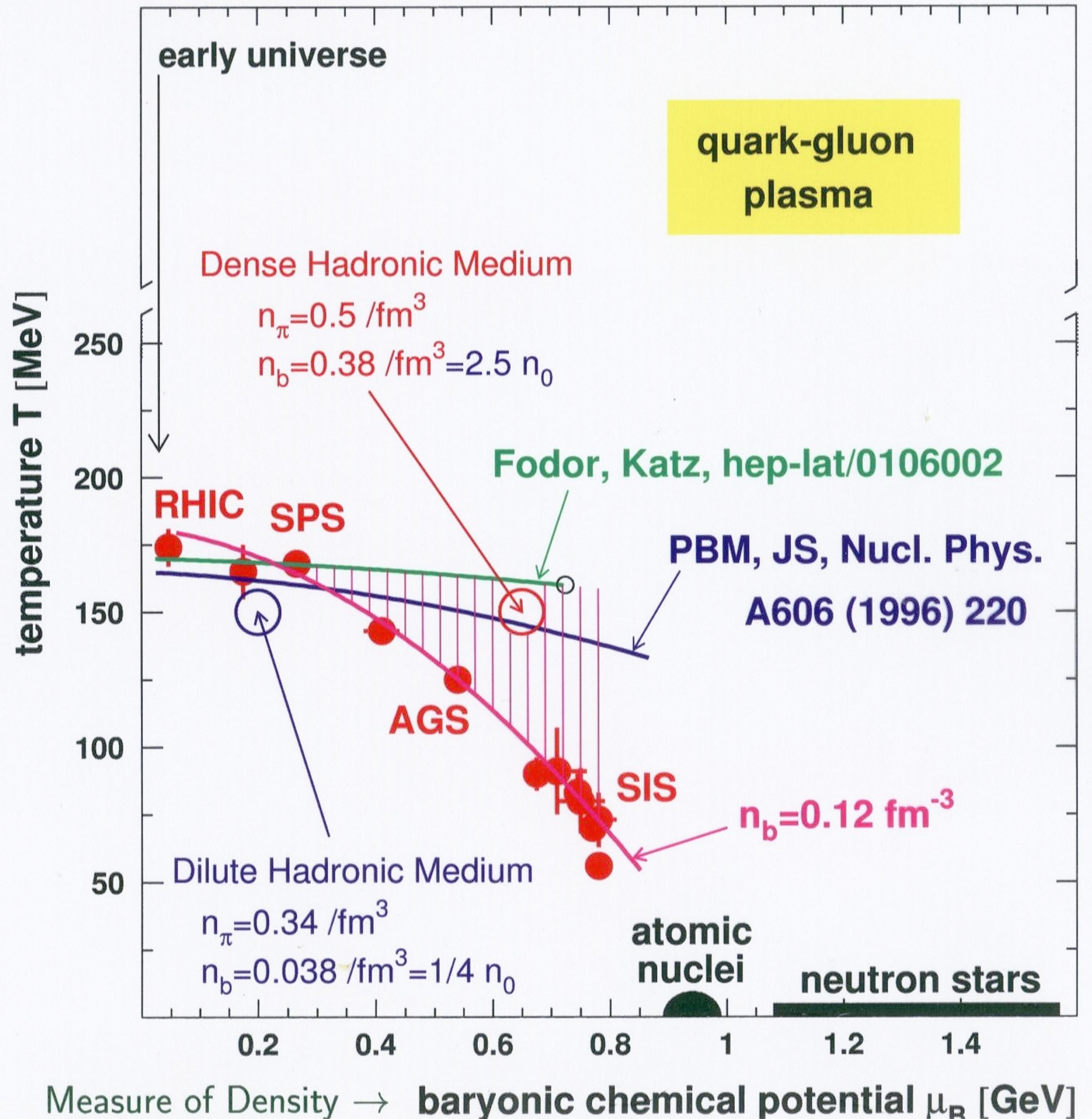
# Evolution of Thermal Parameters with $\sqrt{s}$

Braun-Munzinger, Cleymans, Oeschler, Redlich Nucl. Phys. A697 (2002) 902



# Phase Diagram of Nuclear Matter

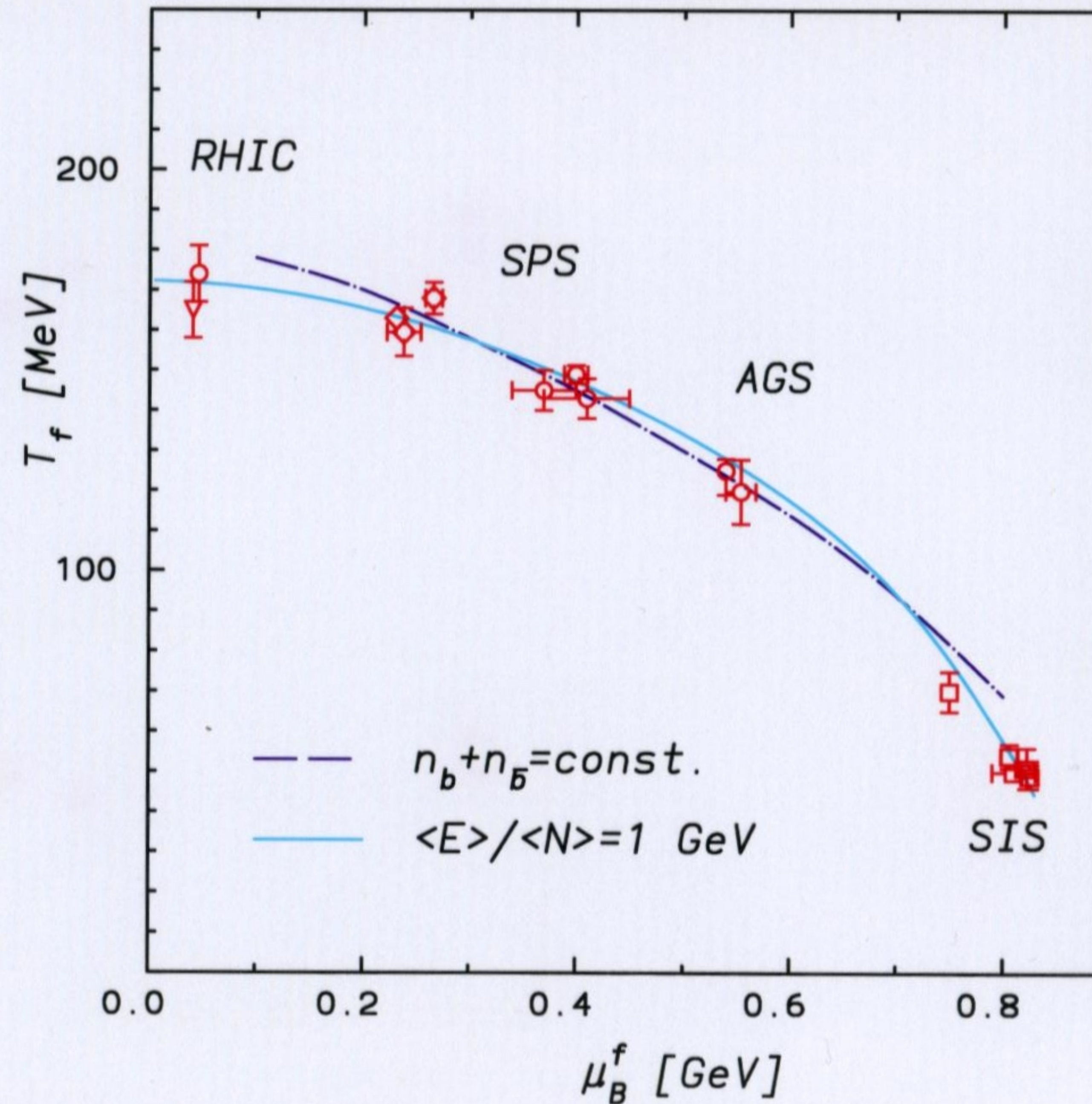
1. hadron yields equilibrated
2. for full SPS energy and above: hadron yields frozen at phase boundary
3. how is equilibrium achieved?  
at SPS and RHIC not with hadronic cross sections  
→ QGP much more efficient equilibrators





# What Characterizes Chemical Freeze-out Curve?

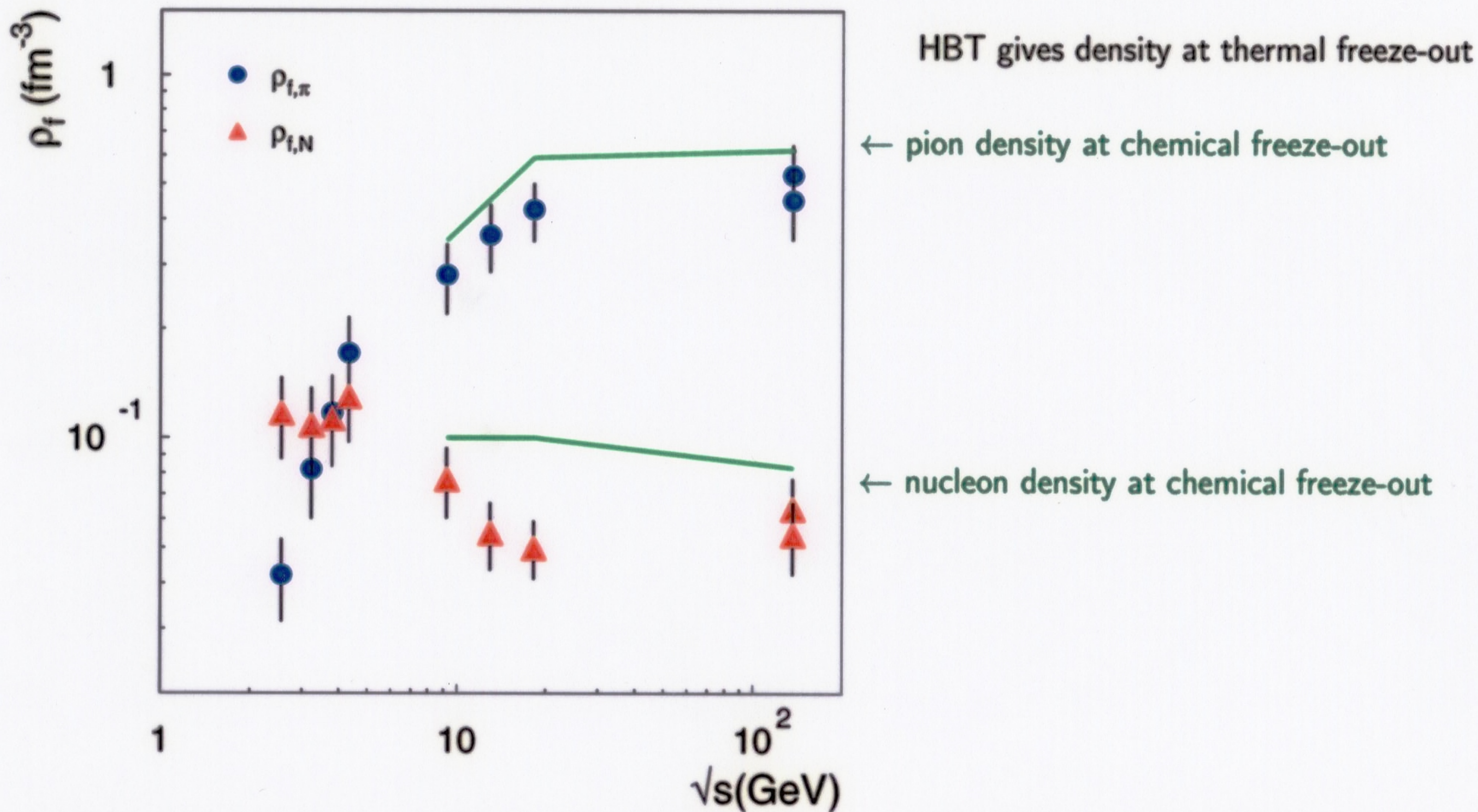
Cleymans/Redlich PRL 81 (1998) 5284: Constant energy per hadron  $E/h = 1$  GeV  
Braun-Munzinger/Stachel JP G28 (2002) 1971: Constant total baryon density  $n_B = 0.12 \text{ fm}^{-3}$



- constant total baryon dens. could indicate bb and mb cross section relevant for chemical freeze-out
- for full SPS energy and above freeze-out points below --- suggestive that actually phase boundary met

# Freeze-Out Density from Pion HBT

CERES PRL 90 (2003) 023001



Volume appears to only grow 30 % between chemical and thermal freeze-out!

## Hadronic Phase after Chemical Freeze-out

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- Densities from HBT as compared to chemical freeze-out:  
indicate growth  $\Delta R_s \approx 15\%$  (at most 40%)  
at average expansion velocity  $0.5c$  this takes  $2 \text{ fm}/c$  (max  $4 \text{ fm}/c$ )
- Role of annihilation for  $\bar{p}/p$  and  $\bar{d}/d$ ; in UrQMD factor 2 change due to annihilation  
in data no room for that  
→ no indication for a longlived hadronic phase
- During hadronization  $\#$  dof increases by factor 3.5  
volume has to grow accordingly  
→ more time at  $T_c$  than in hadronic phase

CERES enhancement understood in this context?

## Excluded Volume to take into account interactions (van der Waals-type)

- different choices in literature
- a thermodynamically consistent correction proposed by Gorenstein, Rischke & Co

$$p^{\text{excl}}(T, \mu) = p^{\text{igas}}(T, \hat{\mu})$$

$$\text{with } \hat{\mu} = \mu - V_{\text{eigen}} p^{\text{excl}}(T, \mu)$$

recursive ... find  $\hat{\mu}$  and then compute any variable  $(T, \mu)$  like  $Z, S, S \dots$

- choice of  $V_{\text{eigen}}$ ? relevant distance is where interaction becomes repulsive  
for nn at 0.3 fm  
mesons?  
our choice  $r = 0.3 \text{ fm}$  for all

→ note: as long as  $r$  is same for all ratios practically not affected by excluded volume correction but absolute densities are!  
caveat for users of freeze-out geometry for e.g. lepton pairs, photons, ...  
for full energy SPS densities reduced by factor 2 - 2.5

fugacities (strangeness suppression  $\gamma_s$ ) in grand canonical

- theoretical limit: this is the canonical strangeness suppression factor

- further (Redlich, ...)  $F_{CS}$  depends explicitly on volume over which quantum number conserv. is enforced

$$\text{for } x_i \ll 1 \quad \frac{I_1(x_i)}{I_0(x_i)} \approx \frac{1}{2} x_i \propto V_0 \leftarrow \frac{k}{\pi}$$
$$V_0^3 \leftarrow \frac{\Omega}{\pi}$$

• for pp and  $e^+e^-$   $V_0 \approx 7 \text{ fm}^3$  volume of hadron  
reason for strangeness supp.

heavy ion coll. (Redlich, ...)  $V \approx V_0 \cdot \frac{N_{part}}{2}$

• Au+Au at SIS  $1400 \text{ fm}^3$

H! for higher energies:  $V$  exceeds this value and loses its meaning; strangeness percolates over large volume

suggestive  $\rightarrow$  (QGP)

$\varphi$  from NA50

N. Willis, proc. Q0799 NPA in print

for central PbPb

$$\varphi / \rho + \omega = 1.10 \pm 0.08 \text{ stat} \pm 0.20 \text{ syst}$$

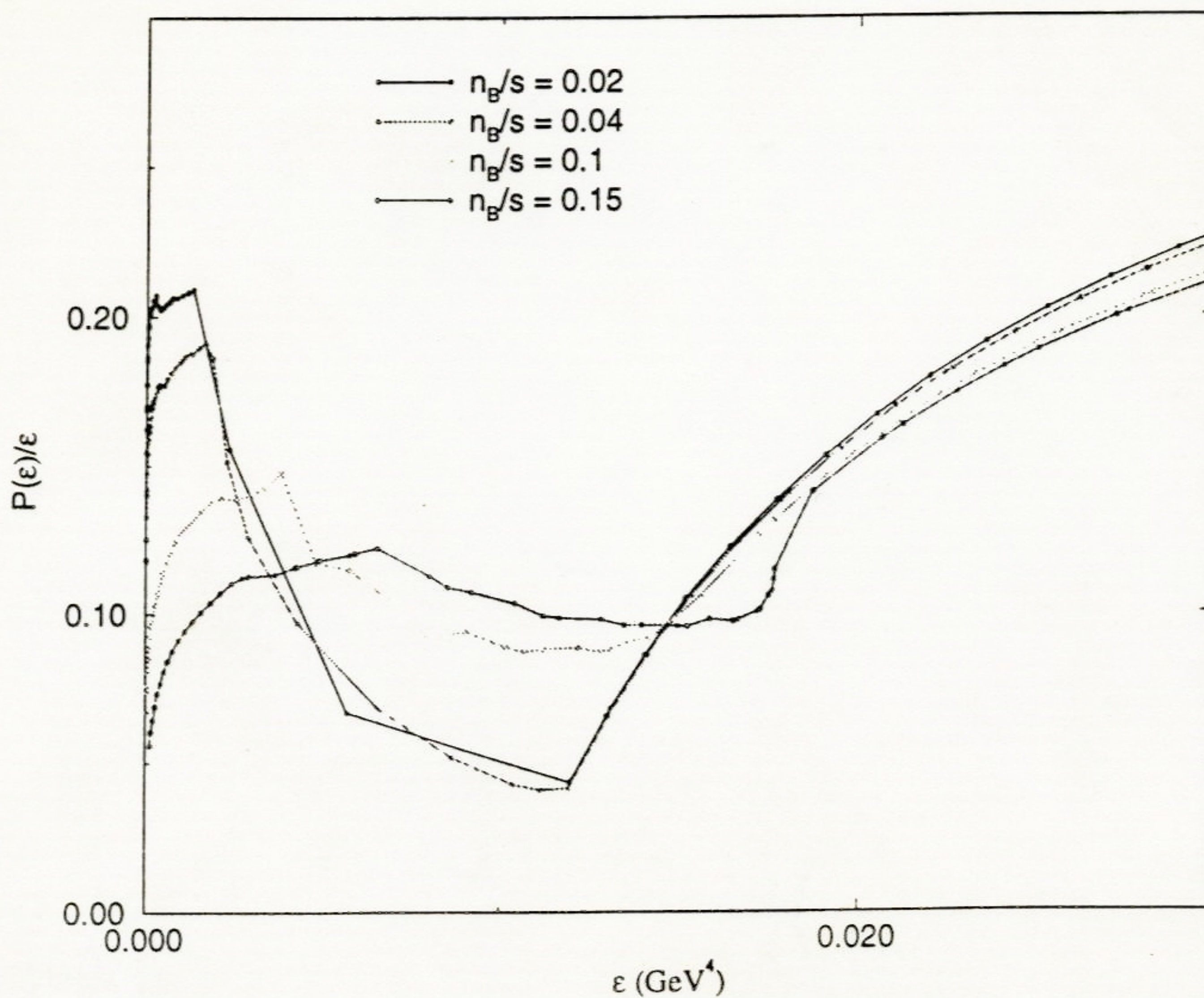
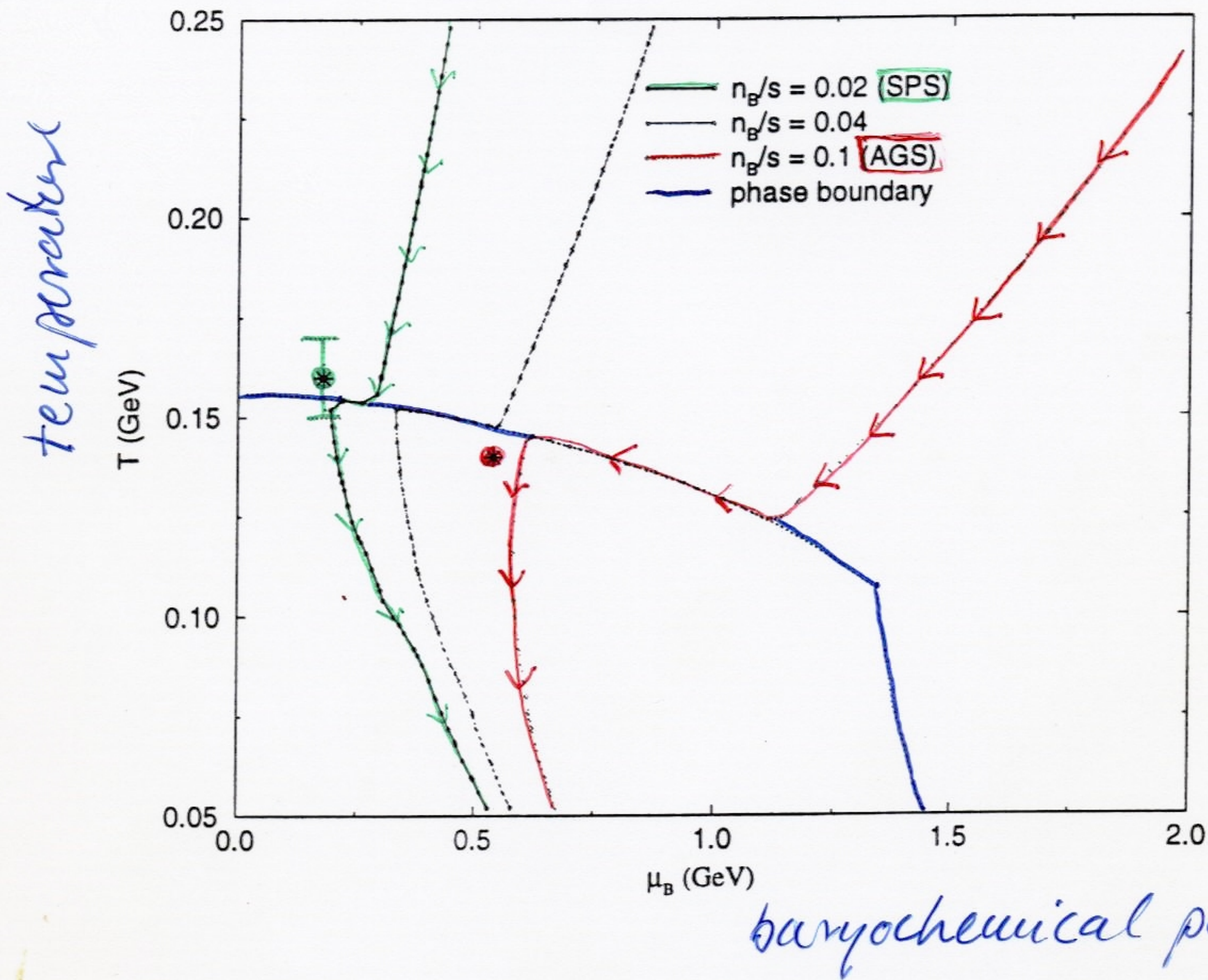
$$T_{\varphi} = 219 \pm 5 \text{ MeV} \quad T_{\rho, \omega} = 222 \pm 6 \text{ MeV}$$

Thermal Model

	yield	$B_{\mu\mu}$	$y \cdot B_{\mu\mu}$	
$\rho_0$	0.0182	$4.60(28) \cdot 10^{-5}$	$8.37 \cdot 10^{-7}$	} $\frac{\varphi}{\rho + \omega} = 0.48 \pm 0.08$
$\omega$	0.0136	$7.07^*(19) \cdot 10^{-5}$	$9.62 \cdot 10^{-7}$	
$\varphi$	0.00345	$2.50(40) \cdot 10^{-4}$	$8.63 \cdot 10^{-7}$	

\* ee

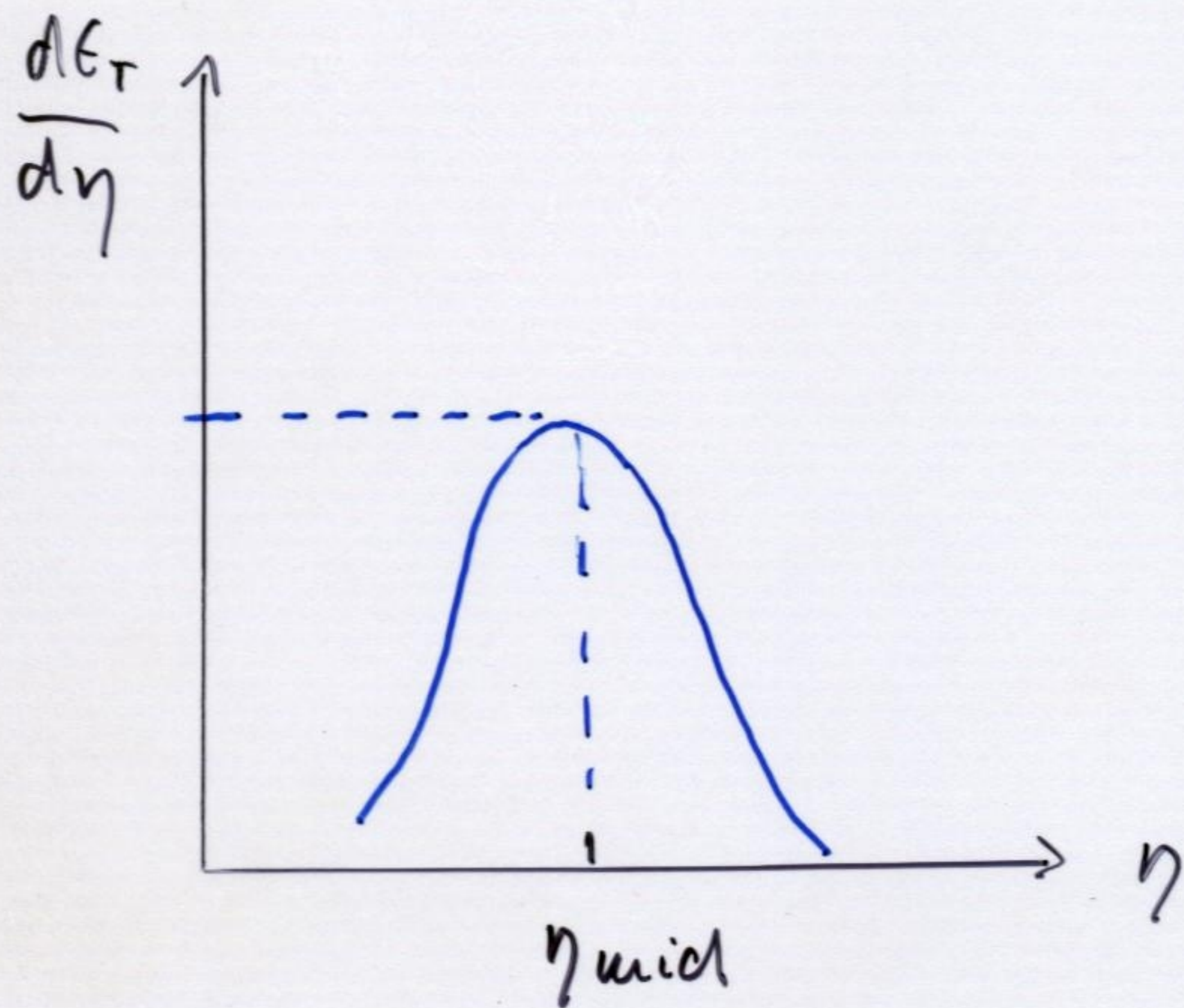
# Phase Diagram and Hydrodynamic Evolution (E.V. Shuryak, Hirscheff 1997)



# estimate of initial energy density

Bjorken 1982

$$\epsilon_0 = \frac{dE_T}{d\eta} \frac{1}{A_{\perp}} \frac{d\eta}{dz} \approx \frac{dE_T}{d\eta} \frac{1}{\pi R_A^2} \frac{1}{\tau_0} \approx 1 \text{ fm}$$



$$\approx \frac{1}{\pi (1.2 \text{ fm } A^{1/2})^2} = 150 \text{ fm}^{-2} \text{ for } A=200$$

similarly initial baryon density  $\rho_0$

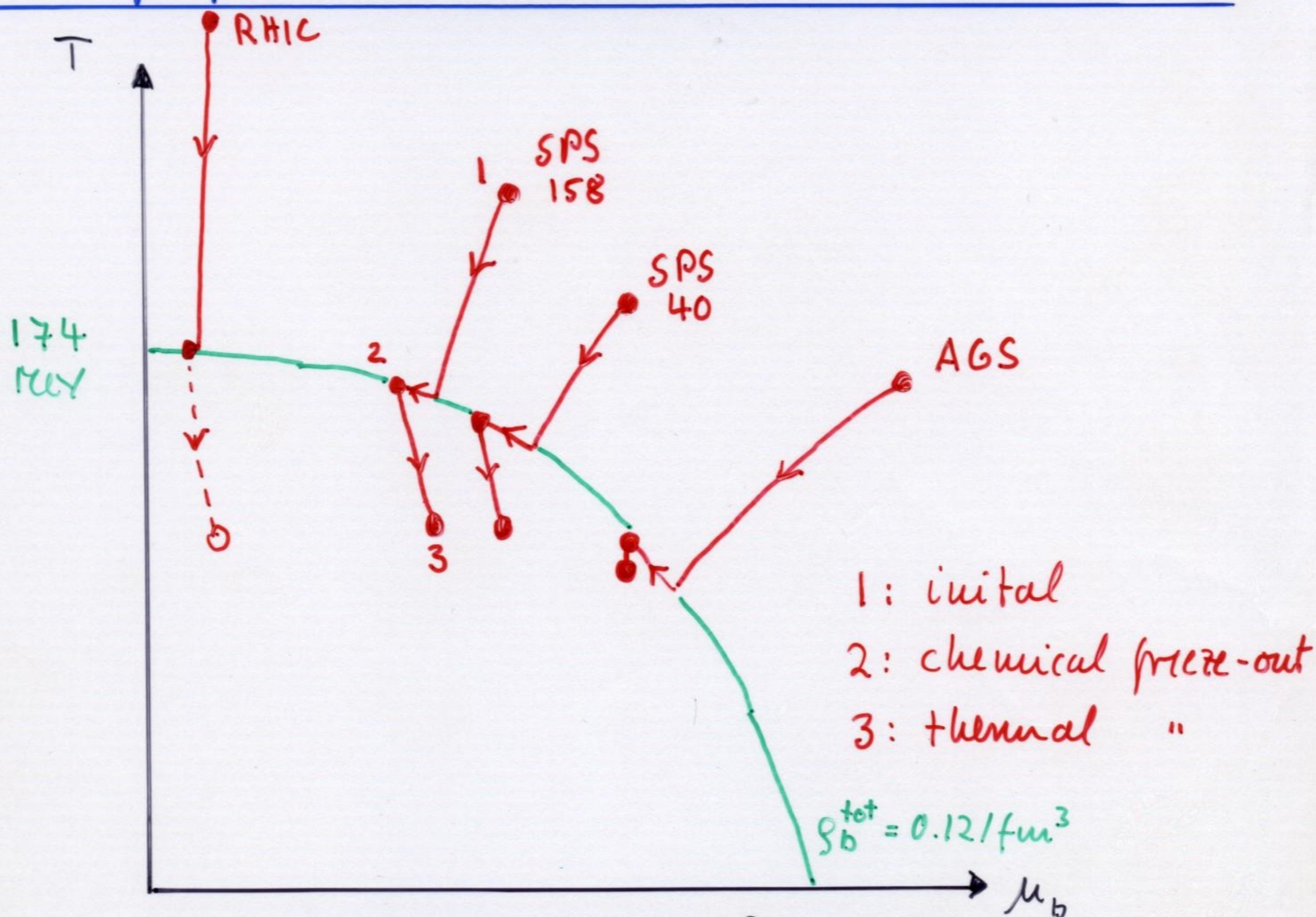
$$\rho_0 = \frac{dN_b}{d\eta} \frac{1}{A_{\perp}} \frac{d\eta}{dz}$$

	AGS	SPS40	SPS158	RHIC
$dE_T/d\eta \text{ max}$	230	-	380	600
$2dN_p/d\eta$	150	114	76	
$\epsilon_0$	1.5	-	2.5	$> 4$ (10-20) GeV/fm <sup>3</sup>
$\rho_0$	1.0	0.75	0.50	1 fm <sup>-3</sup>
$\rho_0/\rho_{nuc}$	$\approx 7$	5	3.3	

$\tau = 0.2 + 0.33 \text{ fm}$



# Temperature and Density at various stages of fireball evolution



	AGS	SPS 40	SPS 158	RHIC	
initial $g_0$	1.0	0.75	0.50		} $dG_T/d\eta$ $dN_p/d\eta$
$\epsilon_0$	1.5		2.5	10-20	
chem $T/\mu_b$	125/540	148/400	170/255	177/41	} stat. model ↔ hadron yields
nuc. dens. $\oplus$	.126	.10	.10	.08	
pion dens	.17	.35	.59	.62	
thermal					
nuc. dens	.106	.077(5)	.063(5)	.060(9)	} HBT $2\pi$ -corr
pion dens	.15	.28(3)	.43(3)	.49(10)	
T	120	120	120	120?	
$V_{\text{therm}}/V_{\text{chem}}$	.86	← = 1.3 →			
note:	≈ equal 3 & 2 coincide	3d expansion at $v/c \approx 0.25$ (known from hydro) → 0.5		$R: 5.5 \rightarrow 6.0$ fm 2 fm/c 1 fm/c	

T	$E_{\text{lat}} \left( \frac{\text{TeV}}{\text{fm}^3} \right)$	$\frac{E}{T^4}$	$E_{\text{HG}}$
105.7	0		20.9
117.3	4.06		35.3
131.3	17.0		80.3
147.9	85.9		173.3
167.1	413.5		368.5
175.6	730.4		489.0
189.3	1471		719.3
214.3	2945		-

P. Chung:  $6895 \Xi^-$  analysis, neural network, &  $\Lambda$   
 $\Xi \rightarrow \Lambda + \Sigma^0$   
 $\begin{matrix} 0.007 \\ 0.01 \\ 0.014 \end{matrix}$  / centrality very good agreement w. statistical model  
 6 GeV

H. Huang:  $K^+$  and  $K^-$  comparison SUDA  
 also does not scale w.  $N_{\text{part}}$

thermal model check:  $\Xi^-$  to  $\Sigma^0$   $\Lambda$  unlikely  
 $\begin{matrix} / & | & 0.013 \\ & 0.0029 & (0.0025 \text{ min}) \\ & (0.0017p) & \\ & 0.0016 & \\ & (0.0009p) & \end{matrix}$

$R_{AA}$   $\Lambda + \bar{\Lambda}$  goes close to 1 drops  $\mu$   $p_T \approx 4 \text{ GeV}$   
 $K_S^0$  0.7